
Analysis of magnetic perturbation in BOUT runs for NSTX

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Magnetic perturbation in BOUT turbulence simulations comes from fluctuating j_{\parallel}

- $\nabla_{\perp}^2 A_{\parallel} = -\frac{4\pi}{c} j_{\parallel} \Rightarrow$ no δB_{\parallel} is kept in BOUT model
- Magnetic perturbation is important for theoretical understanding of regimes of turbulence
- Magnetic perturbation induces anomalous transport

$$\Gamma_r = \langle \tilde{\Gamma}_{\parallel} \tilde{b}_r \rangle, \quad q_r = \langle \tilde{q}_{\parallel} \tilde{b}_r \rangle$$



Magnetic perturbation computational data are analyzed for recent BOUT runs for NSTX

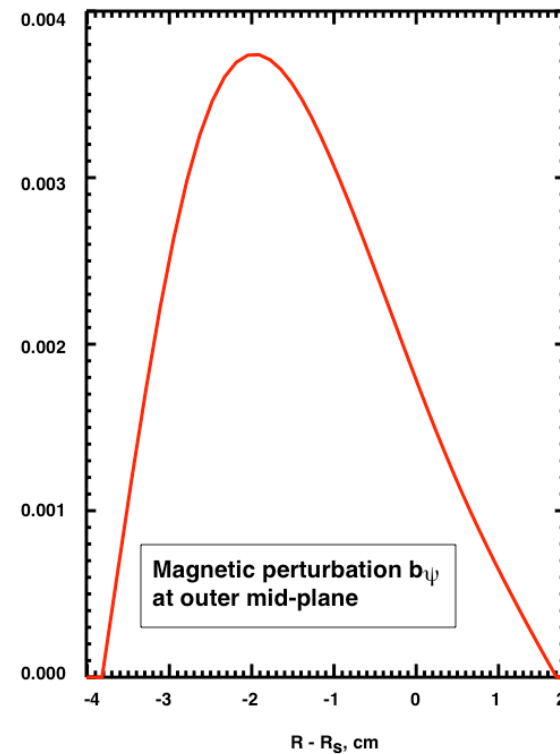
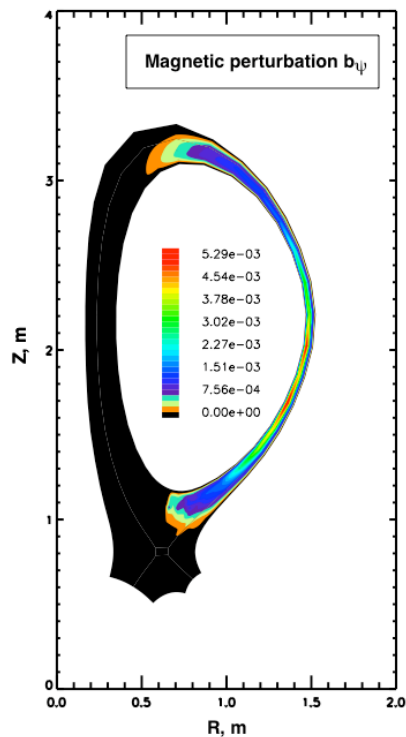
- Solving for $\tilde{N}_i, \tilde{V}_i, \tilde{j}_{\parallel}, \tilde{\phi}_i, \tilde{A}_{\parallel}$
- Using realistic parameters & geometry NSTX
- Matching the spatial scale of turbulence with GPI

