

Application of POCA for NTV Analysis and Field Line Tracing in the Non-Axisymmetric Magnetic Perturbations

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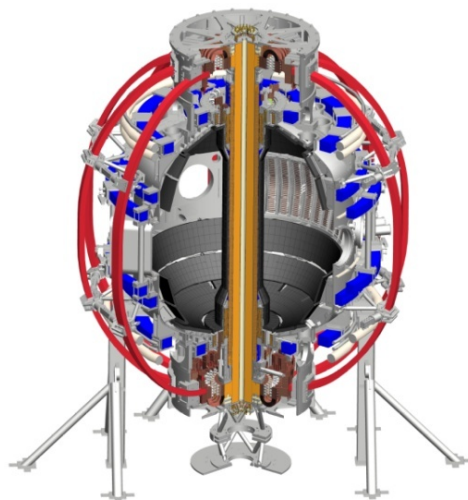
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Background

- Neoclassical Toroidal Viscosity (NTV) plays an important role for control of plasma rotation, stability and performance in perturbed tokamaks
 - Non-axisymmetric magnetic perturbations (NAMPs) can fundamentally change the neoclassical transport by distorting particle orbits on deformed or broken flux surfaces
 - NAMPs drive significant magnetic braking by NTV, thus change plasma rotation impacting on tokamak performance
 - NAMPs are important control elements to actively stabilize locked modes, edge localized modes, and resistive wall modes
- Analytic NTV theories have been successful to show fundamental physics, but lack of quantitative validations
 - $1/\nu$ regime in the high collisionality [K.C. Shaing, POP (2003)], $\nu_{\perp}\nu^{1/2}$ regime in the low collisionality [K.C. Shaing, POP (2008)], superbanana plateau regime in the low ExB precession [K.C. Shaing, PPCF (2009)]
 - Neoclassical offset rotation in the counter- I_p direction [A.J. Cole, PRL (2007)]
 - Approximations by large-aspect-ratio, zero-orbit-width, simplified collision operator, only trapped particles, and/or regime separation
 - Missing physics for rotational resonances

Motivation to δf particle code for NTV calculation

- A δf particle code has been developed to calculate neoclassical transport in perturbed tokamaks
 - Verify basic NTV physics such as quadratic δB dependency, BH resonance, etc
 - Analyze and predict NTV in magnetic braking experiments, using δB spectrum from 3D perturbed equilibrium solver (IPEC)
- Bounce-harmonic (BH) resonance enhancing NTV [K. Kim, PRL (2013)]
 - Predicted by [Linsker and Boozer, PFL (1982)], [H.E. Mynick, NF (1986)], and reformulated by [J.-K. Park, PRL (2009)]
 - Only mechanism to enhance the NTV in the fast rotating plasma
 - Orbit closing by resonant ExB is a critical physics to explain the BH enhanced NTV, which can be captured only by particle simulation w/o approximations
 - First numerical verification of enhancement mechanism of NTV by BH resonances
 - Recent KSTAR experiments of strong magnetic braking by BH resonances
- New capability for field line tracing considering plasma response
 - Enable field line tracing considering ideal plasma response compared to vacuum field

Outline

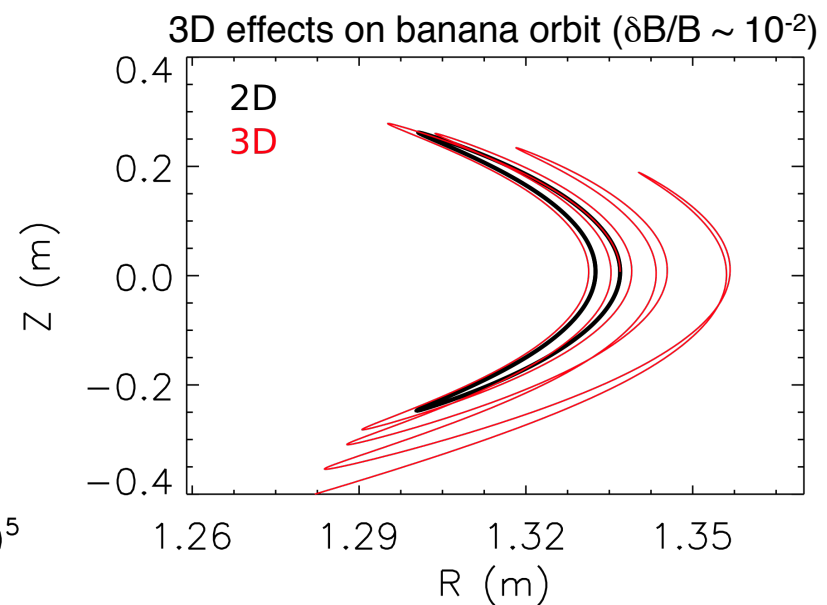
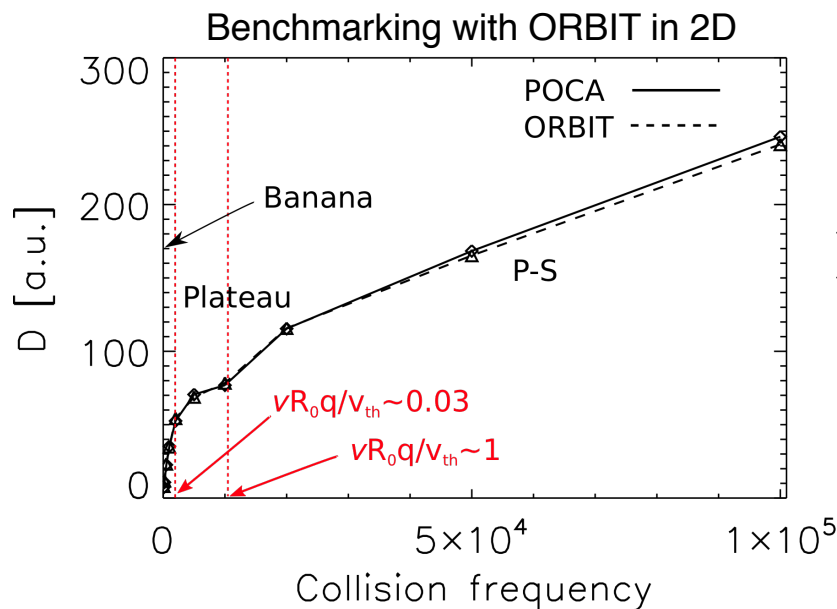
- δf guiding-center particle code – POCA
- Applications to validating theories
- Applications to magnetic braking experiments
- Applications to field line tracing
- Summary

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POCA is a drift-kinetic δf particle code for neoclassical transport calculation in NAMPs

