

# Gas-Puff-Imaging on Alcator C-Mod

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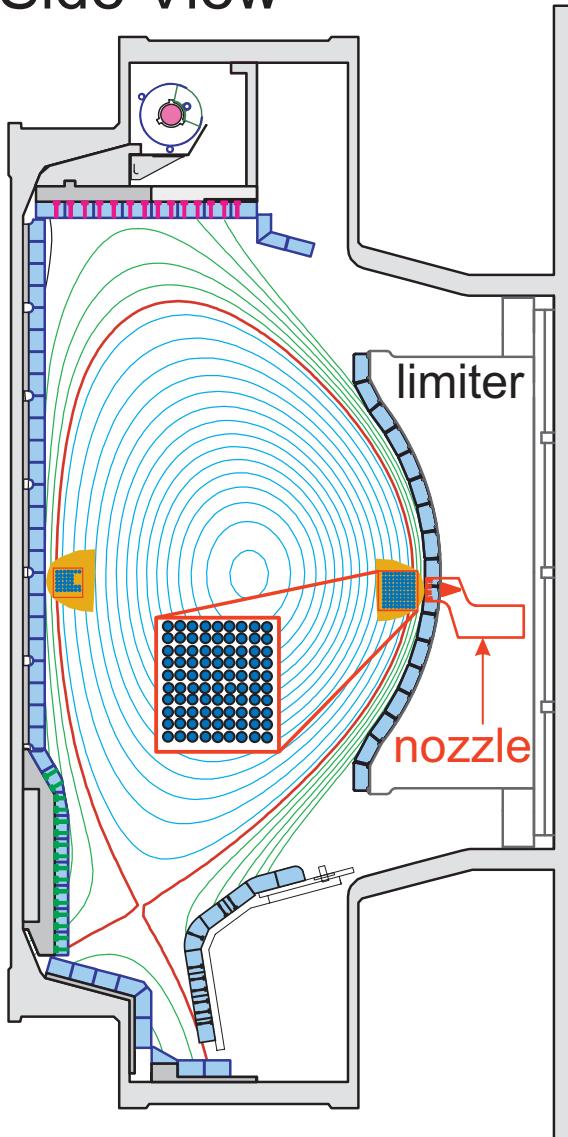
MIT - Plasma Science and Fusion Center

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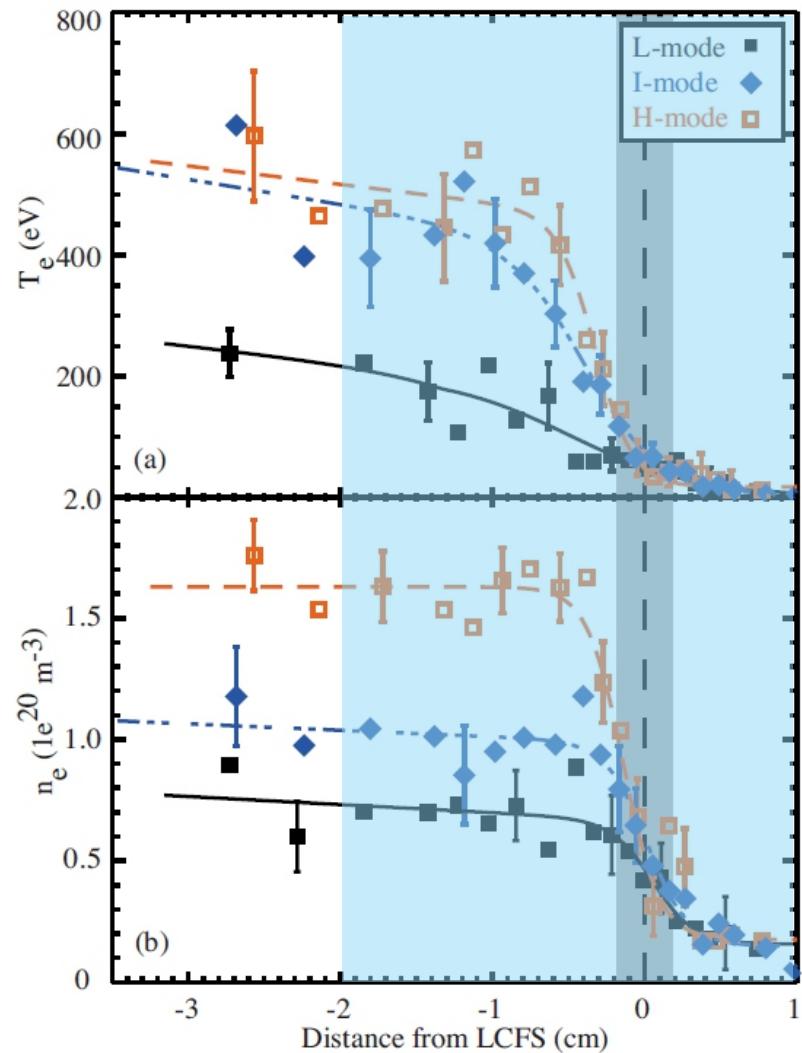
# Experimental setup

