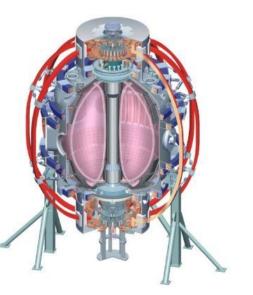


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## **Overview**

College W&M **Colorado Sch Mines** Columbia U CompX **General Atomics** INL **Johns Hopkins U** LANL LLNL Lodestar MIT **Nova Photonics** New York U **Old Dominion U** ORNL **PPPL** PSI **Princeton U Purdue U SNL** Think Tank, Inc. **UC Davis UC** Irvine UCLA UCSD **U** Colorado **U Illinois U** Maryland **U** Rochester **U** Washington **U Wisconsin** 



SPG

#### ASC TSG Meeting, 5/12/09



**Culham Sci Ctr U St. Andrews** York U Chubu U Fukui U Hiroshima U Hyogo U Kyoto U Kyushu U Kyushu Tokai U NIFS Niigata U **U** Tokyo JAEA Hebrew U loffe Inst **RRC Kurchatov Inst** TRINITI **KBSI** KAIST POSTECH ASIPP ENEA, Frascati CEA, Cadarache IPP, Jülich **IPP, Garching** ASCR, Czech Rep **U** Quebec

#### **Milestone Text**

"The high performance scenarios targeted in NSTX Upgrade and next-step ST devices are based on operating at lower Greenwald density fraction and/or lower collisionality than routinely accessed in NSTX. Collisionality plays a key role in ST energy confinement, non-inductive current drive, pedestal stability, RWM stability, and NTV rotation damping. Lower density and/or higher temperature is required to access lower nu\*. HHFW is a potential means of increasing electron temperature and reducing nu\*, and reduced fueling and/or Li pumping are effective and readily available tools for lowering nu\* through lower density. However, while D pumping from lithium has been observed, additional gas fueling is typically required to avoid plasma disruption during the current ramp and/or in the high beta phase of the highest performance (i.e. highest confinement, beta, non-inductive fraction, etc) plasmas of NSTX. The goal of this milestone is to identify the stability boundaries, characterize the underlying instabilities responsible for disruption at reduced density, and to develop means to avoid these disruptions. Possible methods for stability improvement include changes in current ramp-rate (li and q(r) evolution), H-mode transition timing, shape evolution, heating/beta evolution and control, optimized RWM control and error field correction, fueling control (SGI, shoulder injector), and optimized Li *pumping*. This milestone will also aid development of MISK and VALEN stability models and TRANSP and TSC integrated predictive models for NSTX Upgrade and next-step STs."



#### Where the Forum Left Off

- ASC:
  - Combined 3 proposals into a single XP lead by D. Mueller
    - Mueller, Sabbagh, Raman, Jarboe, Bell, Battaglia, Gerhardt,...
    - 3 days allocated.
    - 3 different strategies proposed.
  - Additional 0.5 days for early error field correction optimization.
    - Menard, et al.
    - Dedicated optimization of the OHxTF correction algorithm.
    - Probably needs to focus on a specific scenario.
- MHD:
  - Single proposal to study early rotating MHD that often disrupts shots.
    - S.P. Gerhardt, et al.
  - Was shifted to 1 day of cross-cutting and enabling.