- POCA (Particle Orbit Code for Anisotropic pressures) [K. Kim, POP (2012)]
 - Follows guiding-center orbit motions on the flux coordinates
 - Solves Fokker-Plank equation with modified pitch-angle scattering collision operator conserving toroidal momentum
 - Calculates local neoclassical quantities: Diffusion, flux, bootstrap current
 - Calculates anisotropic tensor pressure and NTV torque
 - Uses DCON/IPEC type routines and parallelized with MPI
 - Reads 2D equilibrium from 20 equilibrium types and 3D magnetic perturbations from IPEC and analytic model



POCA solves Fokker-Planck equation with guiding-center orbit equations

- Guiding-center motion is described by Hamiltonian equations of motion

$$\dot{\theta} = -\frac{2\pi}{q} \frac{1}{g+u} \left\{ \frac{q^2 B^2}{m} \rho_{\parallel} \left[\rho_{\parallel} \frac{\partial g(\psi)}{\partial \psi} - u(\psi) \right] - g(\psi) \left[\left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \psi} + q \frac{\partial \Phi}{\partial \psi} \right] \right\}$$

[R.B. White, PFL B (1990)]
[M. Sasinowski and A.H. Boozer, POP (1997)]

$$\dot{\psi} = -\frac{2\pi}{q} \frac{1}{g+u} \left\{ i(\psi) \left[\left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \psi} + q \frac{\partial \Phi}{\partial \psi} \right] - \left(\rho_{\parallel} \frac{\partial i(\psi)}{\partial \psi} + 1 \right) \frac{q^2 B^2}{m} \rho_{\parallel} \right\}$$

$$\dot{\psi} = -\frac{2\pi}{q} \frac{1}{g+u} \left[g(\psi) \left\{ \left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \theta} + q \frac{\partial \Phi}{\partial \theta} \right\} - i(\psi) \left\{ \left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \phi} + q \frac{\partial \Phi}{\partial \phi} \right\} - \frac{q}{2\pi} V(t) i(\psi) \right]$$

$$\dot{\rho}_{\parallel} = \frac{2\pi}{q} \frac{1}{g+u} \left[\left(\rho_{\parallel} \frac{\partial g(\psi)}{\partial \psi} - u(\psi) \right) \left\{ \left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \theta} + q \frac{\partial \Phi}{\partial \theta} \right\} - \left(\rho_{\parallel} \frac{\partial i(\psi)}{\partial \psi} + 1 \right) \left\{ \left(\frac{q^2 \rho_{\parallel}^2 B^2}{m} + \mu B \right) \frac{1}{B} \frac{\partial B}{\partial \phi} + q \frac{\partial \Phi}{\partial \phi} \right\} - \frac{q}{2\pi} V(t) \left(\rho_{\parallel} \frac{\partial i(\psi)}{\partial \psi} + 1 \right) \right]$$

- δf is calculated from Fokker-Planck equation

[M.N Rosenbluth, PFL (1972)]
[M. Sasinowski and A.H. Boozer, POP (1997)]

- Fokker-Planck equation is written as

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \vec{v} \cdot \frac{\partial f}{\partial \vec{x}} + \frac{\vec{F}}{m} \cdot \frac{\partial f}{\partial \vec{v}} = C(f), \quad f = f_M \exp(\hat{f}) \approx f_M (1 + \hat{f}) \longrightarrow \frac{d \ln f_M}{dt} + \frac{d \hat{f}}{dt} = C_m(f) \equiv \frac{C(f)}{f}$$

- Fokker-Planck equation is reduced to

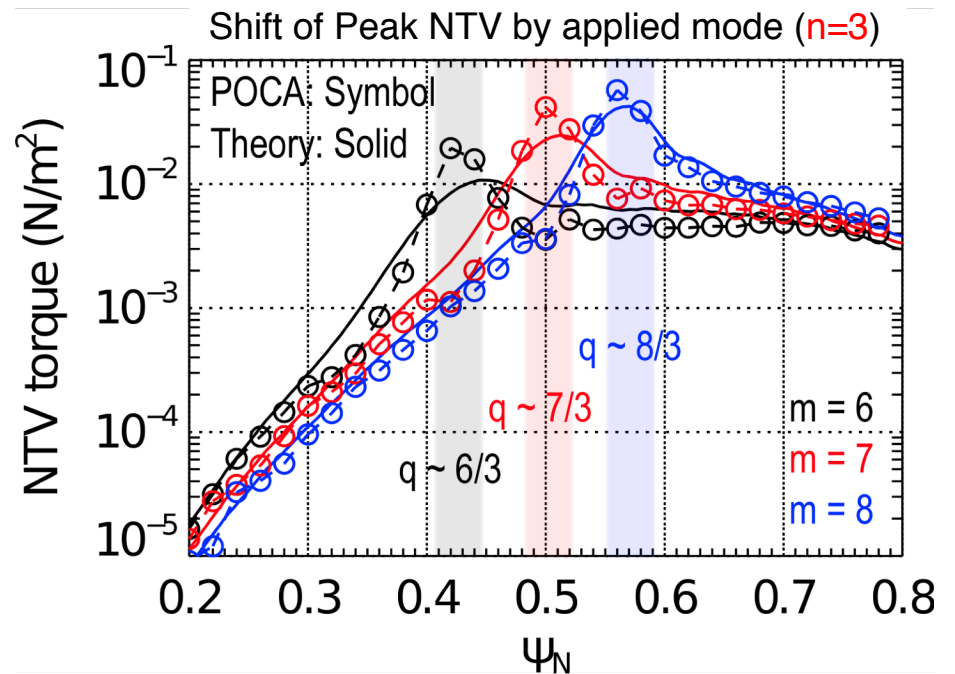
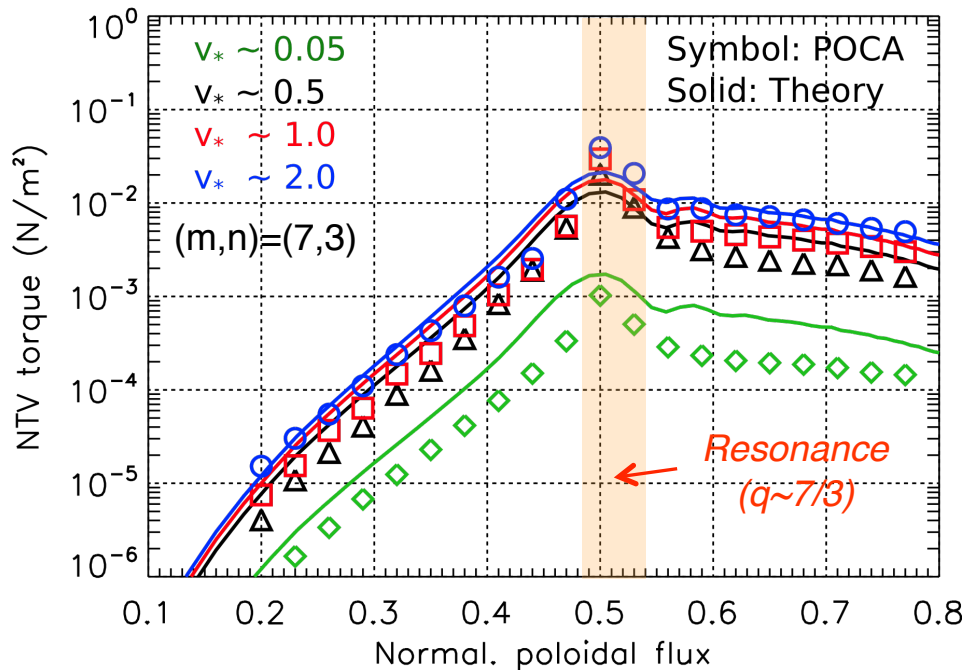
$$\frac{d \hat{f}}{dt} - C_m(f) = -\vec{v} \cdot \nabla \psi \frac{\partial \ln f_M}{\partial \psi} - \vec{F} \cdot \frac{\partial \ln f_M}{\partial \vec{v}}$$

