

Stochastic heating of thermal ions by ETG streamer turbulence

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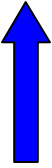
NSTX Motivation 1: Rapid electron transport

- **NSTX χ_e typically 5-20m²/s across profile**
- Simple mixing length estimate for ETG (note $v_{te}^2 \equiv T_e/m_e$) \Rightarrow
 $\chi_e \sim \rho_e^2 v_{te} / L_{Te} \equiv D_{ETG} = 0.075 [T_e(\text{keV})]^{3/2} / B^2 L_{Te}$
- For NSTX near $r/a = 0.5$:
 - $T_e = 0.8 \text{keV}$, $B = 0.3 \text{T}$, $L_{Te} = 0.4 \text{m}$, $D_{ETG} = 1.5 \text{m}^2/\text{s}$
 - For $B = 0.45 \text{T}$, $D_{ETG} = 0.7 \text{m}^2/\text{s}$
- ETG χ_e with streamers $\sim 10\text{-}20 D_{ETG}$ using s- α equilibria at higher aspect ratio (Dorland et al., PRL, December 2000, 5579)
- Crudely applied to NSTX, **streamer ETG** $\Rightarrow \chi_e = 10\text{-}20 \text{m}^2/\text{s}$

Motivation 2: with strong rotation, χ_i can be ≤ 0

- χ_i can be less than neoclassical between $r/a = 0.4$ and 0.8

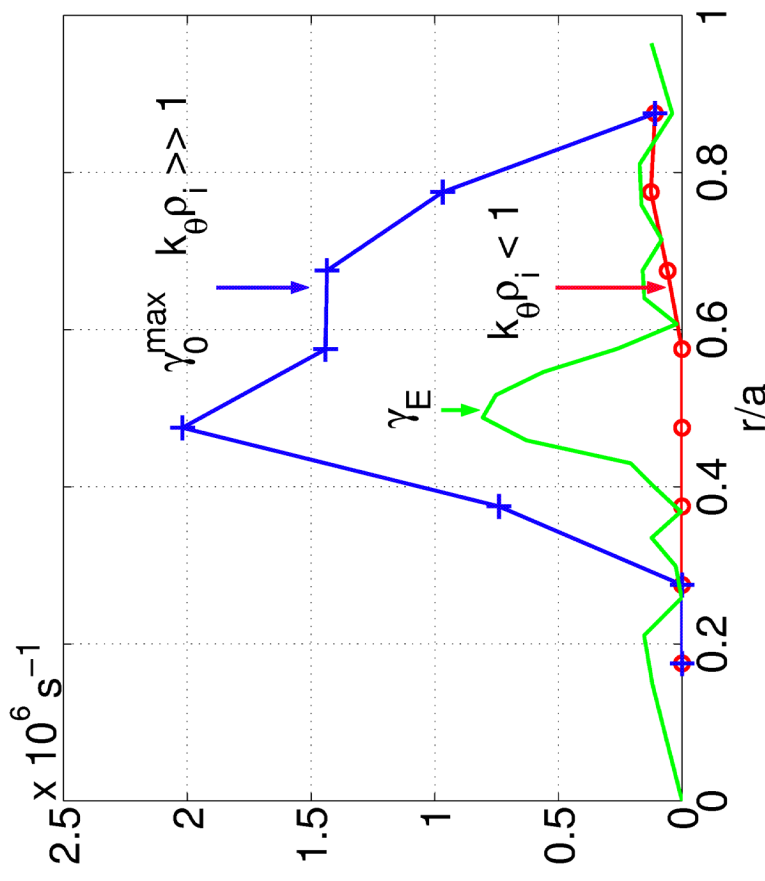
- Only possible with ion heating anomaly or diagnostic error

- Linear ETG simulations predict robust instability in same minor radial region 

- Can ETG streamers heat ions through dissipation?

- If so, what are the thermodynamic implications?