## Gas-Puff-Imaging diagnostics

Side View



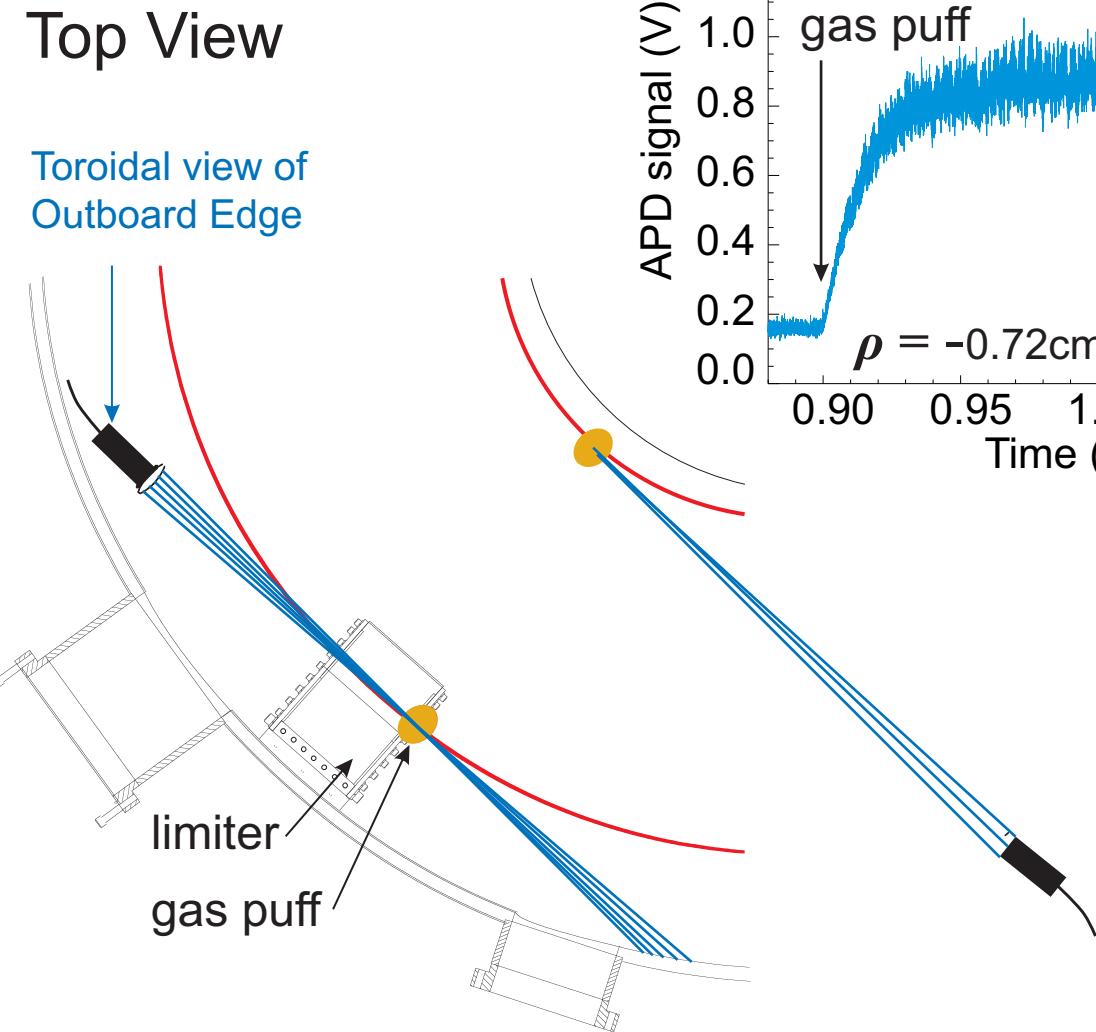
- 90 outboard channels cover ~ 4cm (radial) x 4.4cm (vertical)
- LoS horizontal
- outboard views coupled to APD arrays, sampled @ 2MHz
- 23 inboard views coupled to PIN photodiodes, @ 1MHz
- $k_{\text{pol}} < 7 \text{ cm}^{-1}$
- $\delta k_{\text{pol}} \sim 0.3 \text{ cm}^{-1}$



# Experimental setup

## Gas-Puff-Imaging diagnostics

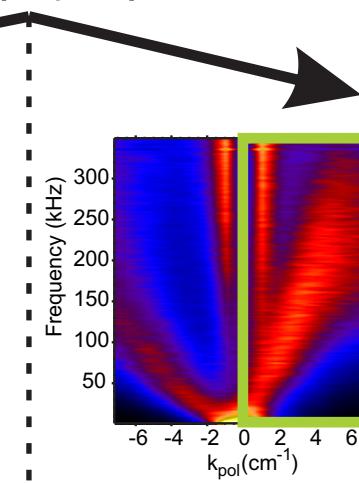
- gas puff injects neutral D<sub>2</sub>, He sensitive to  $n_e$ ,  $T_e$
- small toroidal extent (~5cm) allows localization:
  - ~4mm radial
  - ~5.5mm vertical
- puffing minority species affords high contrasts (x200 instead of x5)