- Using local Maxwellian, δf can be obtained from

$$f_M = \frac{N}{(\sqrt{\pi} v_t)^3} \exp\left(-\frac{U - e\Phi}{T}\right) \longrightarrow \Delta \hat{f} = -\left[\frac{1}{n} \frac{\partial n}{\partial \psi} + \left(\frac{3}{2} - \frac{E}{T} \right) \frac{1}{T} \frac{\partial T}{\partial \psi} \right] \Delta \psi - \frac{e}{T} \frac{d\Phi}{d\psi} \Delta \psi + 2v \frac{u}{v} \lambda \Delta t$$

Neoclassical Toroidal Viscosity (NTV) is calculated using perturbed pressures and magnetic field spectrum

- NTV calculation was benchmarked with analytic NTV theory
 - POCA calculates NTV torque using magnetic field spectrum decomposed to Fourier series and perturbed pressures
 - [J.D. Williams and A.H. Boozer, POP (2003)]
 - NTV torque is calculated in Boozer coordinates by $\tau_\phi = \langle \hat{e}_\phi \cdot \nabla \cdot \vec{P} \rangle = \left\langle \frac{\delta P}{B} \frac{\partial B}{\partial \phi} \right\rangle$
 - Calculated NTV torque shows good agreements with theory, revealing strong resonant features to applied magnetic perturbation [K. Kim, POP (2012)]



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Combined NTV theory is a good guide for benchmarking

- Combined NTV formula reflects fundamental NTV physics [J.-K. Park, PRL (2009)]
 - Bounce-averaged drift-kinetic equation based on BH resonance is analytically solved using Krook collision to connect regimes
 - The combined NTV formula gives the toroidal torque density as

$$\langle \hat{e}_\phi \cdot \nabla \cdot \vec{\Pi} \rangle_l = \frac{\varepsilon^{-1/2} P}{\sqrt{2\pi} R_0^{3/2}} \int_0^1 d\kappa^2 \delta_{\omega,l}^2 \int_0^\infty dx R_{yl} \left[u^\varphi + 2.0\sigma \left| \frac{1}{e} \frac{dT}{d\chi} \right| \right]$$

δB^2 Resonance $\omega_{rot} - \omega_{neo}$

$$\delta_{\omega,l}^2 = \sum_{mmm'} \delta_{mmm'}^2 \frac{F_{nml}^{-1/2} F_{nm'l}^{-1/2}}{4K(\kappa)}$$

$$F_{nml}^y \equiv \int_{-\vartheta_l}^{\vartheta_l} d\vartheta \left(\kappa^2 - \sin^2(\vartheta/2) \right)^y \cos(m - nq \pm l)\vartheta$$

Phase shift of resonant surface by ExB precession

$$R_{yl} = \frac{1}{2} \frac{n^2 \left(1 + (l/2)^2 \right) \frac{v}{2\varepsilon} x \left(x - \frac{5}{2} \right)^y e^{-x}}{\left(l\omega_{l,bounce} - n\omega_{E \times B} - n\omega_{\nabla B} \right)^2 + \left(\left(1 + (l/2)^2 \right) \frac{v}{2\varepsilon} \right)^2 x^{-3}}$$

Bounce-harmonic resonance

Magnetic precession

- Include δB^2 dependency, various resonances between bounce motion, magnetic precession, electric precession, and neoclassical offset rotation

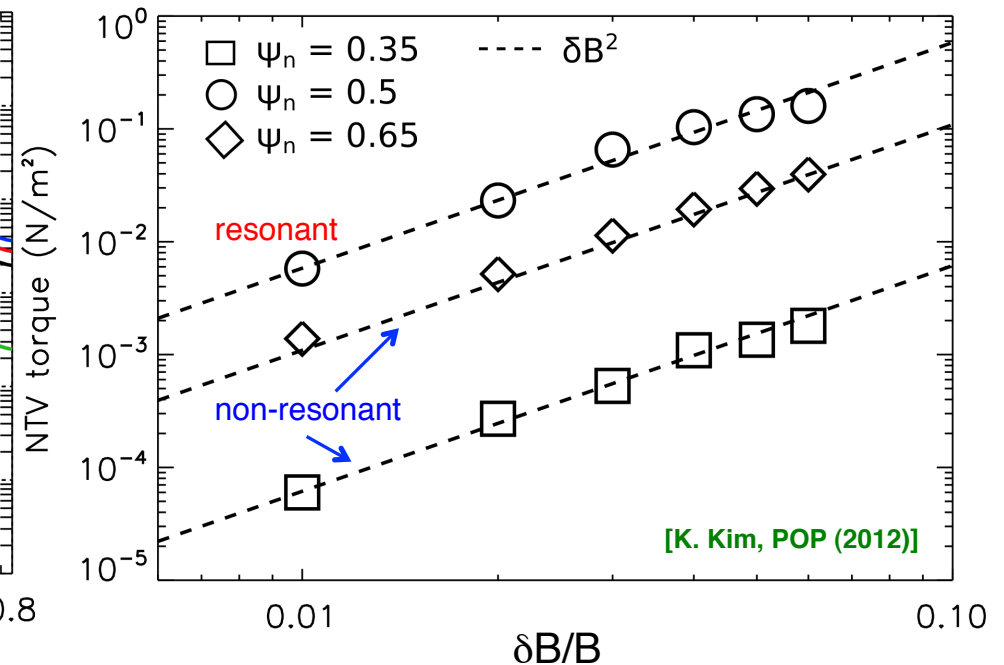
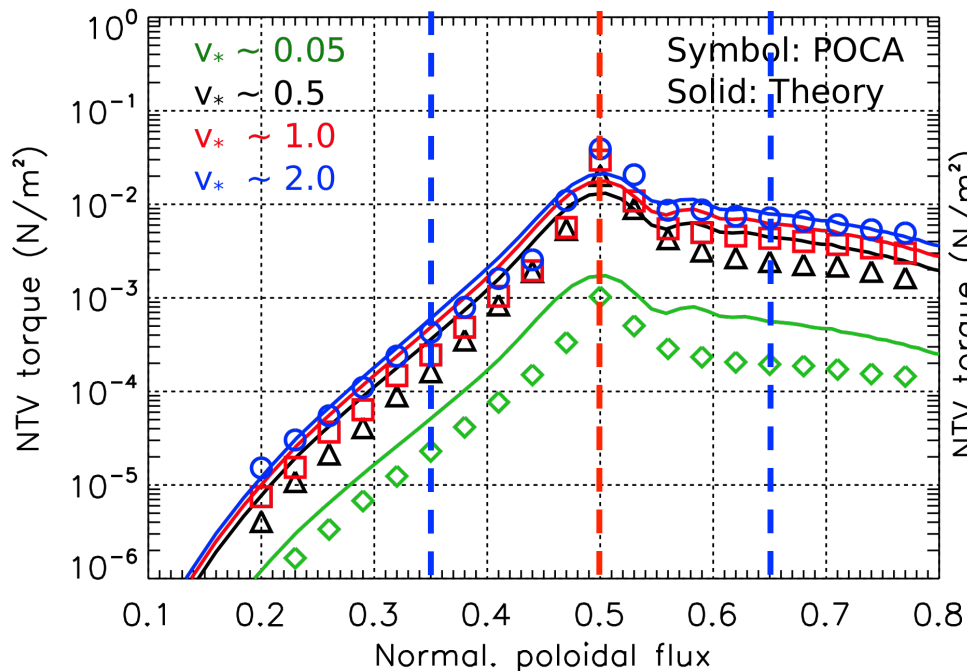
Quadratic δB dependency is the fundamental NTV physics