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## "Controlling" MHD During the Plasma Start-Up

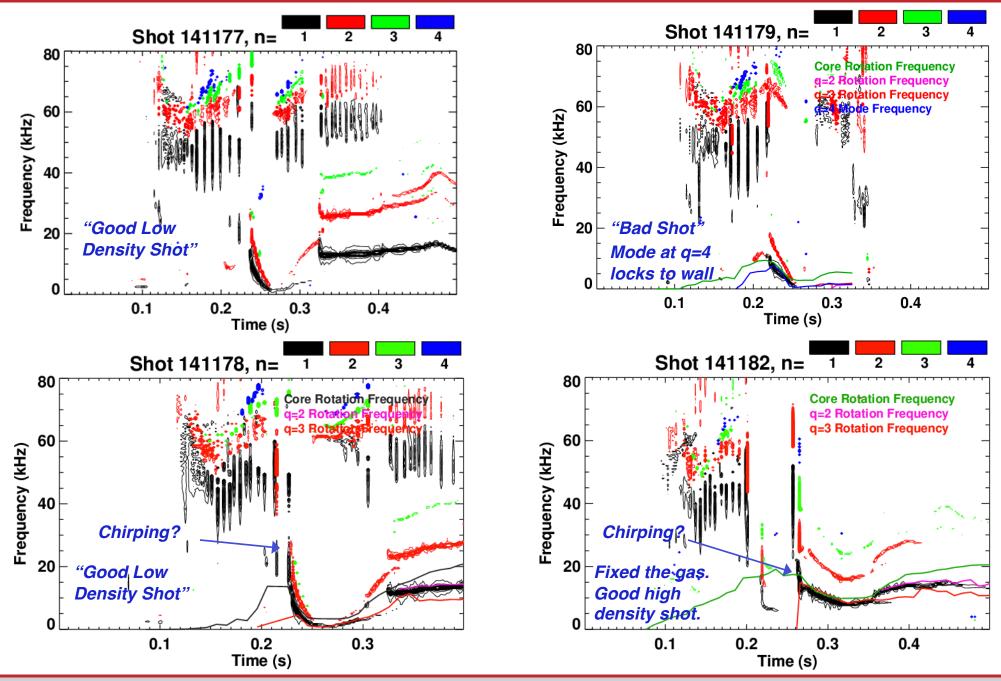


#### **Overview**

- "Low-density startup" should include surviving the entrance of q=4,3, & 2 into the H-mode plasma.
  - typically some few hundred msec. after SOFT.
- MHD associated with those surfaces often locks to the wall, leading to large  $\beta$  collapse or disruption.
  - Important for SGI shots, S. Sabbagh's reduced n<sub>e</sub> discharges
- Any of the current profile, plasma  $\beta$ , rotation, or rotation shear may impact the amplitude locking of these modes.
  - Also all the early EPM/TAE activity.
- We should optimize the discharge to ride through early modes with reduced density. Options include:
  - $-\beta$ : Feed forward or feedback control of NB heating.
  - Current profile: heating timing, ramp rate.
  - Rotation: H-mode timing.
- Results may be un-optimal for NSTX, but useful for NSTX upgrade.
  - Upgrade has 3x the solenoid flux, off-axis NBCD.



#### S. Sabbagh's Low n<sub>e</sub> Shots Failed When Rotating Modes Started to Lock





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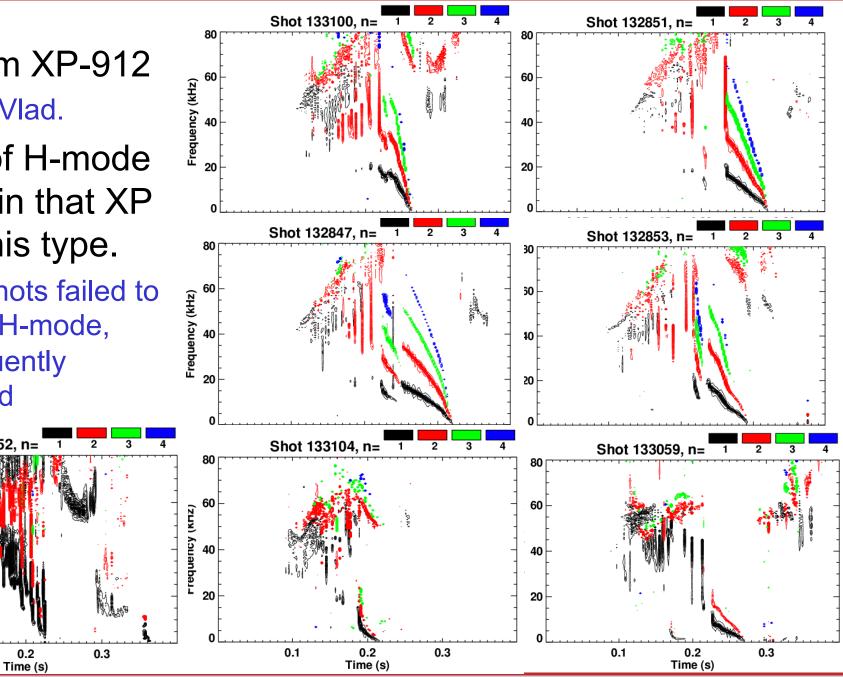
#### Same Physics Leads to Disruption in Many SGI **Optimization Discharges**