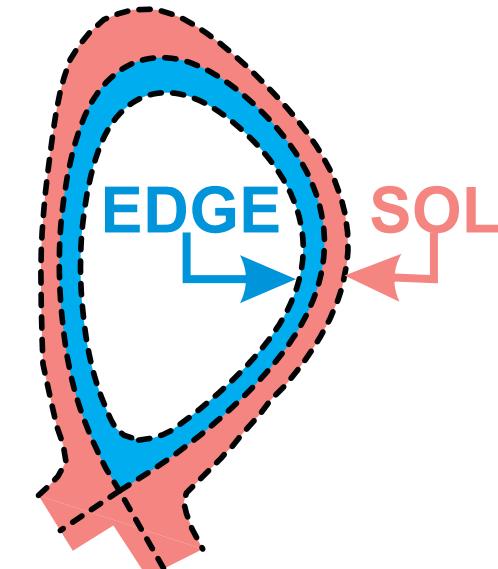
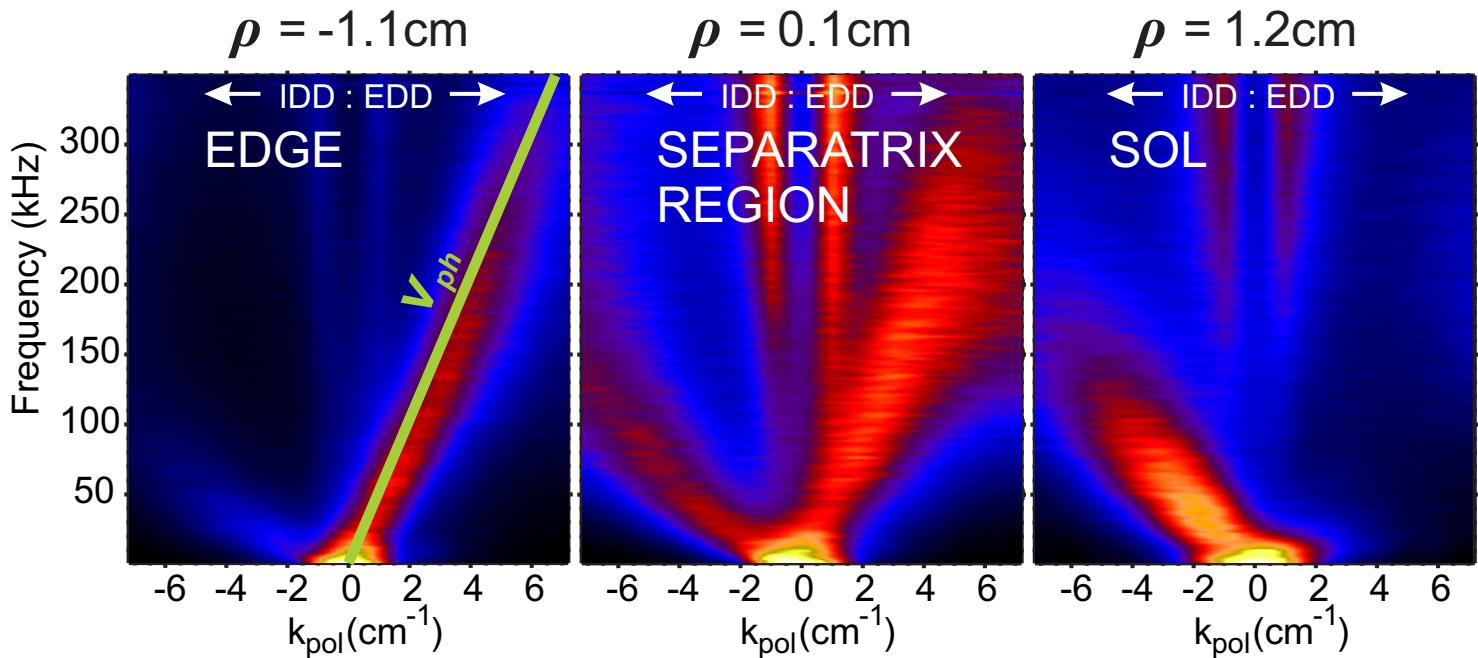
# Complete wavenumber frequency spectra decomposed into empirical dispersion and total power spectra

