# M3D Simulation Studies of NSTX

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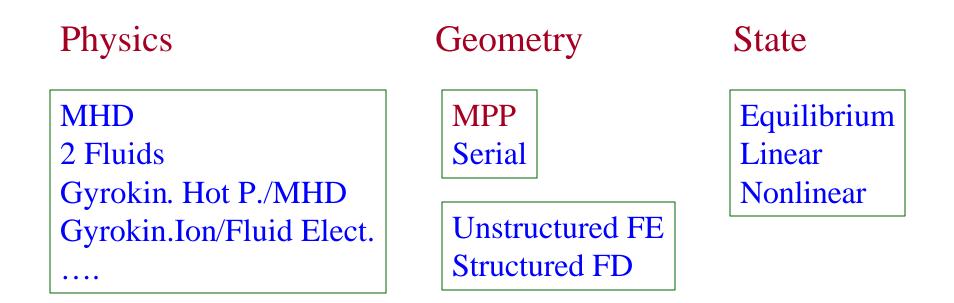
# Outline

- M3D code
  - MHD, two-fluids, hybrid models.
- NSTX studies including flow effects
   2D steady states.
   Evolutions of IRE's.
   BAE modes.



W. Park et al., Phys. Plasmas **6**, 1796 (1999) http://w3.pppl.gov/~wpark/pop\_99.pdf

Multilevel 3D Project for Plasma Simulation studies Various physics levels are needed to understand the physics. The best method depends on the problem at hand.



#### MHD model

#### Solves MHD equations.

 $\rho \partial \mathbf{v} / \partial t + \rho \mathbf{v} \cdot \nabla \mathbf{v} = -\nabla p + \mathbf{J} \times \mathbf{B} + \mu \nabla^2 \mathbf{v}$  $\partial \mathbf{B} / \partial t = -\nabla \times \mathbf{E}, \quad \mathbf{E} = (-\mathbf{v} \times \mathbf{B} + \eta \mathbf{J}), \quad \mathbf{J} = \nabla \times \mathbf{B}$  $\partial \rho / \partial t + \nabla \cdot (\rho \mathbf{v}) = 0$  $\partial p / \partial t + \mathbf{v} \cdot \nabla p = -\gamma p \nabla \cdot \mathbf{v} + \rho \nabla \cdot \kappa \nabla (p/\rho)$ 

The fast parallel equilibration of T is modeled using wave equations;

$$\begin{pmatrix} \partial T / \partial t = s \mathbf{B} / \rho \cdot \nabla u \\ \partial u / \partial t = s \mathbf{B} \cdot \nabla \mathbf{T} + \upsilon \nabla^2 u & s = wave speed / v_A \end{pmatrix}$$

#### Two-fluid MH3D-T

 Solves the two fluid equations with gyro-viscousity and neoclassical parallel viscousity terms in a torus.

#### Equations

$$\mathbf{v} \equiv \mathbf{v}_{i} - \mathbf{v}_{i}^{*} = \mathbf{v}_{e} - \mathbf{v}_{e}^{*} + \mathbf{J}_{\parallel}/\text{en},$$
  
 $\mathbf{v}_{e}^{*} \equiv -\mathbf{B} \mathbf{x} \nabla \mathbf{P}_{e} /(\text{enB}^{2}), \quad \mathbf{v}_{i}^{*} \equiv \mathbf{v}_{e}^{*} + \mathbf{J}_{\perp}/\text{en},$ 

 $\rho \partial \mathbf{v} / \partial t + \rho \mathbf{v} \cdot \nabla \mathbf{v} + \rho (\mathbf{v}_i^* \cdot \nabla) \mathbf{v}_{\perp} = -\nabla p + \mathbf{J} \times \mathbf{B} - \mathbf{b} \cdot \nabla \cdot \Pi \mathbf{i},$ 

 $\partial \mathbf{B}/\partial t = -\nabla \times \mathbf{E}, \quad \mathbf{E} = (-\mathbf{v} \times \mathbf{B} + \eta \mathbf{J}) - \nabla_{\!\!\!\Pi} \mathbf{P}_{\!\!\mathbf{e}} / \mathbf{en} - \mathbf{b} \cdot \nabla \cdot \Pi_{\!\!\mathbf{e}},$  $\mathbf{J} = \nabla \times \mathbf{B},$ 

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\partial \rho / \partial t + \nabla \cdot (\rho \mathbf{v}_j) = 0,
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 $\frac{\partial \mathbf{p}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{p} = -\gamma \mathbf{p} \nabla \cdot \mathbf{v} + \rho \nabla \cdot \kappa_{\parallel} \nabla_{\parallel} (\mathbf{p}/\rho)$  $- \mathbf{v}_{i}^{*} \cdot \nabla \mathbf{p} + (1/en) \mathbf{J} \cdot \nabla \mathbf{p}_{e}$  $- \gamma \mathbf{p} \nabla \cdot \mathbf{v}_{i}^{*} + \gamma \mathbf{p}_{e} \mathbf{J} \cdot \nabla (1/en)$ 

 $\frac{\partial P_{e}}{\partial t} + \mathbf{v} \cdot \nabla P_{e} = -\gamma P_{e} \nabla \cdot \mathbf{v} + \rho \nabla \cdot \kappa_{\parallel} \nabla_{\parallel} (P_{e} / \rho)$ + (1/en)  $\mathbf{J}_{\parallel} \cdot \nabla P_{e} - \gamma P_{e} \nabla \cdot (\mathbf{v}_{e}^{\star} - \mathbf{J}_{\parallel} / en)$ 