- Shots from XP-912 Thanks Vlad.
- Majority of H-mode "failures" in that XP were of this type.
  - Some shots failed to get into H-mode, consequently disrupted

Shot 133052, n=

0.1

0.2





80

60

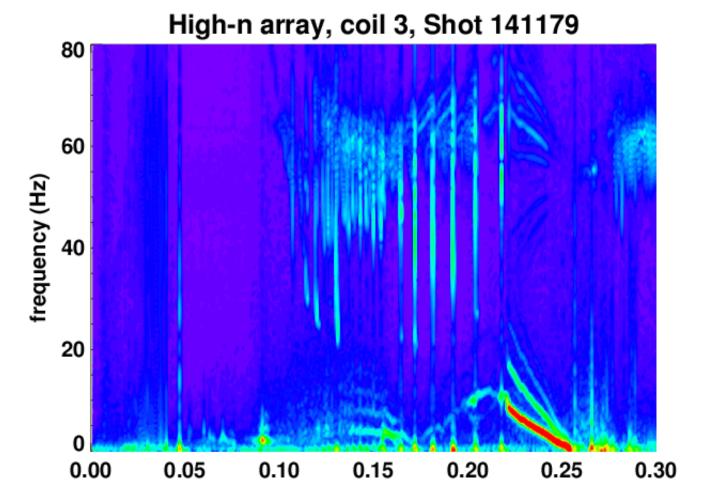
20

Frequency (kHz)

ASC TSG Meeting: Controlling Early MHD Modes (Gerhardt)

#### Appears that EPM May Have Triggered the q=4 Mode That Ultimately Locks During Sabbagh's Day

q=4 enters the plasma at ~0.19, but mode does not strike till 0.22. Mode strike coincident with large EMP.



# Does some triggering physics explain why previous shots didn't have large q=4 modes?



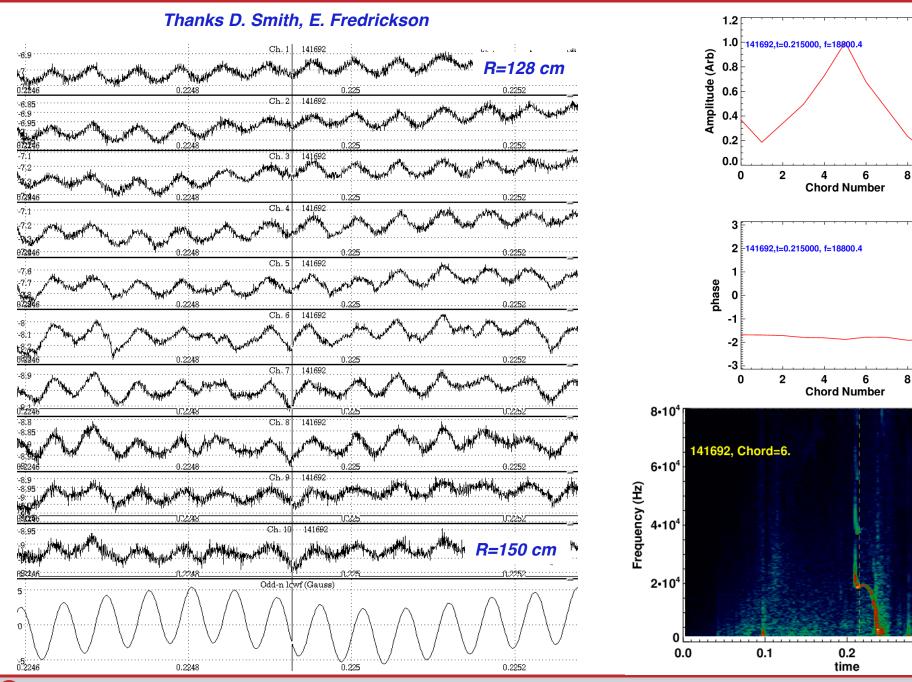
#### BES Shows That There is No Inversion Radius...Appears Not to be (double) Tearing...