- Particle simulation confirms the quadratic δB dependency of NTV

$$\langle \hat{e}_\phi \cdot \nabla \cdot \vec{\Pi} \rangle_l = \frac{\varepsilon^{-1/2} P}{\sqrt{2\pi}^{3/2} R_0} \int_0^1 d\kappa^2 \delta_{\omega,l}^2 \int_0^\infty dx R_{1l} \left[u^\varphi + 2.0\sigma \left| \frac{1}{e} \frac{dT}{d\chi} \right| \right]$$

δB^2

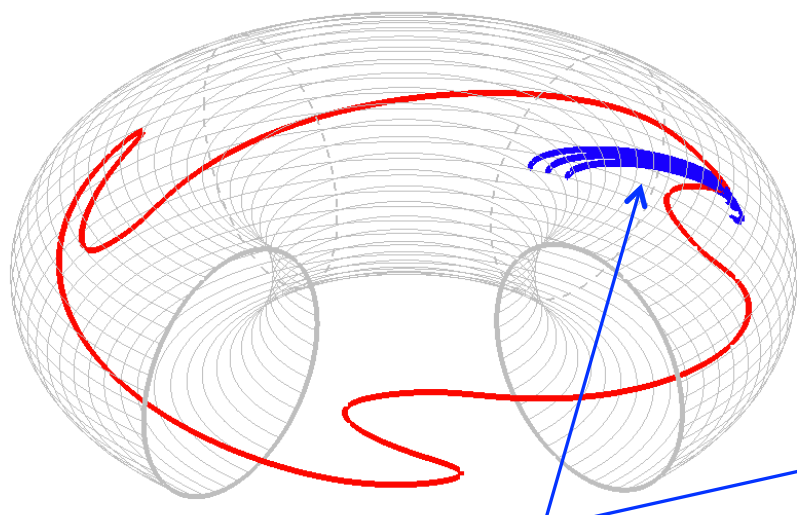
- NTVs at both resonant and non-resonant flux surfaces increase as δB increases following $\tau_\phi \sim (\delta B)^2$
- Useful to particle simulation, enable NTV scaling by δB to improve statistics



New type of closed orbit is formed by rotational resonance

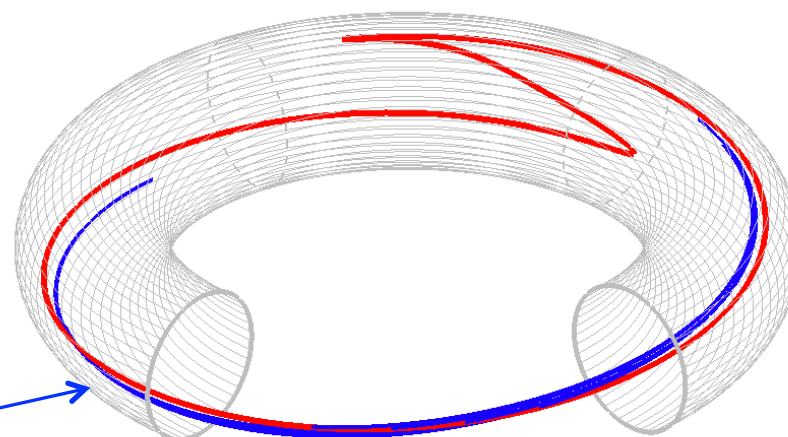
- Parallel/perpendicular ExB drifts can form a closed loop at resonance
 - Resonant field is shifted to follow path of the closed orbit satisfying
Resonant field condition: $m - nq \pm l = 0$
 - Resonance condition is represented by particle's bouncing class ℓ
Resonant condition: $l\omega_b \approx n\omega_E$
 - Banana orbits can be closed, if ExB is fast enough to reach BH resonant frequency
 - Various closed orbits can exist depending on energy and pitch of particle, magnetic field configuration, and ExB precession

Modified closed orbit ($n=3, \ell=1$)



Original bounce orbit w/o rotation

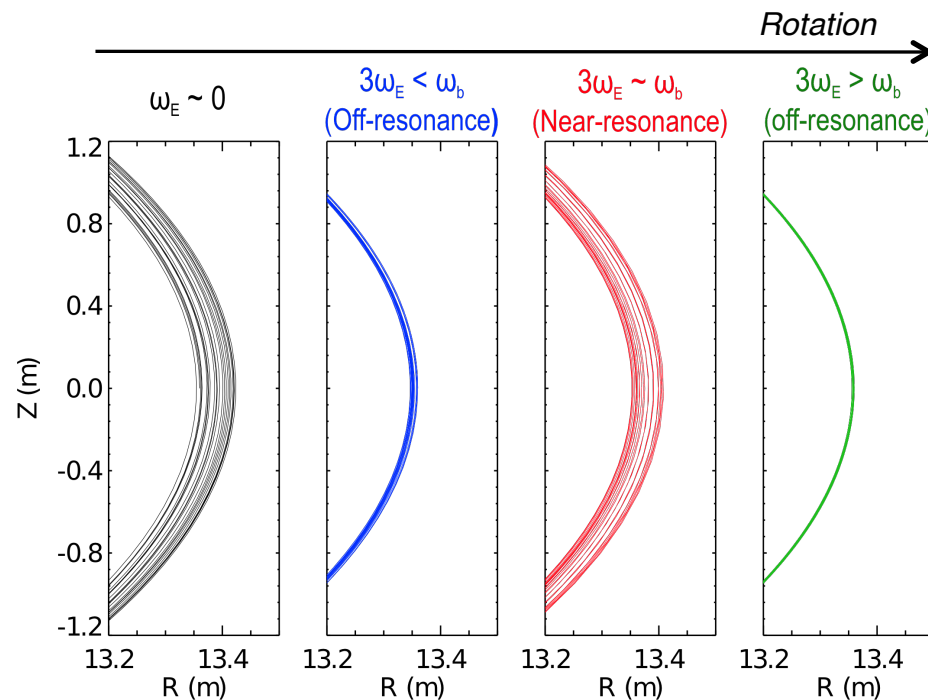
Modified closed orbit ($n=1, \ell=1$)