#### GK Hot Particle /MHD Hybrid MH3D-K

#### Fluid equations

$$\partial \mathbf{B}/\partial t = -\nabla \times \mathbf{E}, \quad \mathbf{E} = \mathbf{v} \times \mathbf{B} - \eta (\mathbf{J} - \mathbf{J}_h), \quad \mathbf{J} = \nabla \times \mathbf{B}$$

 $\partial \rho / \partial t + \nabla \cdot (\rho \bm{v}) = 0$ 

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\partial \mathbf{p} / \partial t + \mathbf{v} \cdot \nabla \mathbf{p} = -\gamma \mathbf{p} \nabla \cdot \mathbf{v} + \rho \nabla \cdot \kappa \cdot \nabla (\mathbf{p} / \rho)
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Gyrokinetic equations for energetic particles

 $d\mathbf{R}/dt = \mathbf{u}[\mathbf{b} + (\mathbf{u}/\Omega)\mathbf{b} \times (\mathbf{b}\cdot\nabla\mathbf{b})] + (\mathbf{1}/\Omega)\mathbf{b} \times (\mu\nabla\mathbf{B} - q\mathbf{E}/m),$  $d\mathbf{u}/dt = -[\mathbf{b} + (\mathbf{u}/\Omega)\mathbf{b} \times (\mathbf{b}\cdot\nabla\mathbf{b})] \cdot (\mu\nabla\mathbf{B} - q\mathbf{E}/m).$ 

#### **GK Particle Ion / Fluid Electron Hybrid**

#### Pressure coupling

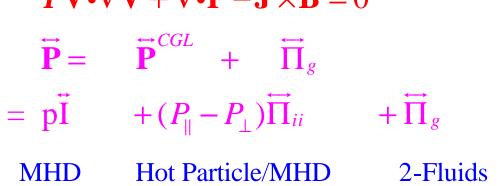
$$\rho \partial \mathbf{v} / \partial t + \rho \mathbf{v} \cdot \nabla \mathbf{v} = -\nabla \cdot \mathbf{P} \mathbf{i} - \nabla \mathbf{P} \mathbf{e} + \mathbf{J} \times \mathbf{B}$$
$$= -\nabla \cdot \mathbf{P} \mathbf{i}^{CGL} - \nabla \cdot \Pi \mathbf{i} - \nabla \mathbf{P} \mathbf{e} + \mathbf{J} \times \mathbf{B}$$

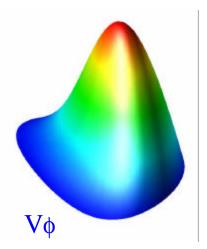
V·Pi<sup>CGL</sup>: from particles following GK eqns.
 V·Πi : fluid picture as 2 fluid eqns, or from particles.

# Fluid electrons E = - Ve × B + ηJ + ∇· Pe /ne = -Ve × B + ηJ + ∇Pe /ne + bb· ∇·Πe /ne ∂B/∂t = -∇×E, J= ∇×B Pe eqn currently, but P<sub>µ</sub> and P<sub>⊥</sub> eqns are planned.

## 2D steady state with toroidal sheared flow

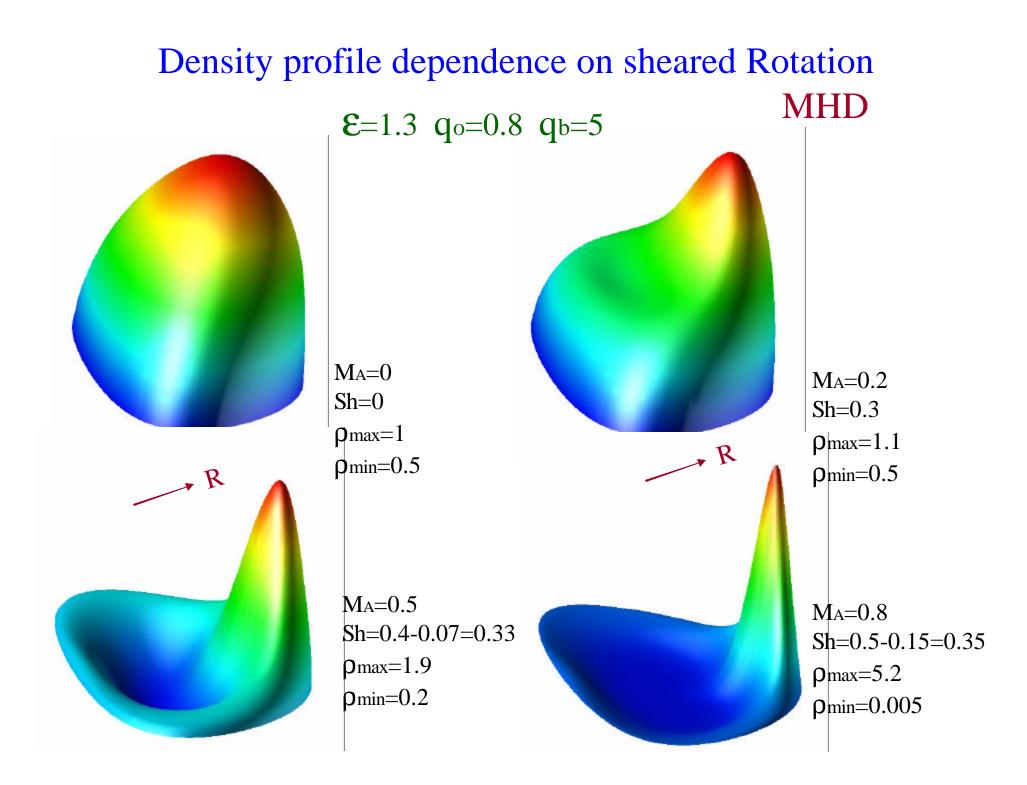
Quasi neutrality:  $\mathbf{r}\mathbf{V}\cdot\nabla\mathbf{V} + \nabla\cdot\mathbf{\ddot{P}} - \mathbf{J}\times\mathbf{B} = 0$ 