10

10

0.3

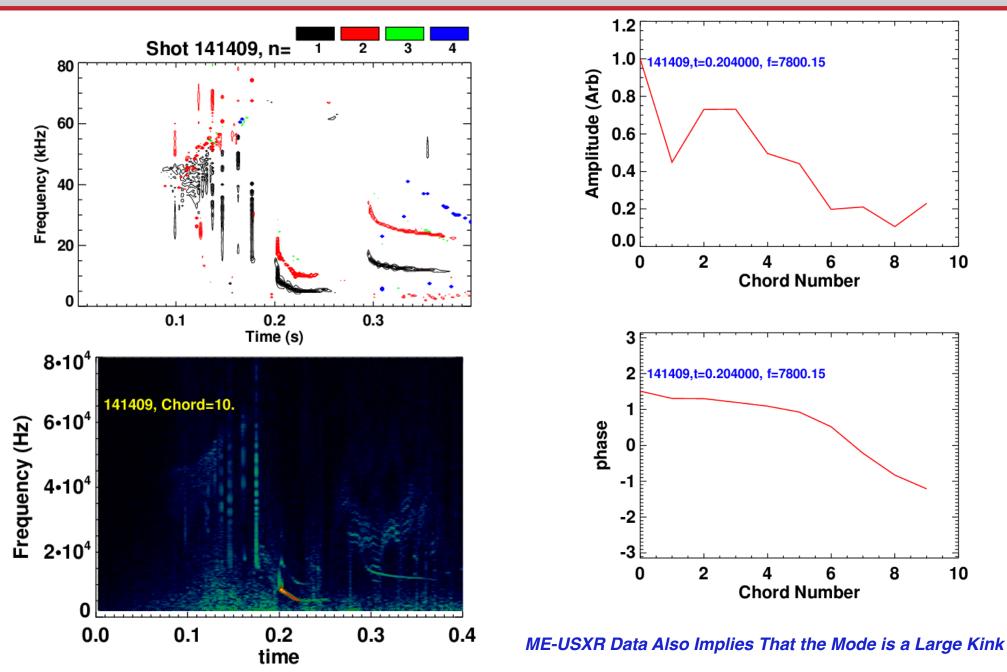
0.4



()) NSTX

ASC TSG Meeting: Controlling Early MHD Modes (Gerhardt)

#### BES Shows That the Modes Are Global, Without Sharp Inversion Layer





### **Preliminary Assessment for ME-USXR Corroborates Kink Explanation**

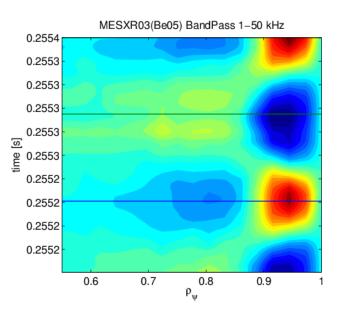
- JHU Multi-Energy **Tangential X-ray** diagnostic.
- Analysis by A. Bortolon.
- Again, key rational present before modestrike.