From GS2 – C. Bourdelle



#104001 at 0.28 s

Estimate of ETG streamer dissipation

- First, estimate normalized fluctuation amplitude $\Phi = e\phi/T_e$
 - Balance KH against streamer growth:

$$\Phi = \rho_e/L_{Te} \Gamma^{-1} \tau / (1+\tau) \gamma / k_{\perp}^4 \quad (\text{See Dorland PRL})$$

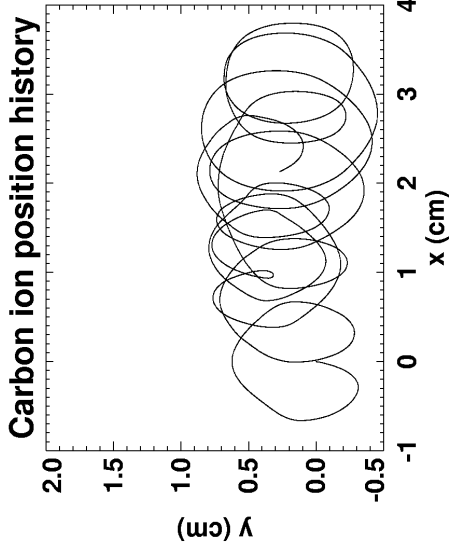
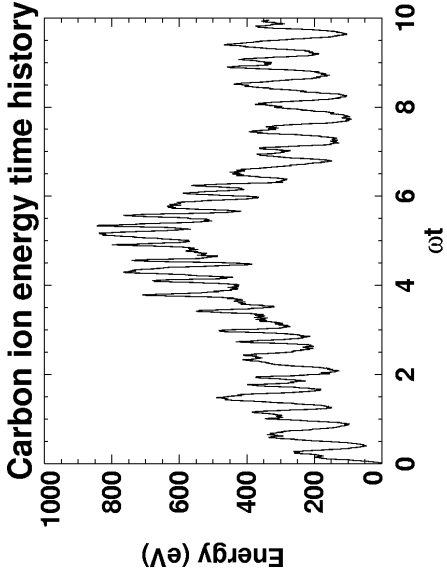
- $\Gamma =$ KH growth rate coefficient $\approx 0.2\text{-}0.3$
- $\tau = Z_{\text{eff}} T_e / T_i \approx 1\text{-}2$
- $\gamma / k_{\perp}^4 = 10\text{-}30$ for positive shear tokamaks
- For these coefficients, $T_e = 800\text{eV}$, and NSTX B-fields, $\Phi = 0.5\text{-}5\% \Rightarrow \phi \approx 4\text{-}40\text{V}$ assuming $L_{Te}=0.4\text{m}$
- Saturated turbulence predicted to have $k_{\theta}\rho_e \approx 0.2 \Rightarrow$
large fine-scale E-fields $\approx 10\text{-}20\text{kV/m}$

Details of ion heating simulations

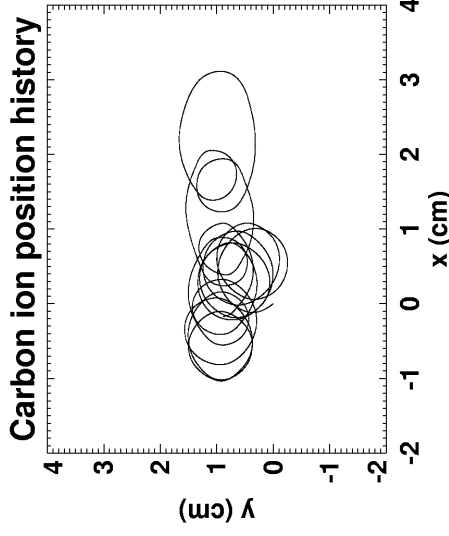
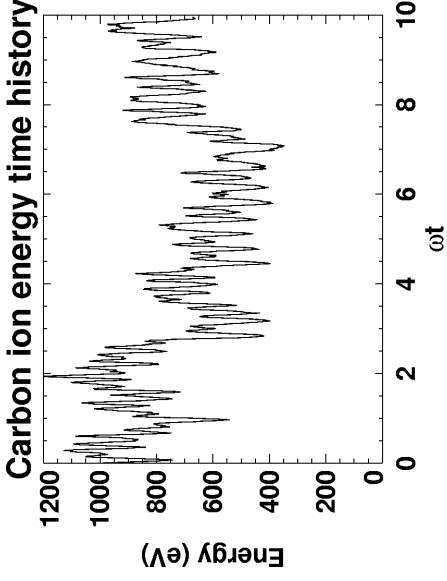
- Solve orbit equation: $m\mathbf{d}\mathbf{v}/dt = \mathbf{q}\mathbf{v}\times\mathbf{B} + \mathbf{q}\mathbf{E} + \mathbf{F}_{\text{collision}}$ using full test-particle collision operator assuming background Maxwellian ion distribution *with constant T_i*
 - For long integration times, any test-particle heating from wave will be dissipated by the background ions.
 - Equivalently, **the test particle temperature is always close to background T_i** \Rightarrow many runs needed to determine dT_i/dt vs. T_i
- Simulation parameters:
 - $T_e = T_i = 1\text{keV}$
 - $B = 0.3\text{T}$, $\phi = 20\text{V} \Rightarrow e\phi/T_e = 2\%$
 - $n_e = 4 \times 10^{19}\text{m}^{-3}$ in deuterium, $Z_{\text{eff}} = 2$, Carbon impurity ($Z=6$)
 - $k_{\theta}\rho_e = 0.2 \Rightarrow \alpha \approx 3$
 - $v = 0.11$ ($f \approx 250\text{kHz}$)