[K. Kim, PRL (2013)]

Rotation scan clarifies transport mechanism by rotational resonances

- BH resonances enhance radial particle transport by closed orbits preventing random phase-mixing
 - Radial drift is large when $\omega_E \sim 0$ as orbits are nearly closed without ExB precession, equivalent to superbanana plateau resonance
 - When ω_E is off-resonant ($3\omega_E \neq \omega_b$), particle drift decreases due to phase-mixing
 - When ExB approaches the first BH resonant frequency ($3\omega_E \sim \omega_b$), modified closed orbit prevents phase-mixing, thus enhances radial transport and NTV



[K. Kim, PRL (2013)]

Peak NTV is found at the resonant ExB precession frequency, verifying BH resonances in NTV

- BH resonance enhanced NTV is confirmed by Er scan

- Theory indicates bounce-harmonic resonances occur when

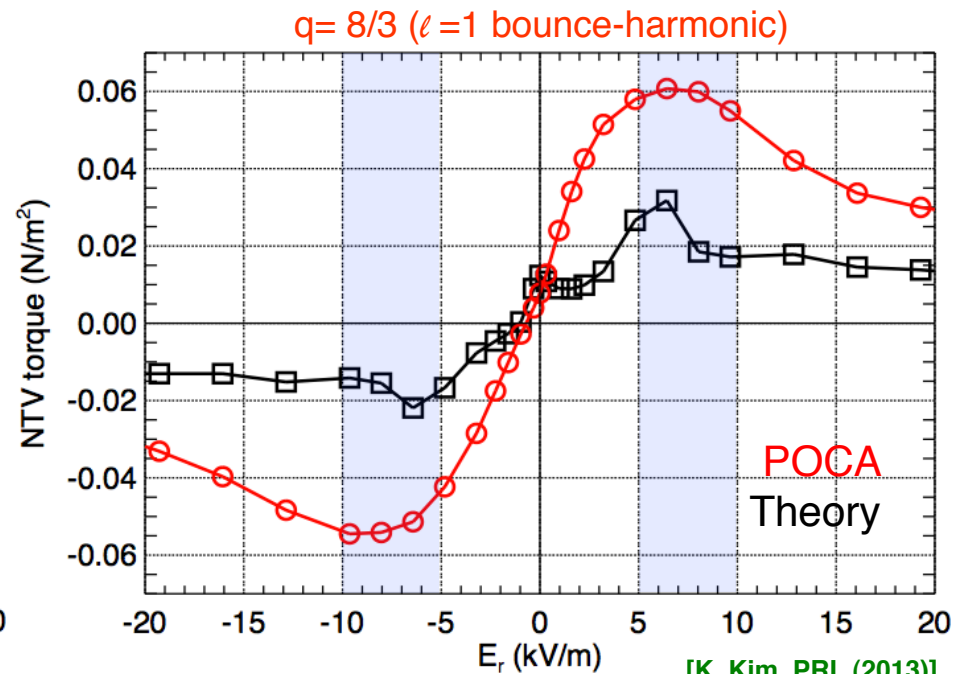
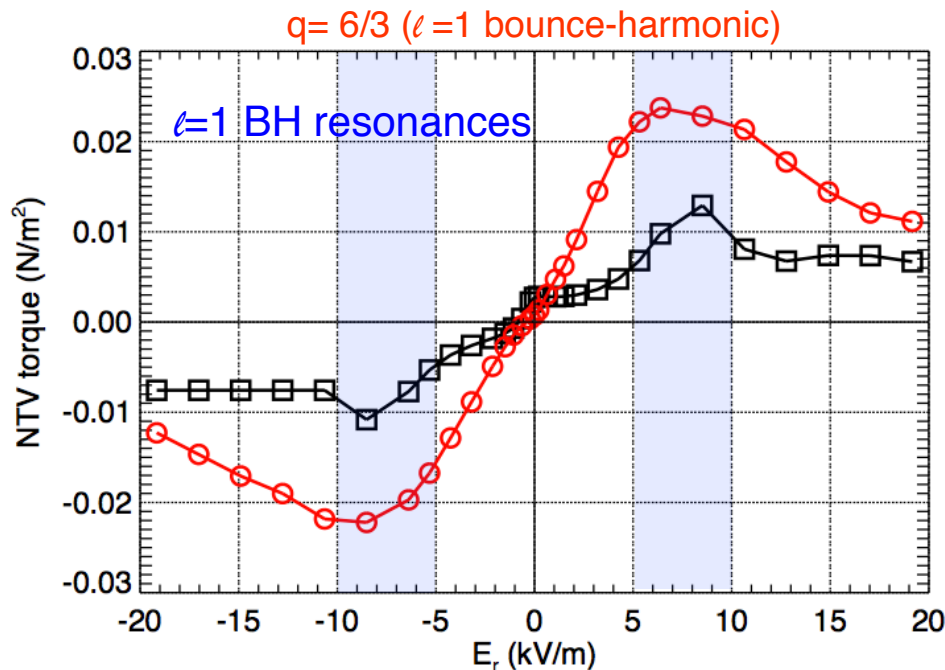
$$l\omega_b - n\omega_E - n\omega_{\nabla B} = 0 \quad \rightarrow \quad l\omega_b \approx n\omega_E$$

- Resonance flux surface is shifted to $m - nq \pm l = 0 \quad \rightarrow \quad q = (m \pm l)/n$

- Bounce-harmonic resonance by $\ell = 1$ class particles with $(m=7, n=3)$ is predicted at

$$\omega_E \approx E_r / RB_p \quad \rightarrow \quad E_r \approx \pm 7 \text{ kV/m}$$

- POCA reproduces strong peak NTV around predicted ExB precession frequency

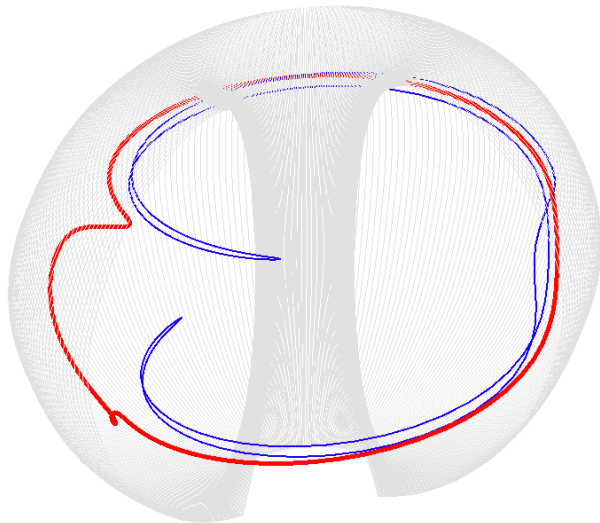


[K. Kim, PRL (2013)]