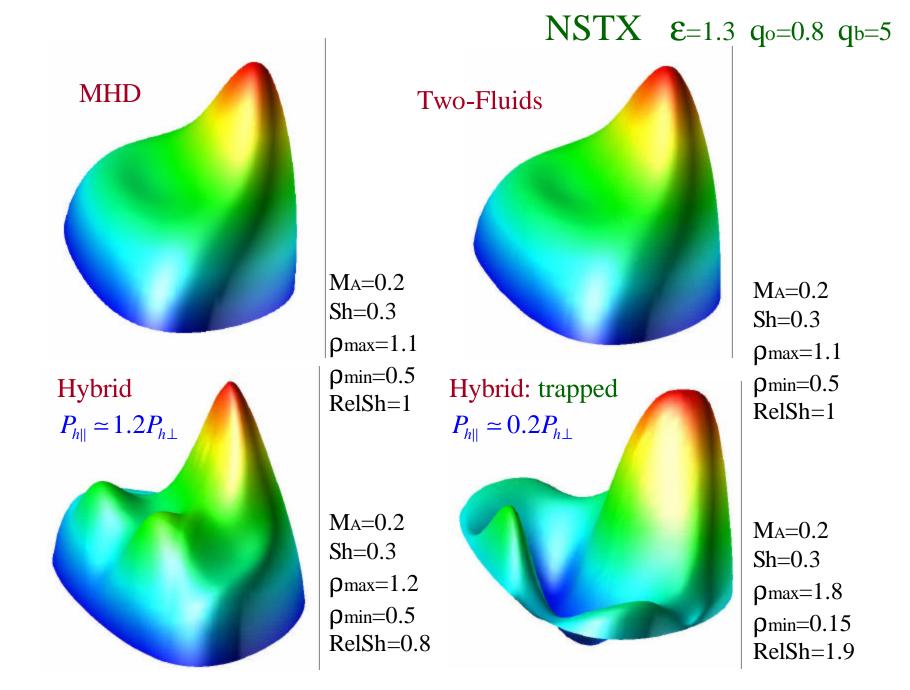


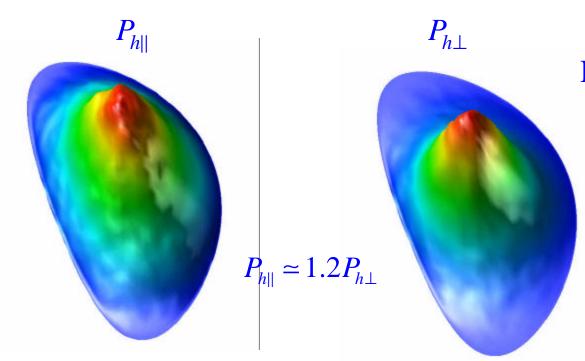
MHD:

At the magnetic axis:  $\mathbf{J} \times \mathbf{B} = 0$   $-\frac{\mathbf{r} V_f^2}{R} + \frac{T \partial \mathbf{r}}{\partial R} = 0$ Relative shift of  $\mathbf{r} \equiv \frac{R \partial \mathbf{r}}{r \partial R} = \frac{V_f^2}{T} = \frac{2M_A^2}{b}$ 



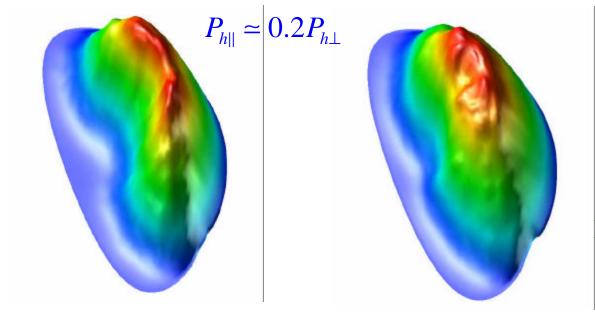
## Density profile dependence on Physics model





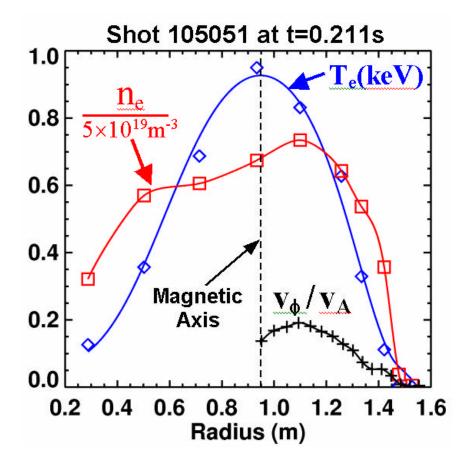
Hot particle pressure  $\mathbf{P}_h$ in the hybrid simulation