0.2

Time (s)

0.3

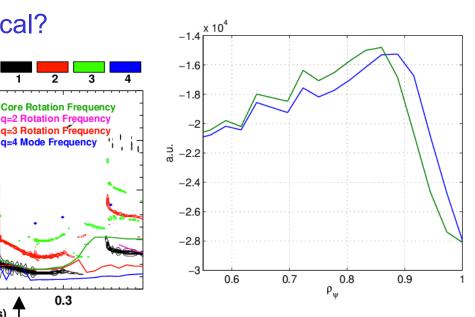
MESXR03(Be05) NSTX 142293 0.2554 0.2553 0.2553 0.2553 0.2553 time [s] 0.2553 0.2552 0.2552 0.2552 0.2552 0.6 0.7 0.8 0.9 ρ<sub>w</sub>

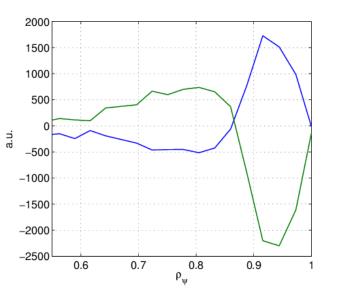




Shot 142293, n=

0.1







80

60

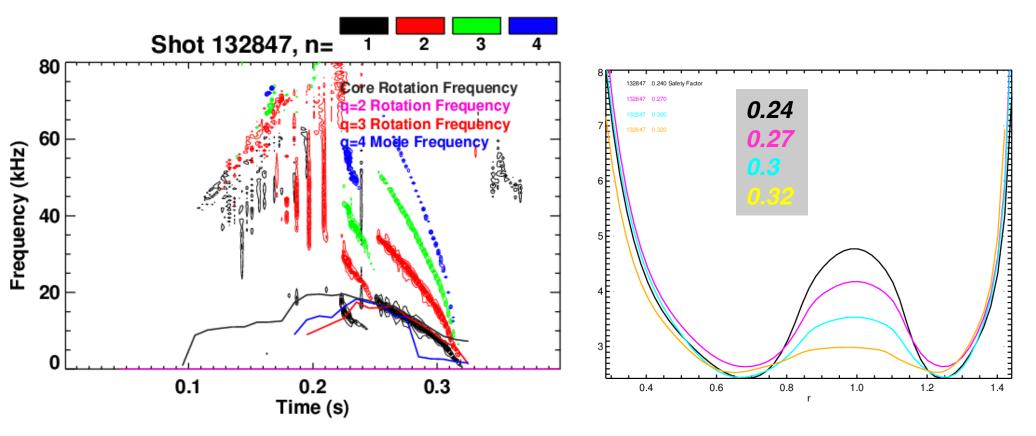
40

20

0

Frequency (kHz)