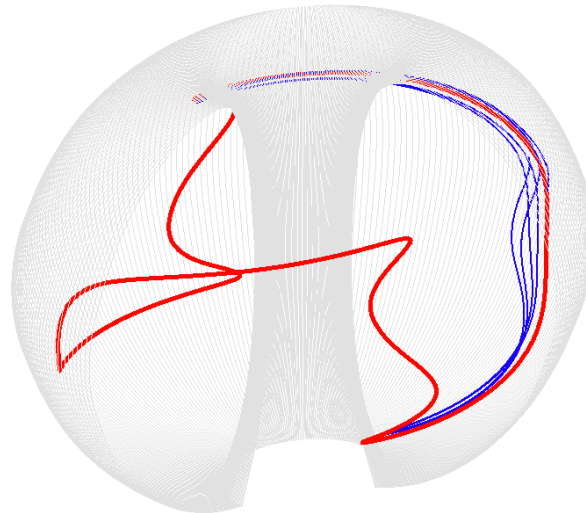
Closed orbits by BH resonances can be found in NSTX

- Bounce-harmonic resonance almost always exist in perturbed tokamaks
 - BH resonance always exist in the finite ExB due to Maxwellian energy distribution, and on every surface due to multi-harmonic magnetic perturbations
 - Modified closed orbits, theoretically predicted and numerically reproduced in the simple configuration, can be also found in the complicated NSTX configuration
 - Identical features in orbit-closing by resonance

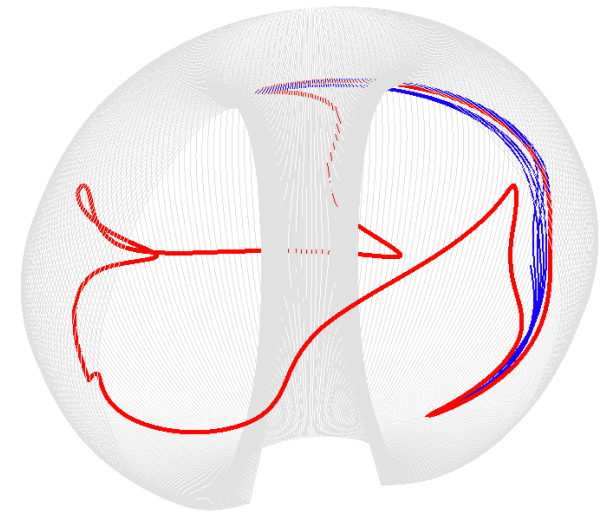
Closed orbit ($n=1, \ell=1$)



Closed orbit ($n=2, \ell=1$)



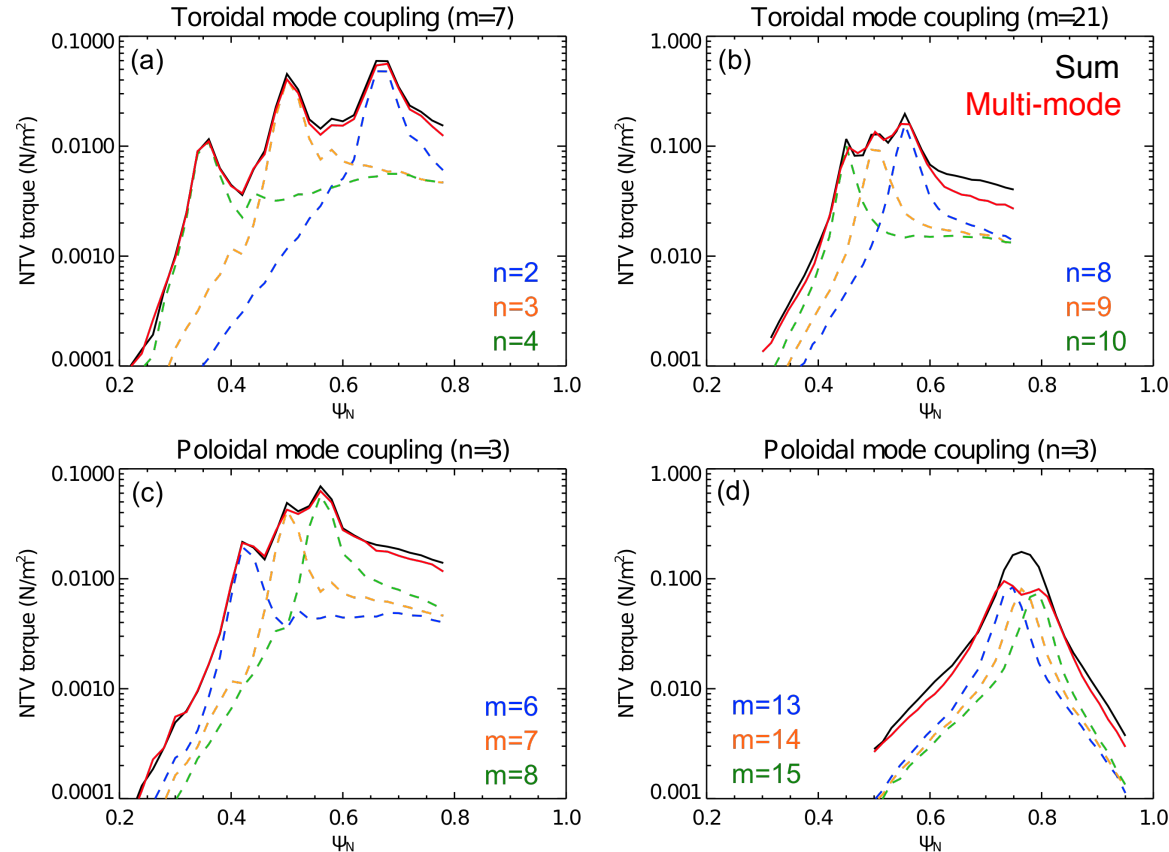
Closed orbit ($n=3, \ell=1$)



————— Original bounce orbits w/o rotation

Mode coupling effect of NTV was tested using analytic δB

- Poloidal modes can be coupled but toroidal mode cannot
 - Three-mode magnetic perturbation applied and compared with linear NTV sum of three single-mode perturbations of the same harmonics
 - Indicate no or weak toroidal mode coupling effect regardless gap
 - Poloidal modes can couple when the gap is small