BUT

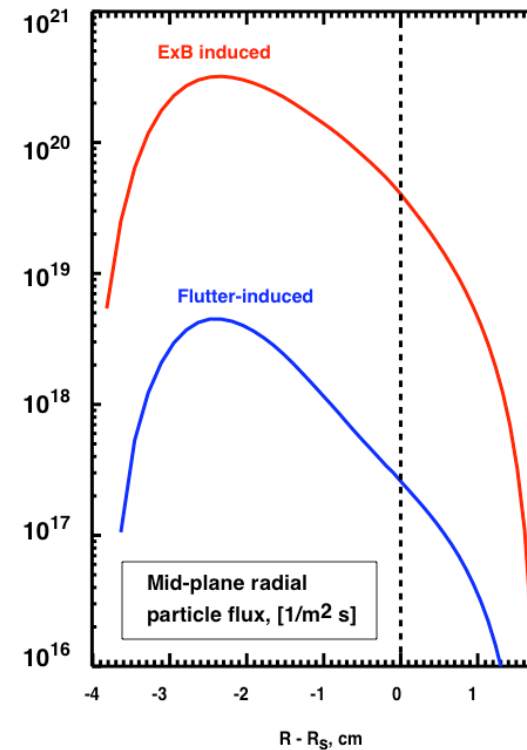
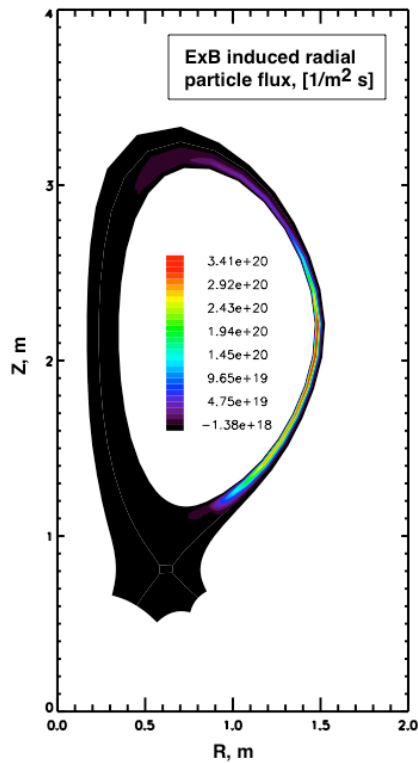
- Not including temperature fluctuations
- Assuming cold ions for background plasma
- Not all terms of the form $\tilde{b} \cdot \nabla$ are included



Magnetic perturbation is below 1% level, with ballooning-like poloidal dependence



Flutter induced transport is at least an order of magnitude below ExB flux, both are mainly at outer mid-plane



Conclusions

- **Based on recent BOUT runs for NSTX**
 - δb is below 1% level
 - magnetic flutter induced radial particle flux is at least an order of magnitude below ExB flux
- **These BOUT runs for NSTX still miss some physics, however present findings are consistent with previous BOUT experience,**



Backup slides

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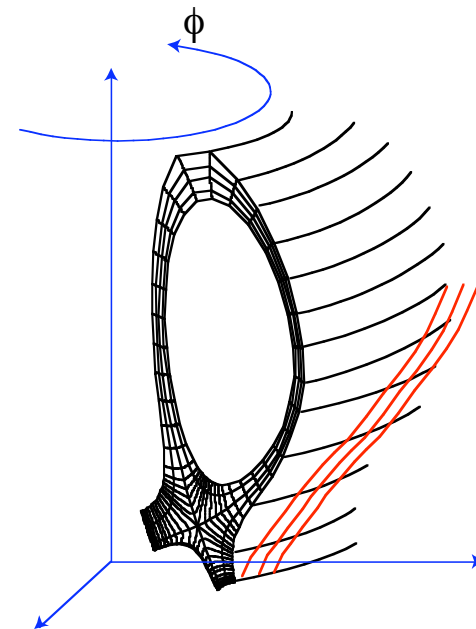


BOUT (BOUdary Turbulence) is a unique modeling tool for tokamak edge plasma turbulence*

- Fluid equations based on Braginskii equations for N_i , T_e , T_i , $V_{||e}$, $V_{||i}$, and ϖ
- Spatial discretization on 3D mesh in real geometry
- Time integration by implicit ODE solver PVODE
- Parallel implementation with MPI

**Xu et al, Contrib. Plasma Phys. 38, 158 (1998)*

Field following grid



Convective part of turbulent heat flux can be extracted from these BOUT runs, no conductive part present (no δT are used)

$$q_r = \frac{5}{2} \Gamma_r T_0$$

- **Convective heat flux is very small, to match the heat flux need to use T fluctuations - this is a high priority**