> Similar to Experimental situation



Mostly trapped particles

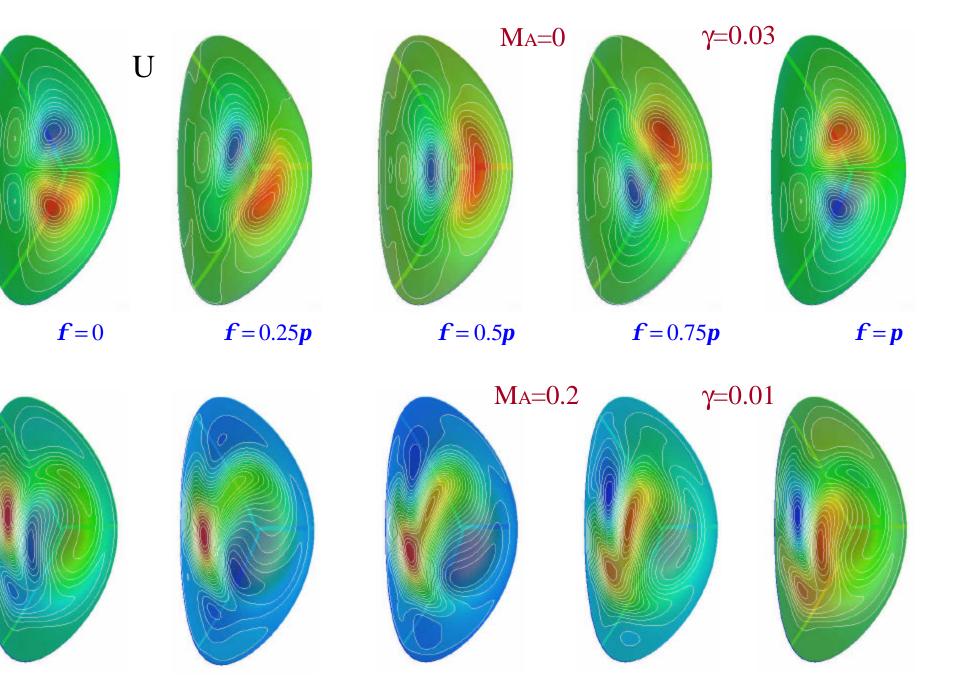
#### NSTX experimental data



Relative shift of  $\boldsymbol{r}$  $\frac{R\partial \boldsymbol{r}}{\boldsymbol{r}\partial R} = \frac{2M_A^2}{\boldsymbol{b}}$ 

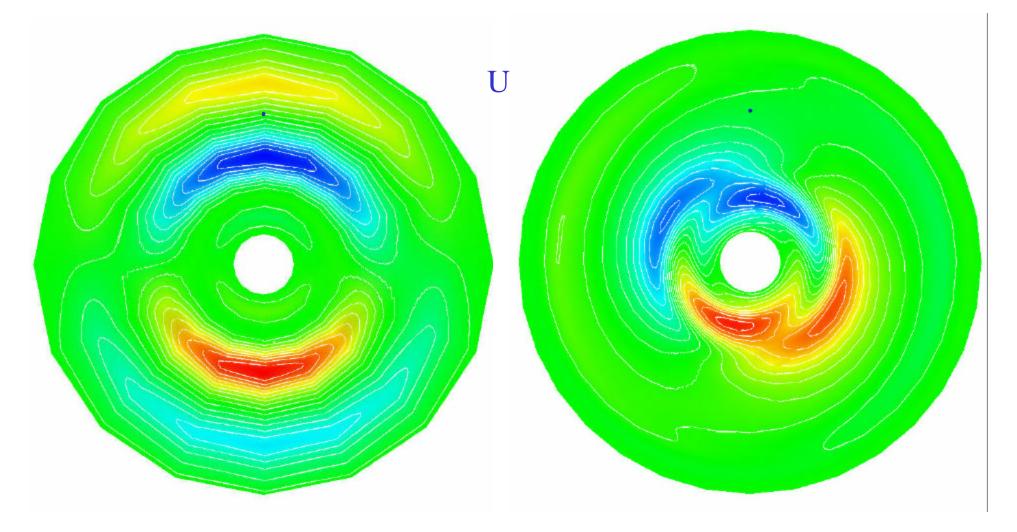
## Hot particle centrifugal force ~ Bulk plasma