#### **Reversed Shear May Play a Role in Setting the Mode Size?**



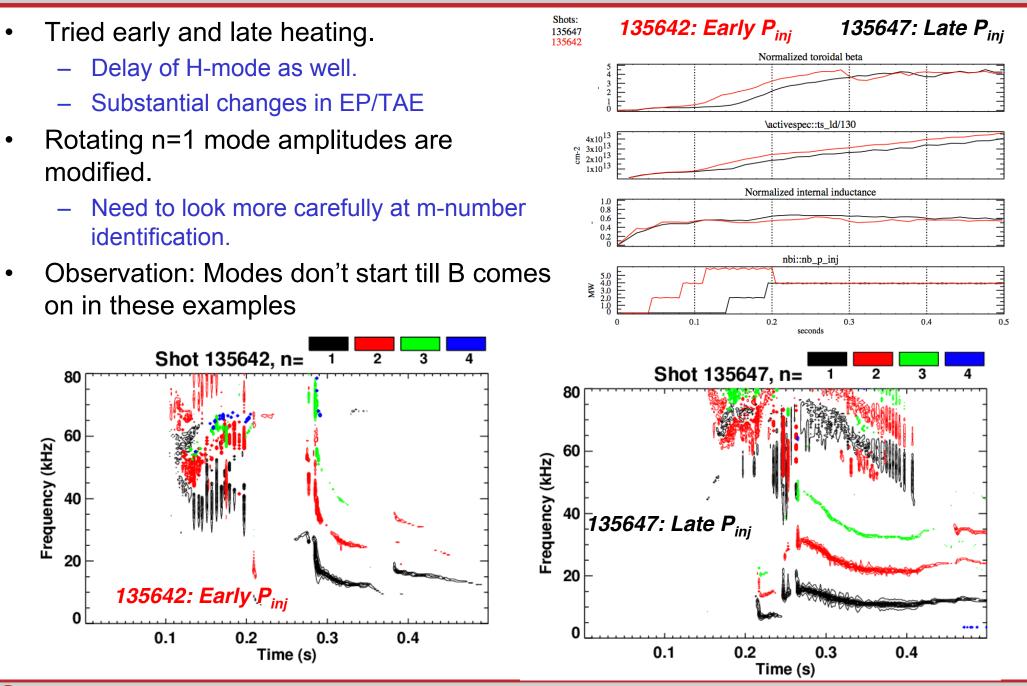
 Are larger modes more likely to coencide with flatter qprofiles?

- BES may tell us.

• Provides a mechanism for heating timing to impact evolution.



#### We Have Evidence That Changing the Heating Profile Can Impact the Evolution of These Modes.



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#### Possible shotlist needs refinement, but basics are...

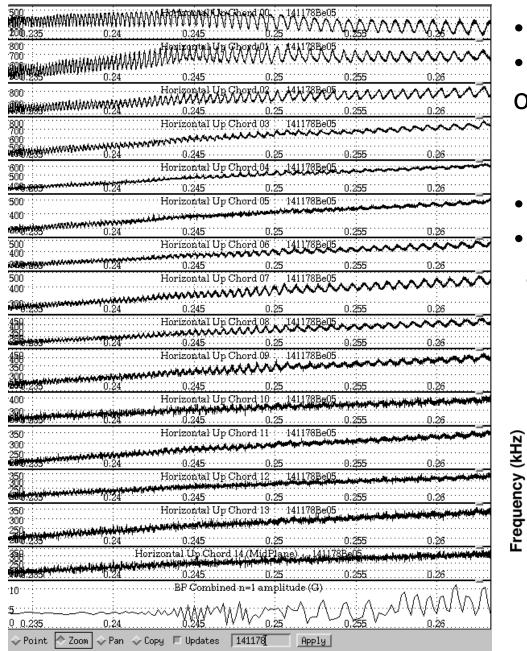
- Reload a scenario such as 141178 (S. Sabbagh's reduced density target).
  - Don't change early ramp-rate, shape evolution.
  - Develop robust mode evolution with reduced fuelling.
- Study the space of rotation,  $\beta$ , and  $n_e$  vs. mode dynamics.
  - Test low power early, delayed H-mode.
    - Reduced  $\beta$  and density with single source will allow surfaces in quickly, should have strong rotation.
    - Elimination of EPMs may be important...lower voltage pre-heating beam?
  - Study timing of  $\beta$  (P<sub>inj</sub>) ramp.
    - When is the earliest time that we can ramp  $\beta$  without EPMs and large modes.
    - Can we prevent too-rapid J evolution if we only heat strongly after q=3 enters?
  - Vary the ramp rate:
    - If we slow the ramp rate, can we prevent some unstable current profiles?
    - And eliminate some irreproducibility of the modes?
    - Target the q=2 & 3 surfaces entering just after the I<sub>OH</sub> zero crossings?
  - What is the impact of reduced input power?
    - Onset of "IREs" will be unacceptable.
    - Unsustainably high I<sub>i</sub> will be unacceptable.
    - Earlier onset of RWMs or the "late" rotating MHD is OK...
      - not in scope of XP
- Diagnostic considerations.
  - Need source A for important MSE and BES measurements.
    - Tangential USXR Camera, ME-USXR and Hup/Hdown for USXR mode identification.