Ion-ion collisions cause significant energy diffusion

Initial $E = 1\text{eV}$

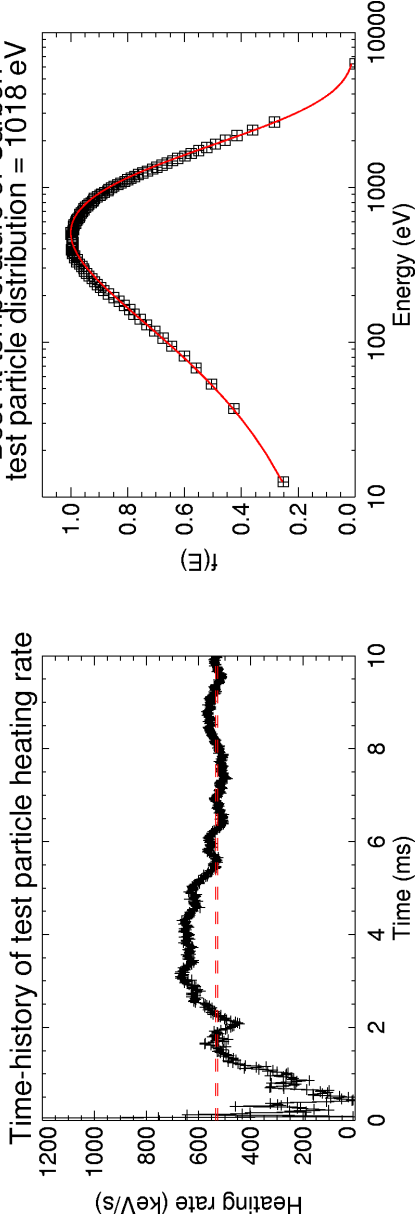
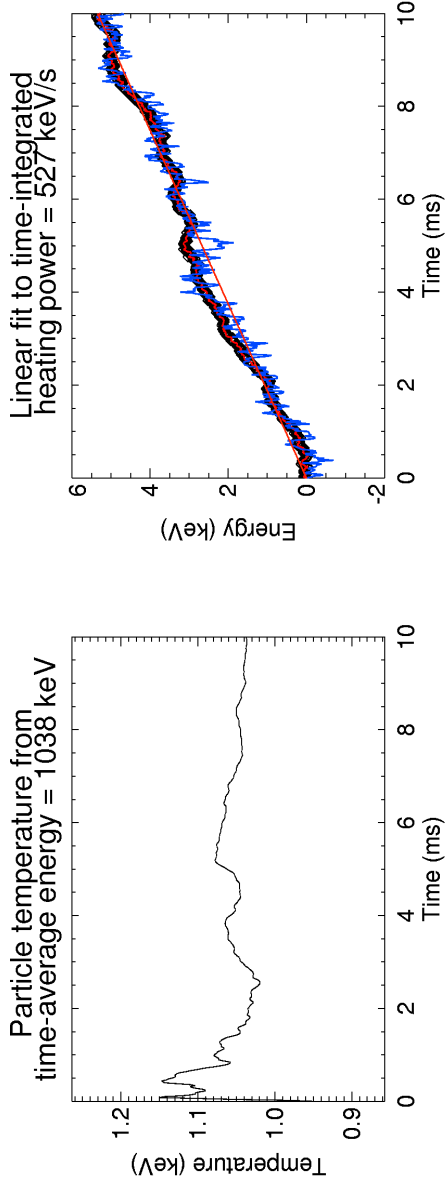


Initial $E = 1\text{keV}$



- Collisions can readily move ions in and out of regions of orbit stochasticity
- Collisionality can change mean energy significantly over few wave periods \Rightarrow
- **Collisions should be strong enough to cause significant decorrelation and heating**