### Linear Eigenmodes: shear flow reduces growth rate

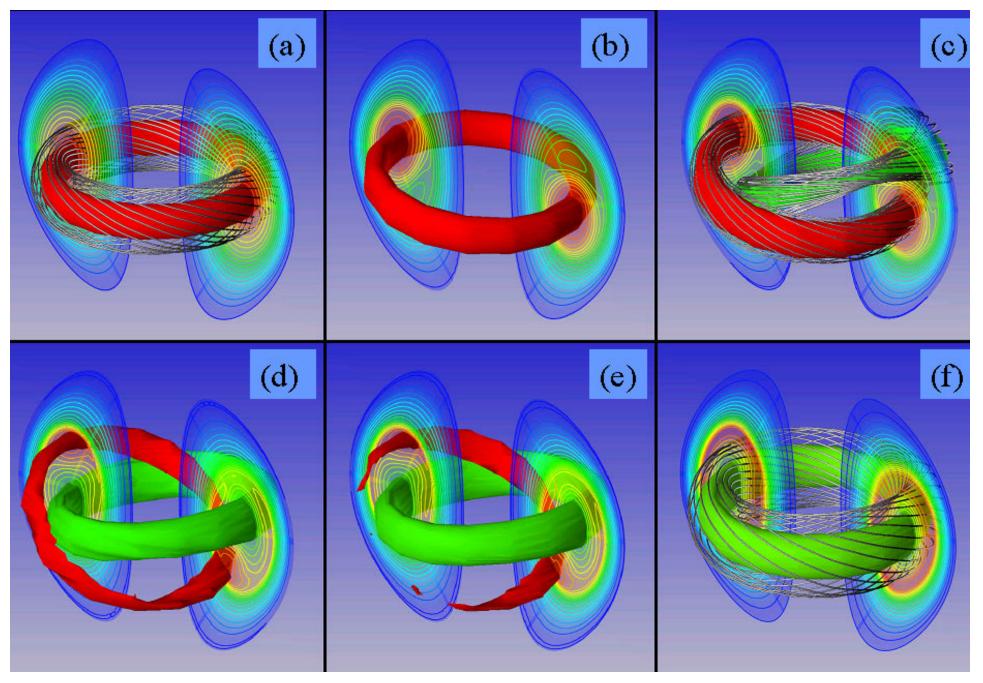


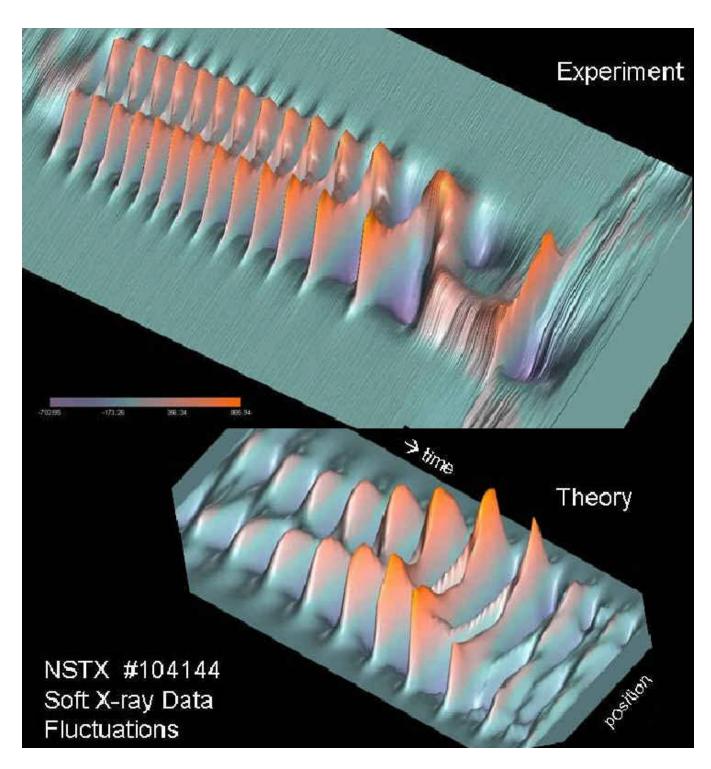
## Linear Eigenmodes Top view on the mid-plane