poloidal arrays + time history:  $S(k_\theta, \nu)$

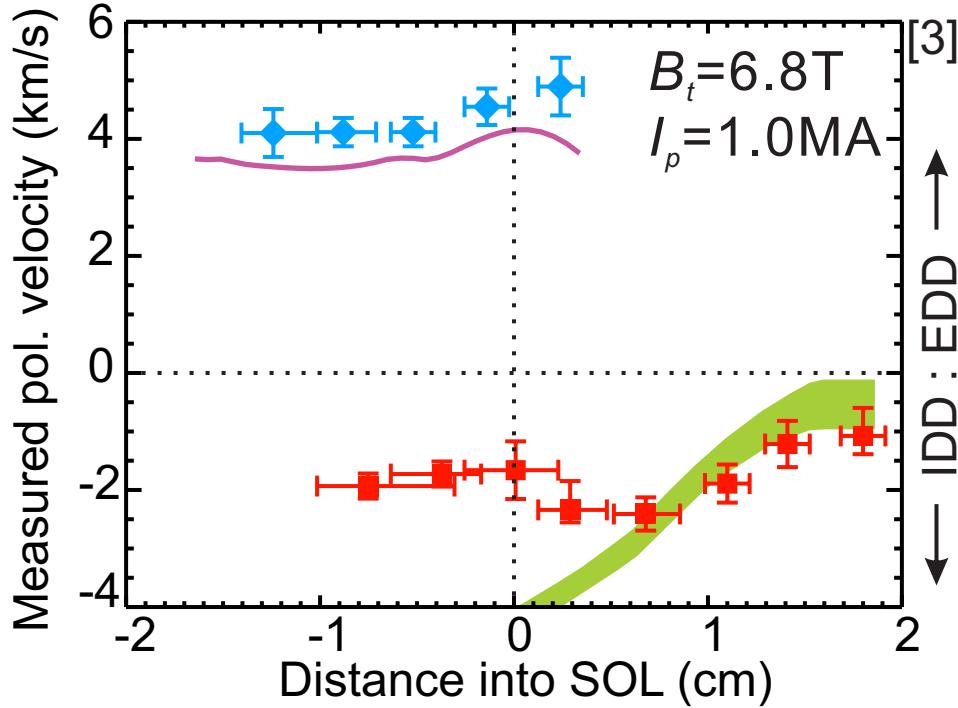
- $S(k_\theta|\nu) = S(k_\theta, \nu)/S(\nu)$
- Shows dispersion and dominant frequencies
- Dispersions are found to be very nearly linear:
- $k_\theta(\nu) \rightarrow \nu_{ph} = 2\pi\nu/k_\theta$



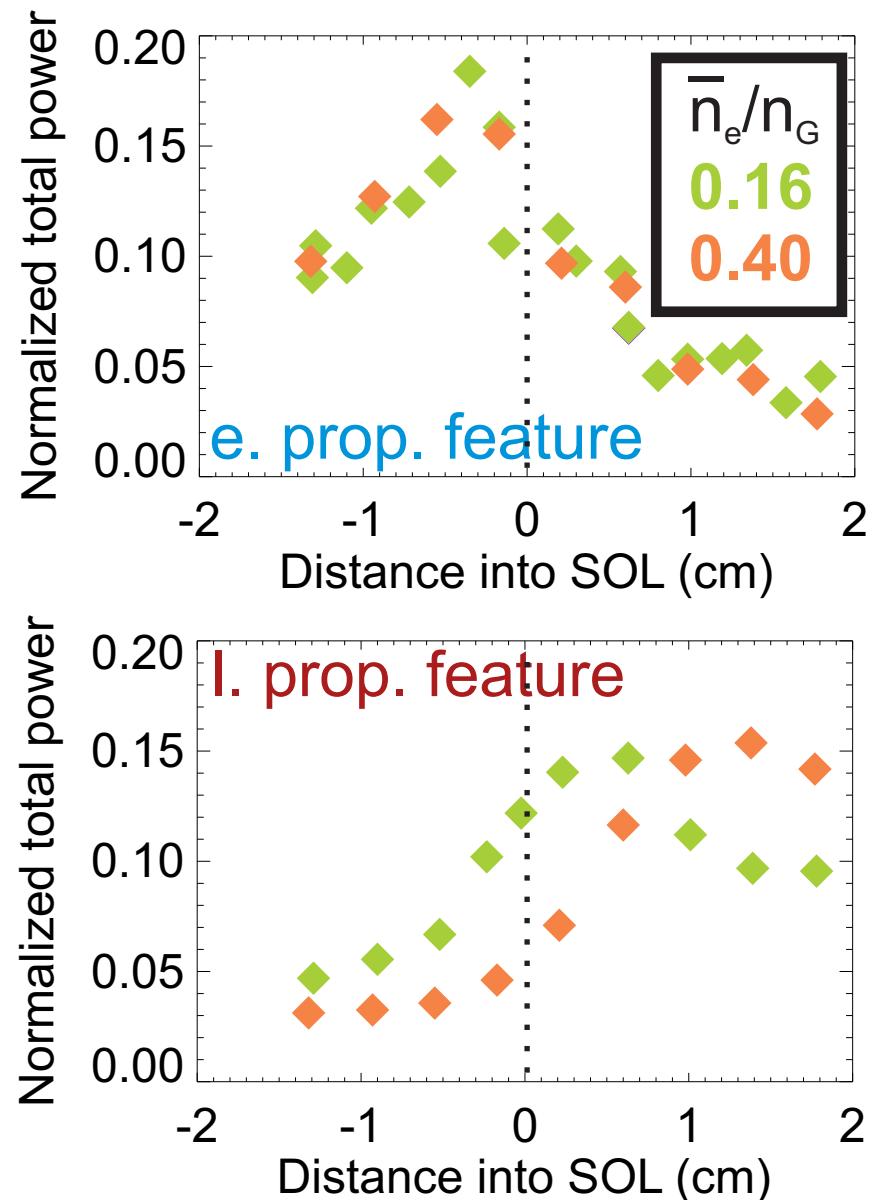
- $S(\nu) = S_{k<0}(\nu) + S_{k>0}(\nu)$
- Shows power spectra
- Characteristic scales can be seen