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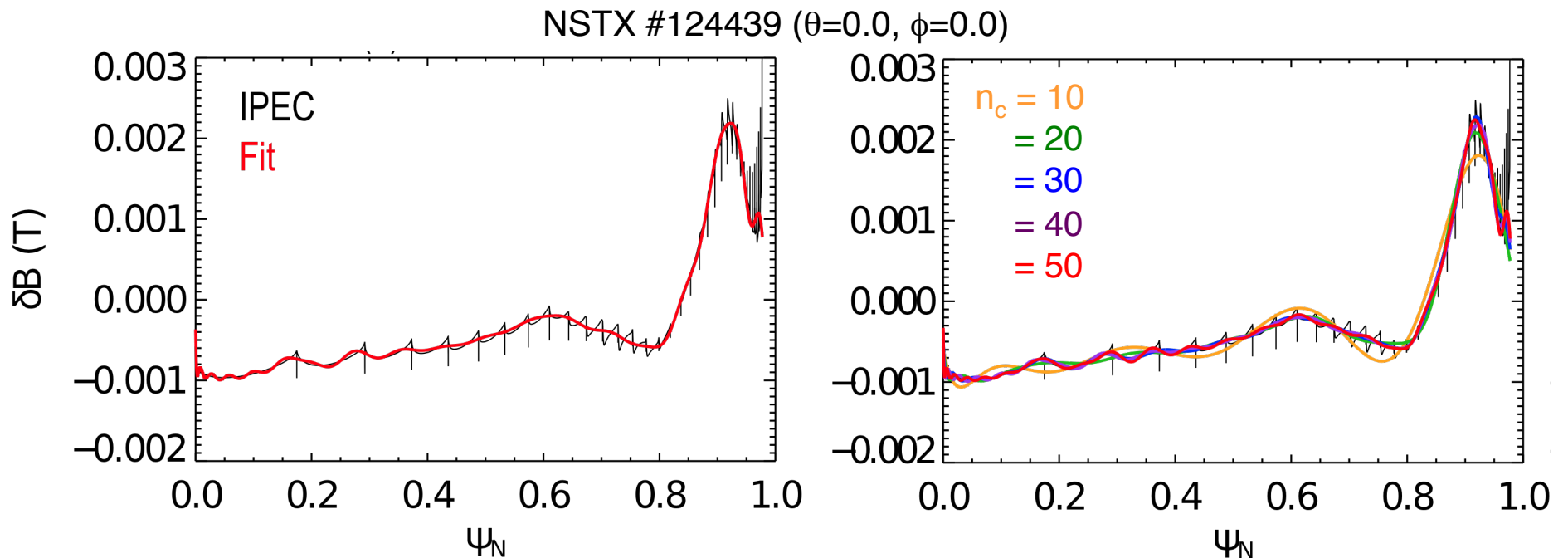
POCA is being actively applied to experimental analysis

- Perturbed magnetic field spectrum provided by IPEC
 - Original IPEC output contains nonphysical peaks at the rational surfaces
 - Fitting technique (i.e. Chebyshev polynomials) will be used in POCA as

$$\delta B_{mn}(\psi_n) = \sum_m a_{mn}(\psi_n) \cos(m\theta - n\phi) + b_{mn}(\psi_n) \sin(m\theta - n\phi) \quad \leftarrow \text{IPEC}$$

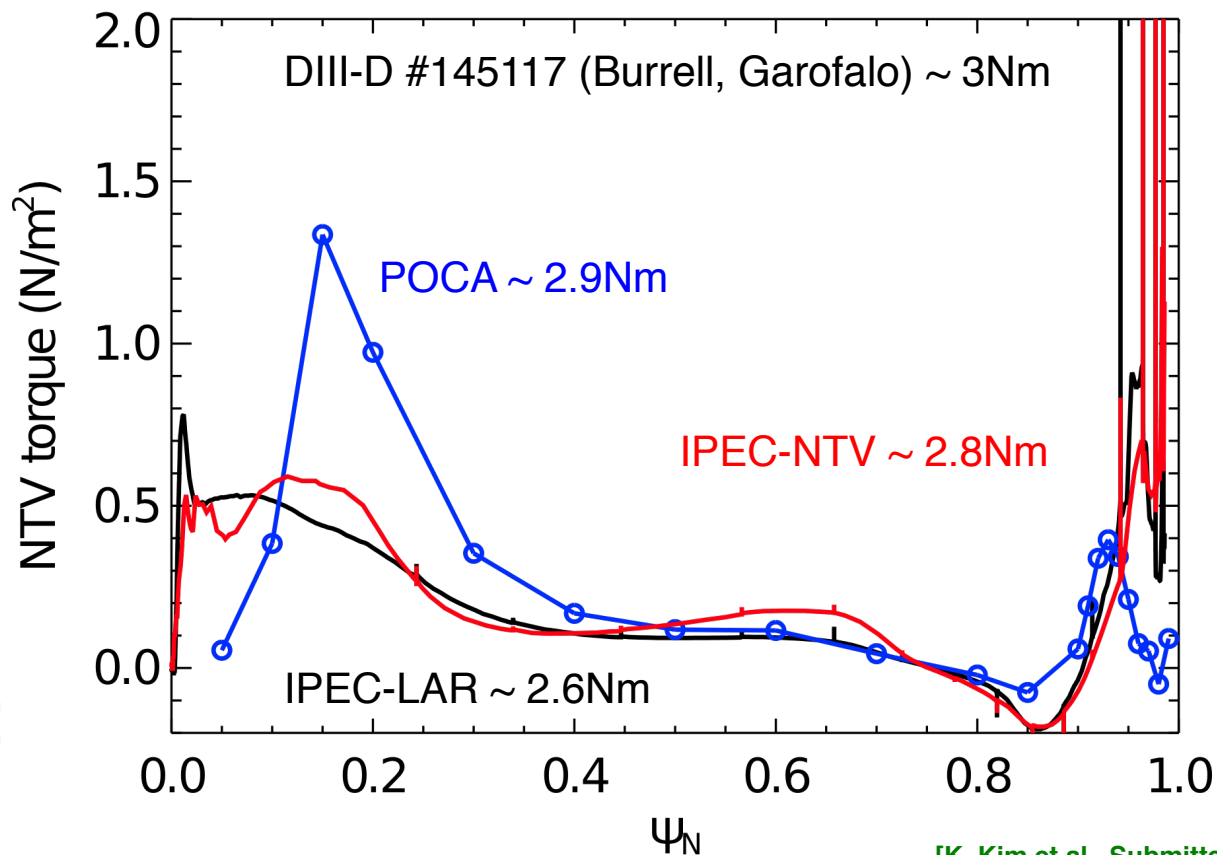
$$\delta B_{mn}(\psi_n) = \sum_m \left[\sum_j^{n_c} a_j \cos(j \cos^{-1}(x)) \cos(m\theta - n\phi) + b_j \cos(j \cos^{-1}(x)) \sin(m\theta - n\phi) \right] \quad \rightarrow \text{POCA}$$

- Fitting follows overall features of IPEC δB , and effectively smoothes the peaks