MA=0 Ωm=0 With shear flow: MA=0.2 Rotating mode:  $\Omega$ m=0.13



### Nonlinear Evolution without strong flow: similar to a sawtooth crash

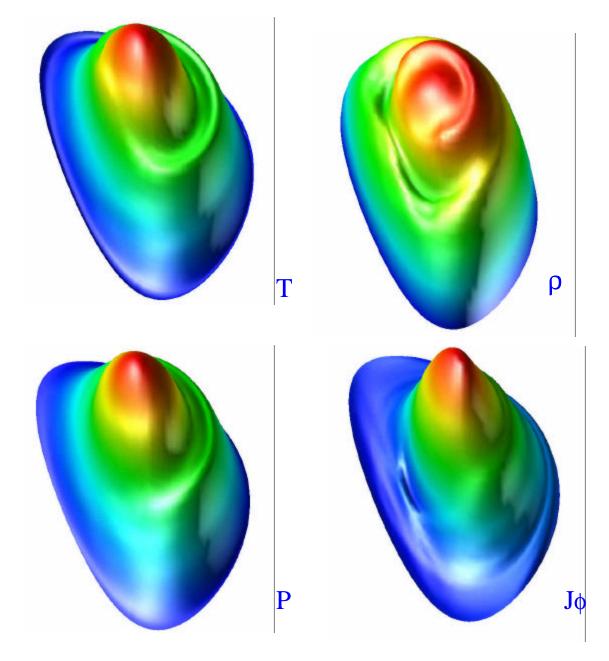


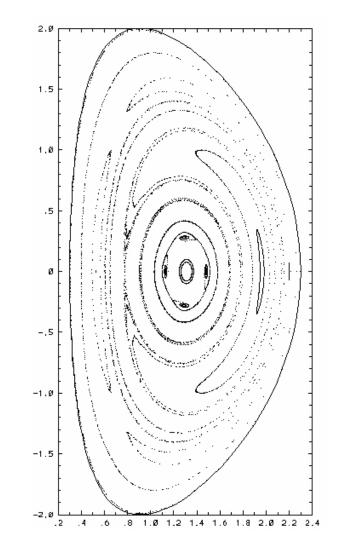


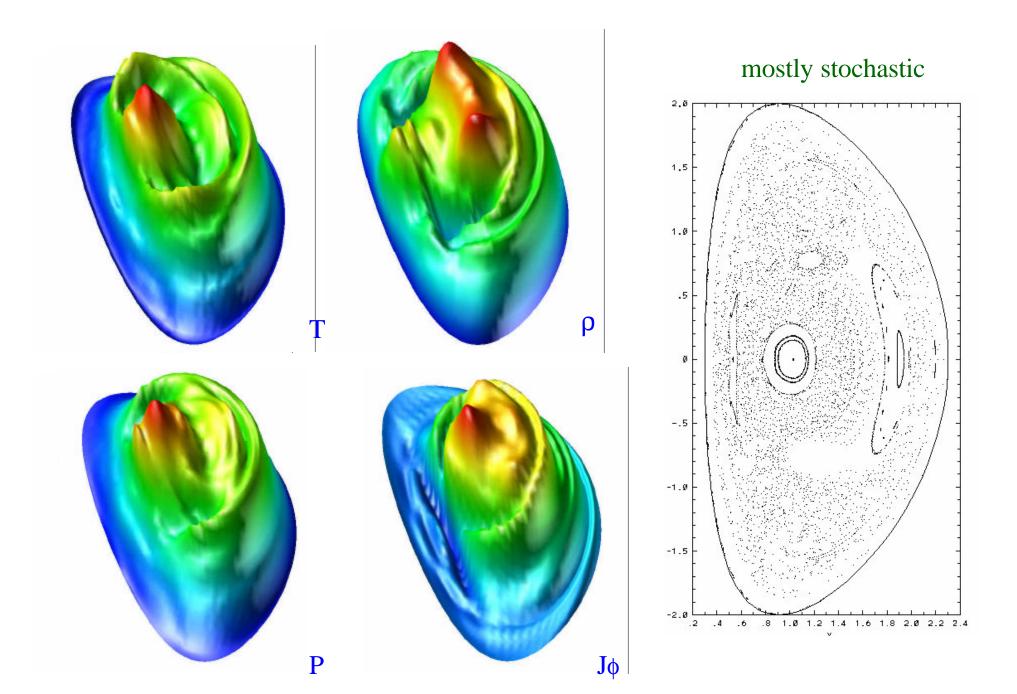
# Soft X-ray signals compared:

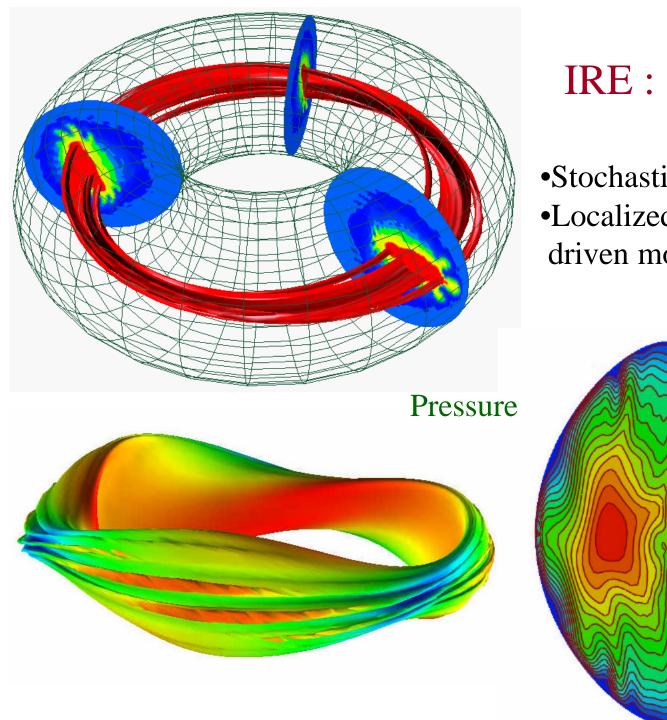
Theory agrees with experiment on general characters, but does not have wall locking and a saturation phase.

## Nonlinear Evolution without strong flow





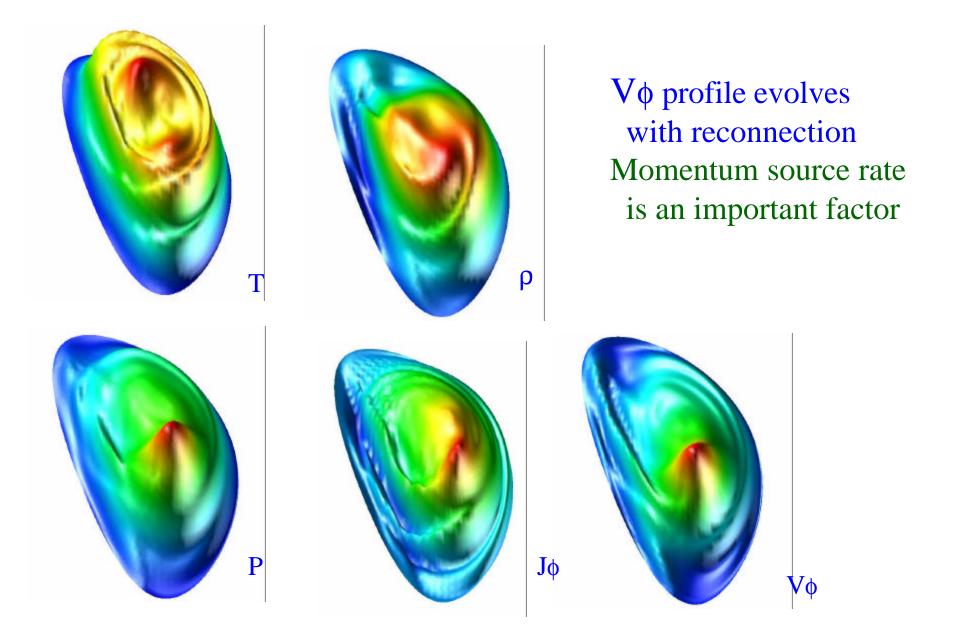




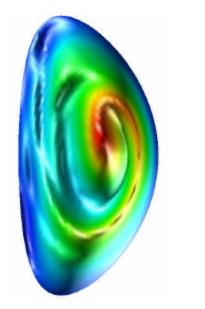
IRE : Disruption

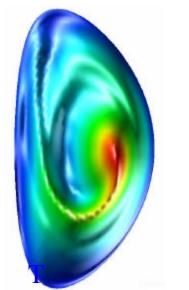
Stochasticity as shown before.
Localized steepening of pressure driven modes as shown here.