#### Backup

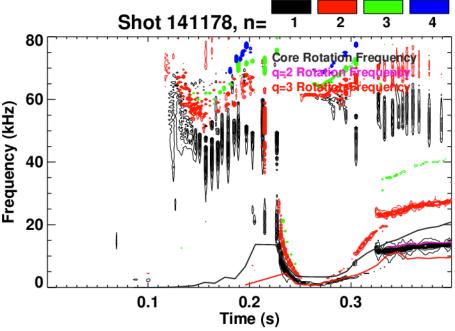


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#### Chirping q=3 Mode Settles With Inversion Radius In the Outer Plasma

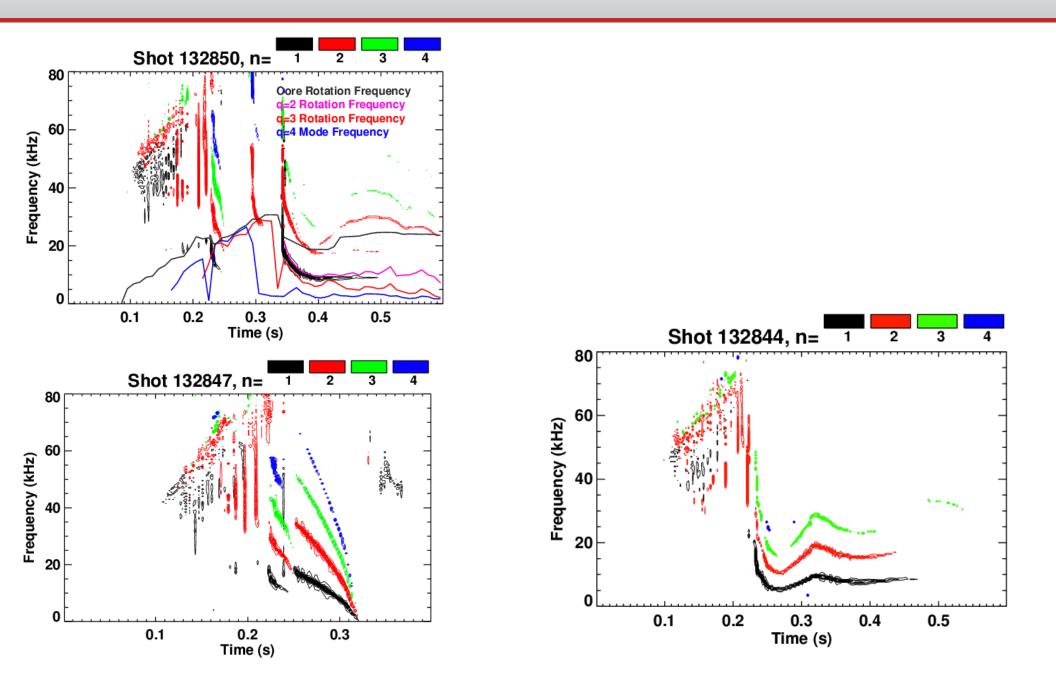


- Mode chirps down.
- Inversion layer develops in the outer plasma.
  - q=3 island?...need bit more analysis.
- Core kink-like component as well?
- ME-USXR may be quite useful analyzing this data.





#### **Difference Between "Good" and "Bad" SGI Shots**





#### Not Easy To Find Difference Between Shots Using 0D Parameters