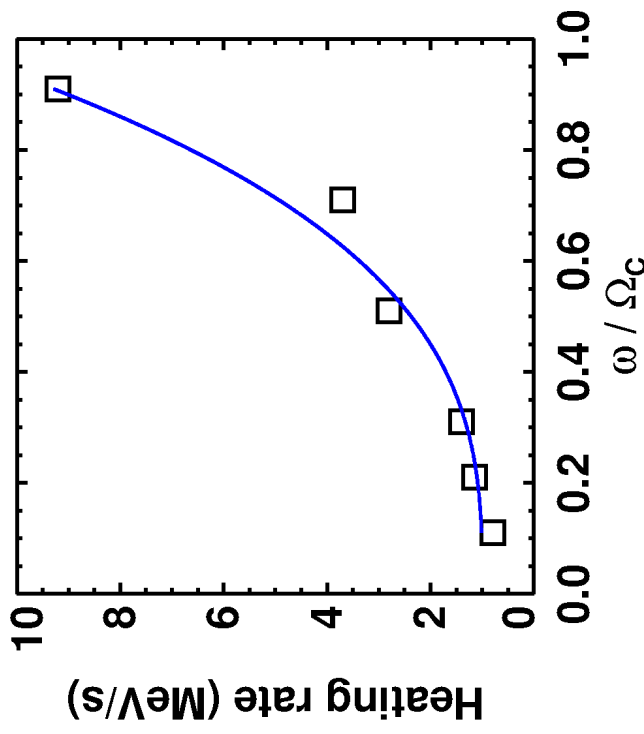
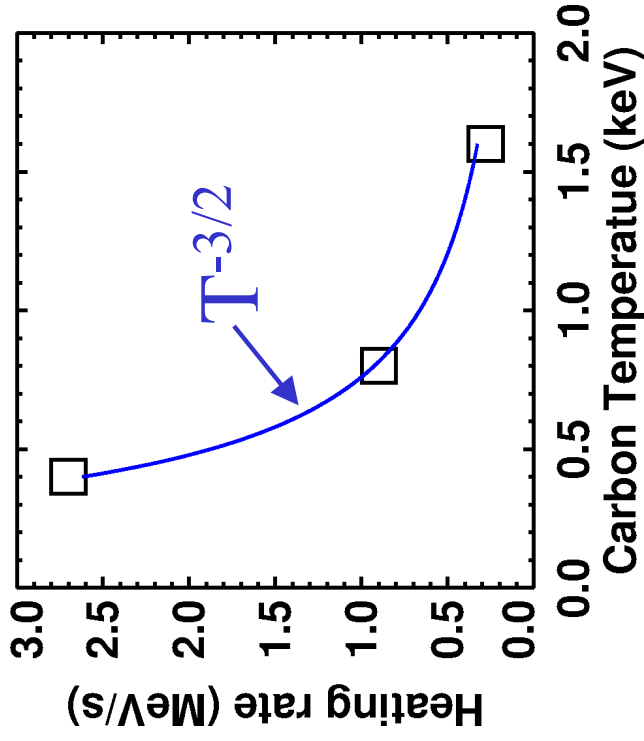
Carbon test-particle energy distribution function becomes Maxwellian with temperature expected from heating



- Fully converged Carbon heating rate is $\approx 600\text{keV/s} \Rightarrow 130\text{kW/m}^3$

Parametric scaling of heating rates

- Heating proportional to $v_{\text{collision}}$
- Heating susceptible to other de-correlation processes
- Heating rate increases with ω
 - Does stochastic energy increase?
- Cyclotron resonance evident



- Presently working on determining role of non-zero k_{radial}
 - Initial indications are that higher k_r results in weaker dissipation

Experimental Considerations:

- Density/potential fluctuation characteristics:
 - Wavelength: $k_{\theta}\rho_e \approx 0.2$, $k_r\rho_e \ll 1$
 - Amplitude: 0.5-2%, maximum near $\frac{1}{2}$ radius, $\sim 1/B$
 - Frequency: broadband $f \sim 100$ -400kHz
- Possible techniques for understanding physics:
 - Strong reverse shear should suppress ETG
 - Z_{eff} would influence ETG heating power density
 - Higher B-field might reduce heating anomaly
 - Heat electrons in regimes w/ NC ions, but w/o beams
 - Can V_p driven E_r shear in NSTX stabilize ITG ?
 - If so, NBI followed by HHFW might eliminate CAE in hot ion regime