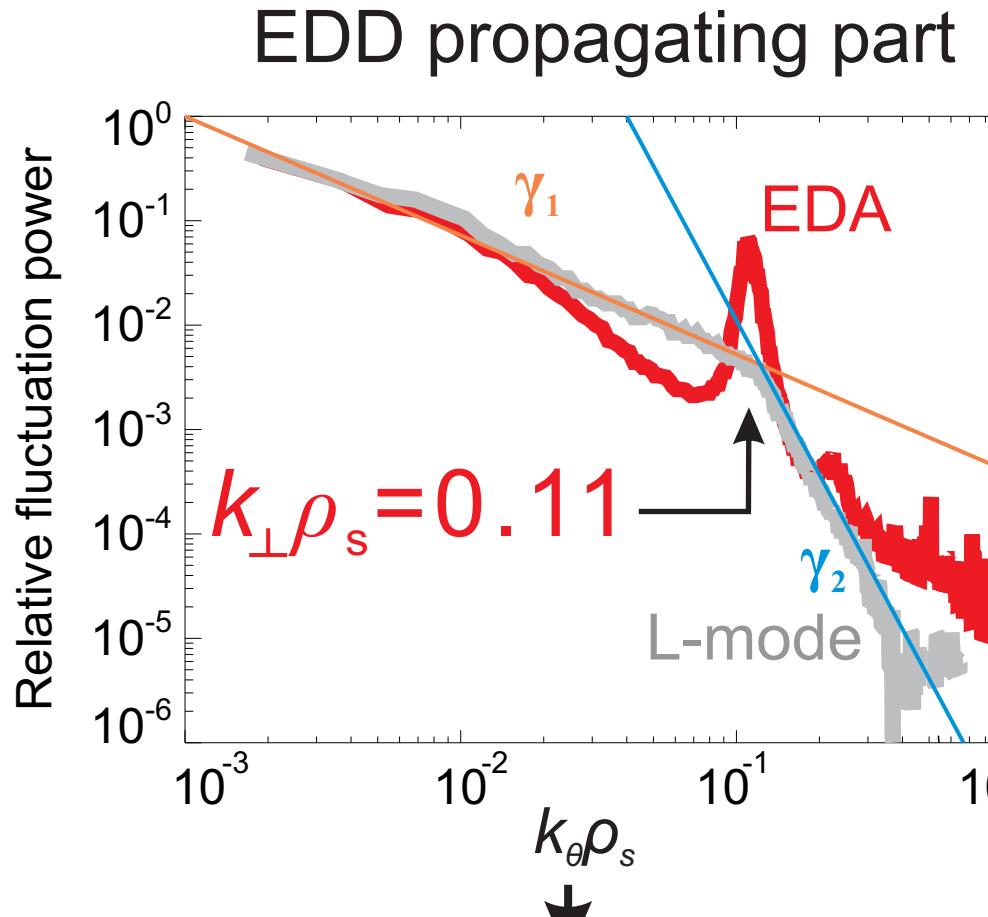
EDD velocity in edge is close to  $\omega_*/k_{\text{pol}}$   
 far SOL IDD velocity matches ExB (expected for blobs)



- **EDD propagation in the edge** is close to the **electron diamagnetic flow velocity estimated from TS profile**
- Poloidal velocity in the far SOL where **blobs** are seen matches **ExB velocity estimated from probe measurements**
- Overlap region between the two features