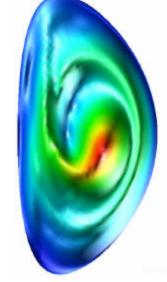
## Nonlinear Evolution with peak rotation of $M_A=0.2$

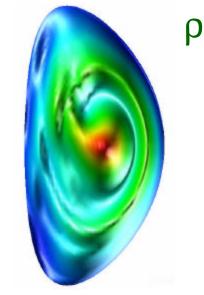


## $\rho$ (P) and T out of phase in a saturated case







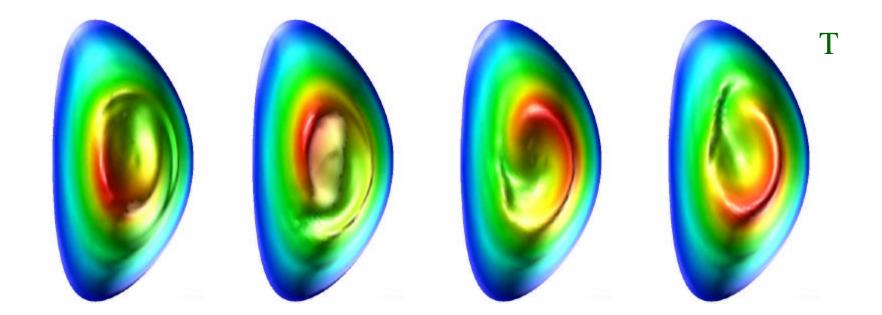


 $\mathbf{f} = 0$ 

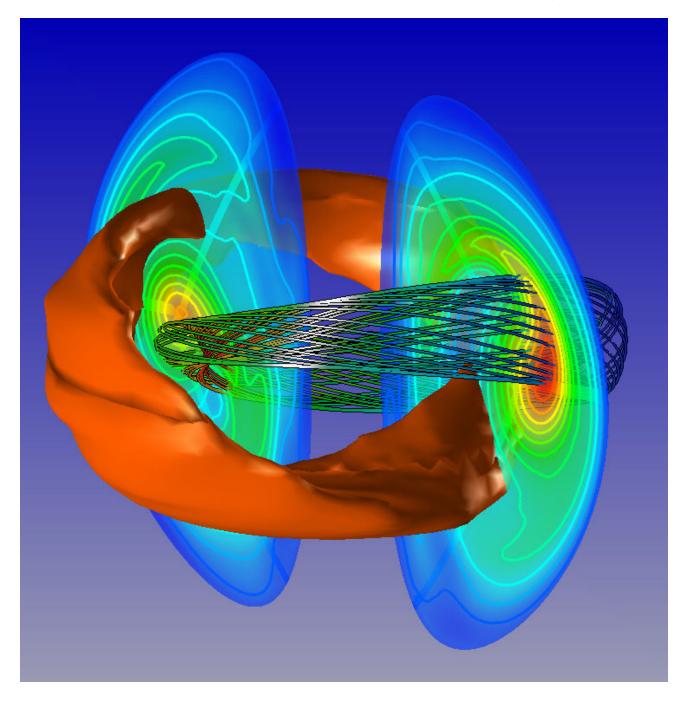
f = 0.5p

**f**=1.5**p** 

**f** = **p** 

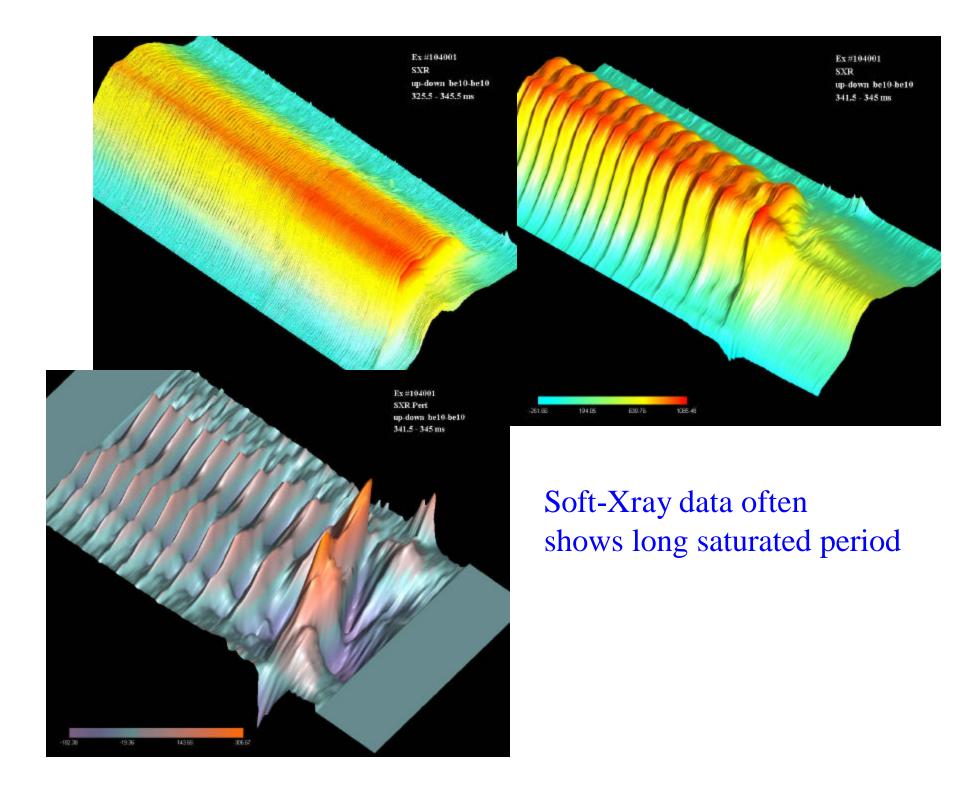


## Saturated steady state with strong sheared flow



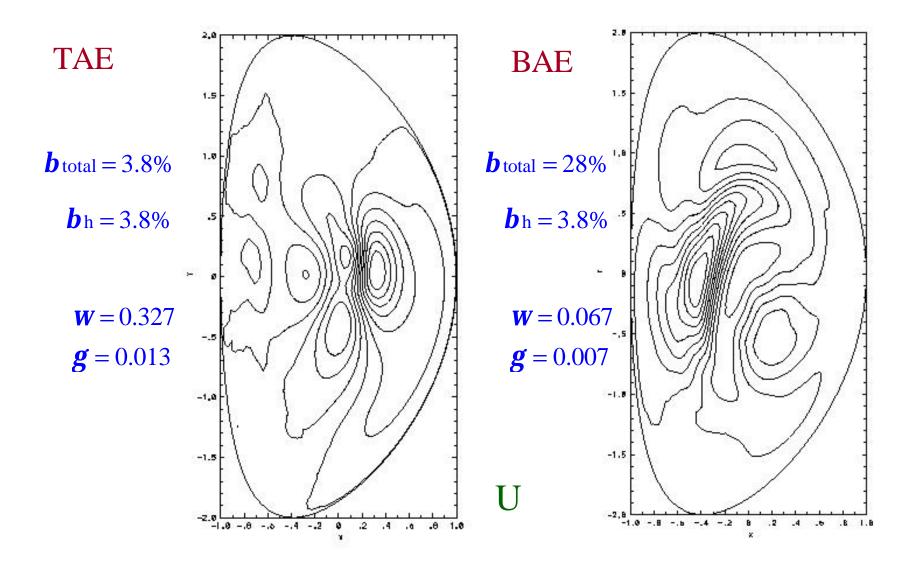
**B** Field line in the island Density (Pressure) contours Temperature isosurface

Pressure peak inside the island together with shear flow causes the mode saturation.

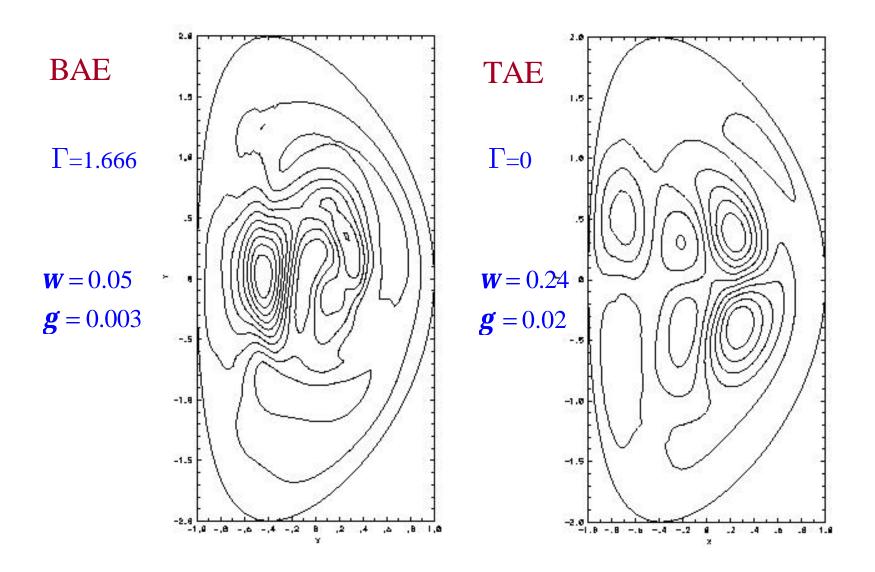


#### EPM (BAE) is excited at high beta in hybrid simulations

More coupling to sound wave due to stronger curvature and high beta. May explain experimental data.



#### BAE changes to TAE when $\Gamma$ is set to zero



# Summary

- M3D code with MHD, Two-fluids, and Particle/Fluid Hybrid levels is used to study NSTX.
- The relative density shift relation holds both in the simulation and experiment, with the centrifugal force of the hot component included.
- Toroidal sheared rotation reduces linear growth and can saturate internal kink.
- IRE:Disruption can occur in at least in two ways; due to stochasticity, and due to localized steepening of pressure driven modes.
- BAE mode is found which may explain experimental data.
- Resistive wall, vacuum region, and external coils are being added to M3D code to expand the regime of applicability.