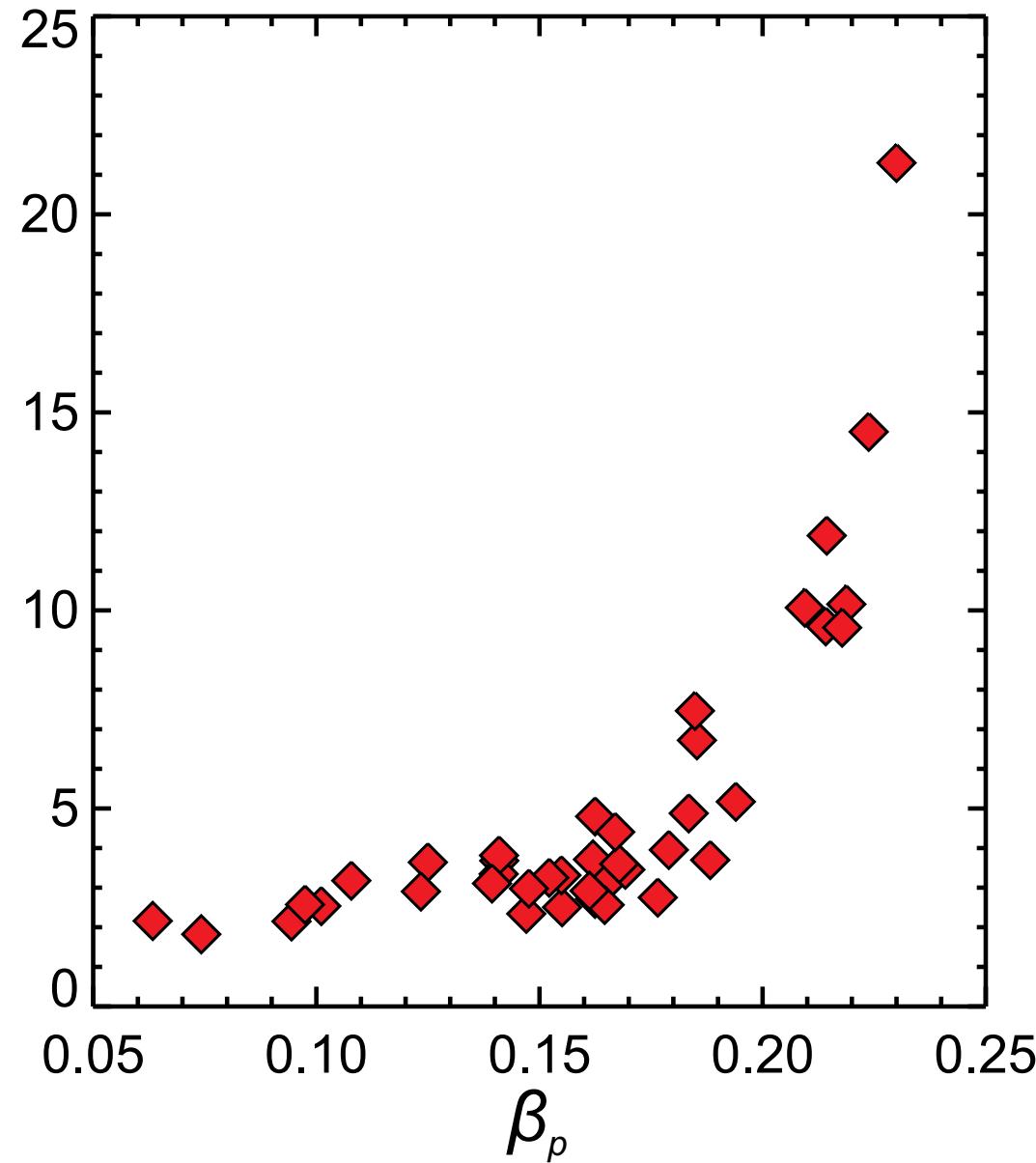
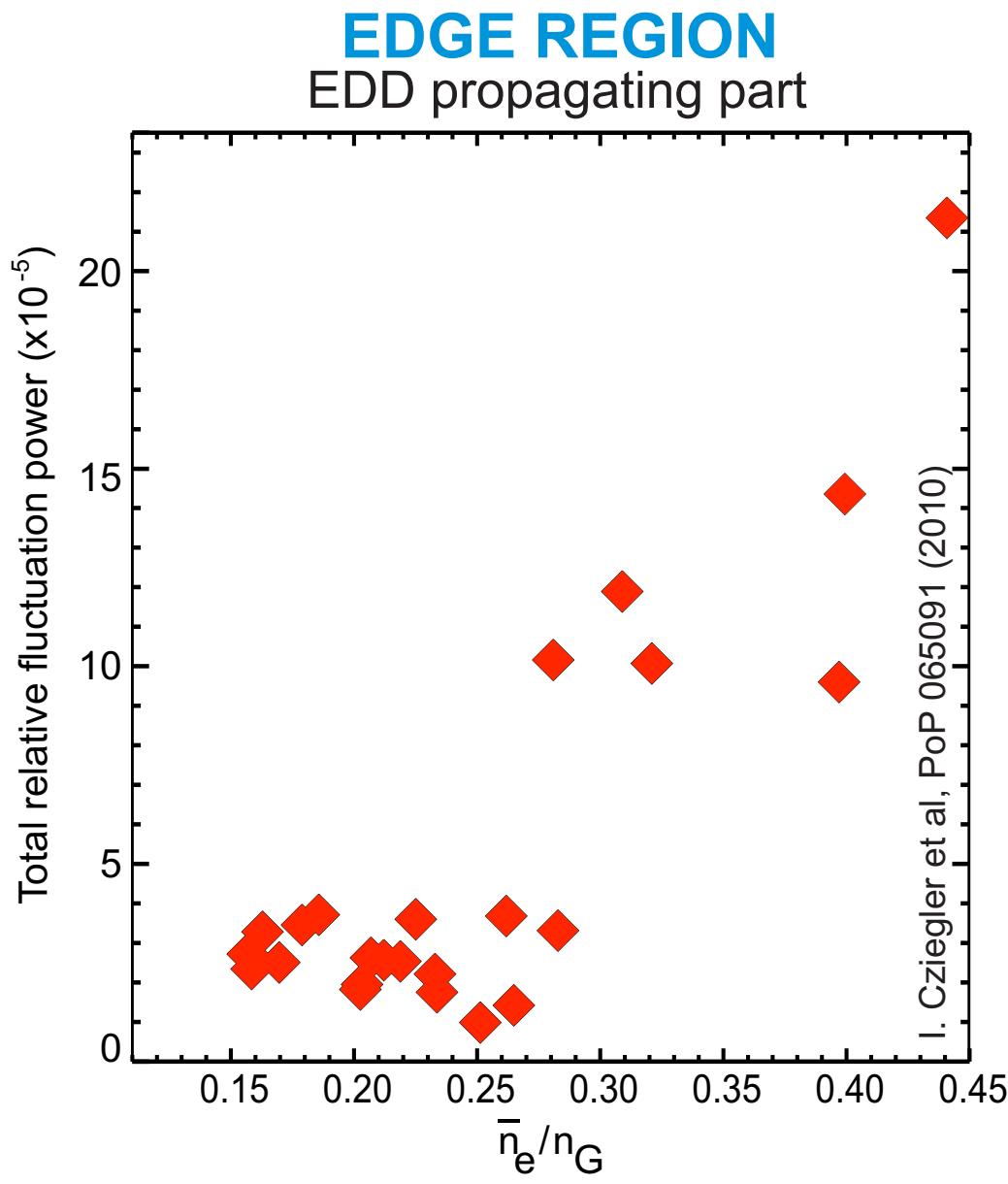
Wavenumber filtered spectra  $S_{k>0}(\nu)$  can yield information on underlying dynamics



isotope tests would be interesting

- Edge spectra often power laws
- Spectral indices are fairly reproducible - spectral transfer?
- Break-in-slope  $\nu^c$  may indicate:
  - dissipation scale
  - scale of energy input
- In the latter case:
  - $\gamma_1$  indicates “inverse cascade”
  - $\gamma_2$  indicates “forward cascade”

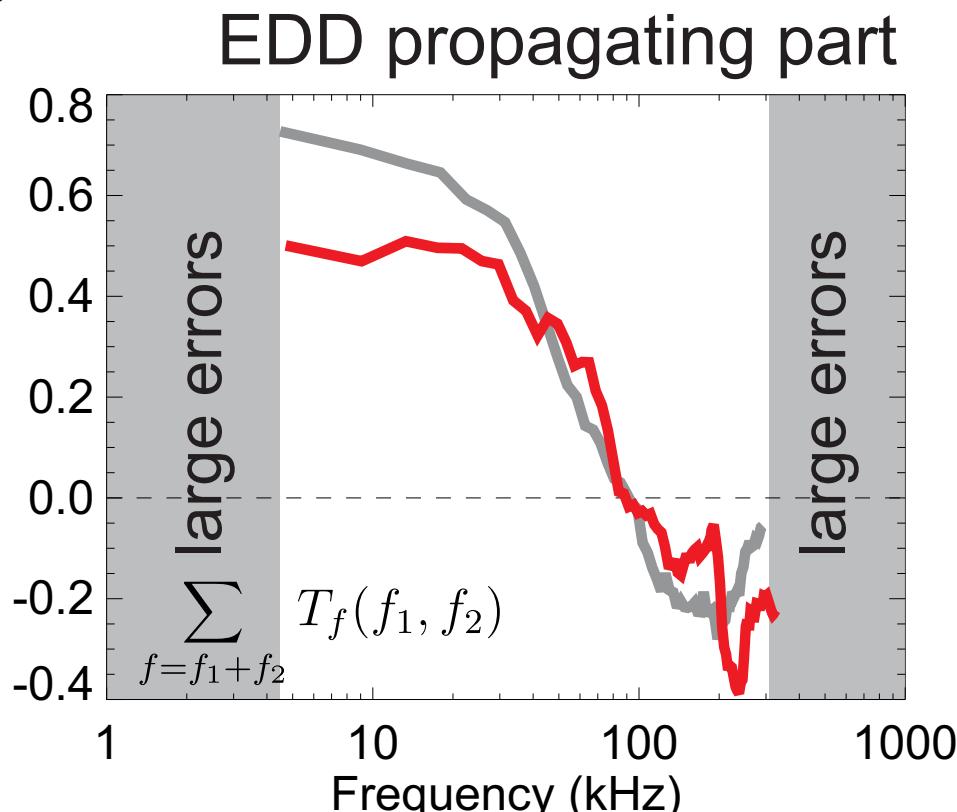
Total power of EDD turbulence observed to increase as a function of  $\beta_p$ , or Greenwald fraction, consistent with turbulence driven transport



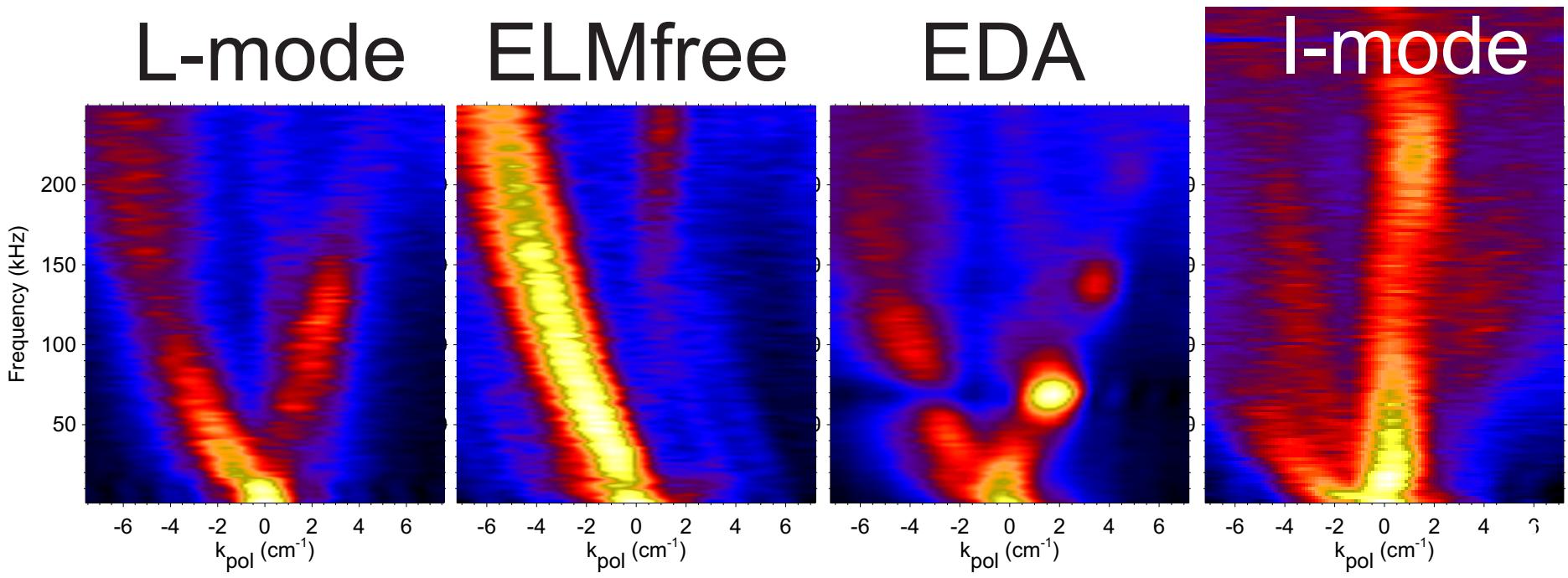
Low noise levels allow bispectral analysis over a large spectral region: energy input at the same scale for H-mode and L-mode

From  $S(k_\theta|\nu)$  we can move to spatial separation rather than time using  $k_\theta(\nu) \rightarrow \nu_{ph} = 2\pi\nu/k_\theta$ , and the appropriate Galilean transformation for the velocities.

viewing chans  $\left\{ \begin{array}{|c|} \hline \text{ } \\ \hline \text{ } \\ \hline \text{ } \\ \hline \end{array} \right\} \Delta z \left\{ \begin{array}{|c|} \hline \text{ } \\ \hline \text{ } \\ \hline \text{ } \\ \hline \end{array} \right\} S(k_\theta|\nu)$   $\longrightarrow$  time average instead of ensemble average



All confinement regimes found to have “fingerprint” turbulence patterns in the pedestal - JFT-2M shape, ELM regime is still missing



Due to the vertical arrays of views spanning the pedestal, ELM precursors are expected to be visible and their poloidal wavenumbers to be resolved - what characteristics of the models (P-B, KBM, etc) are testable?