Consistent NTV profile and total NTV was obtained in DIII-D n=3 magnetic braking

- Application of POCA to DIII-D n=3 magnetic braking (in QH mode experiments) using IPEC δB gives a consistent NTV
 - Consistent with combine NTV formula in profile, and good agreement with measurement in total NTV torque

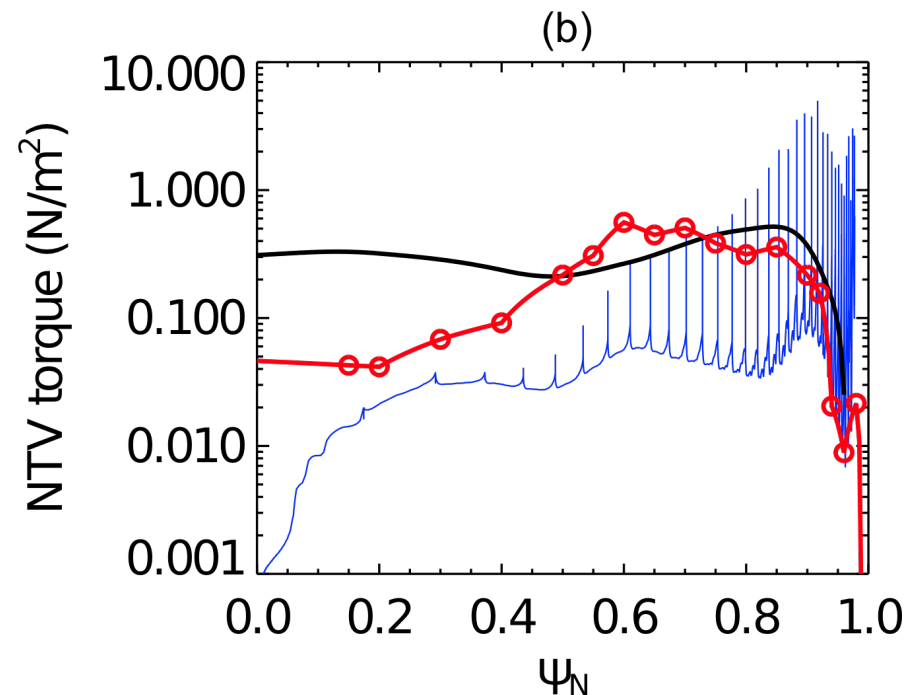
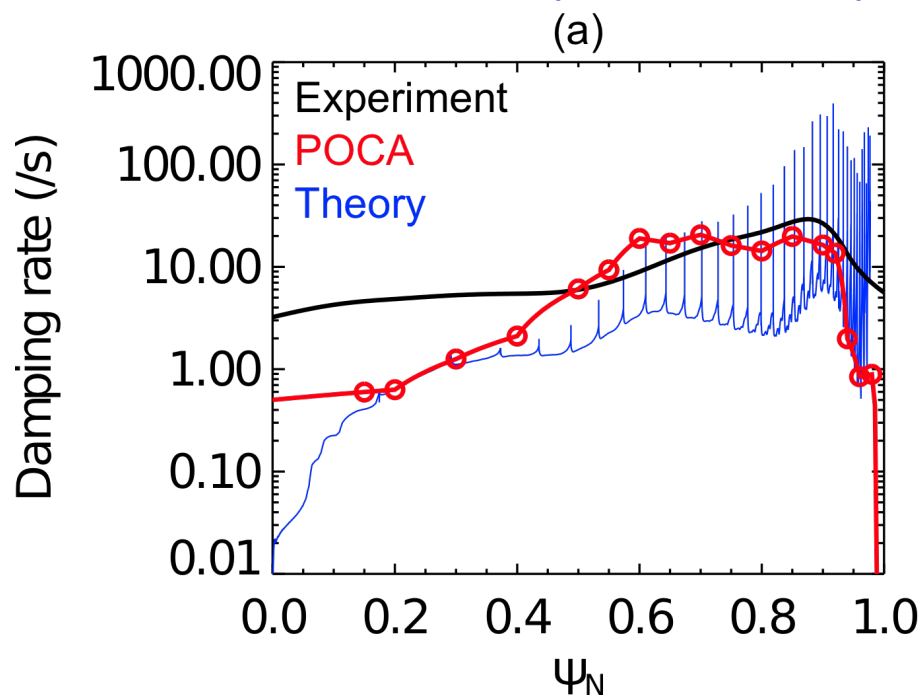


[K. Kim et al., Submitted to Nucl. Fusion (2012)]



Prediction of NTV in NSTX can be improved with POCA

- Improved agreement with measurement in NSTX n=3 discharge
 - POCA calculates damping rate from NTV / Experiment gives NTV from damping rate
 - Both damping and NTV torque profiles show improved agreements with measurements
 - Combined theory is valid only within an order of magnitude, due to the large aspect-ratio expansion and ignoring finite-orbit-width
 - Total NTV torque agrees well: Exp. $\sim 3.5\text{Nm}$, POCA $\sim 2.7\text{Nm}$
 - Room for further improvement in experimental estimation and numerical simulation



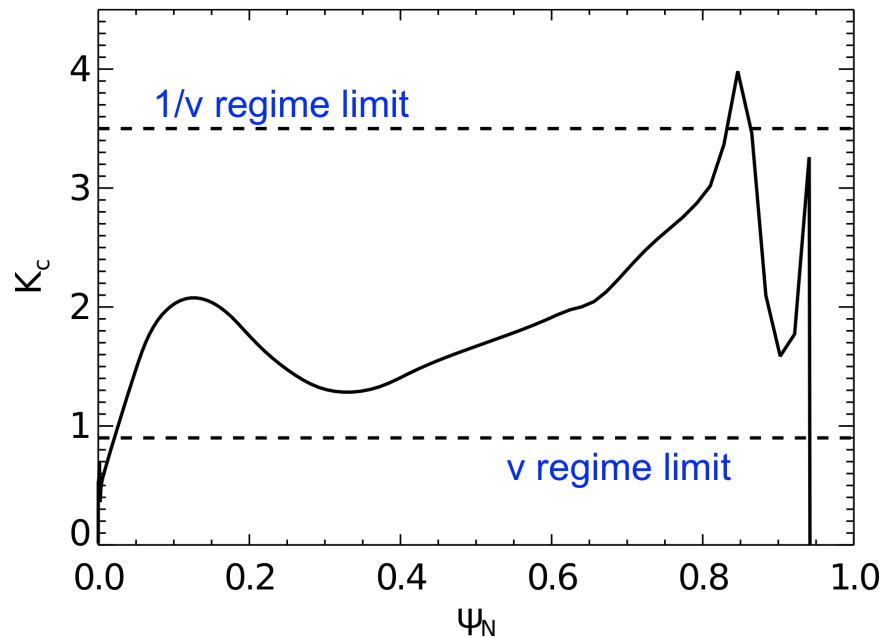
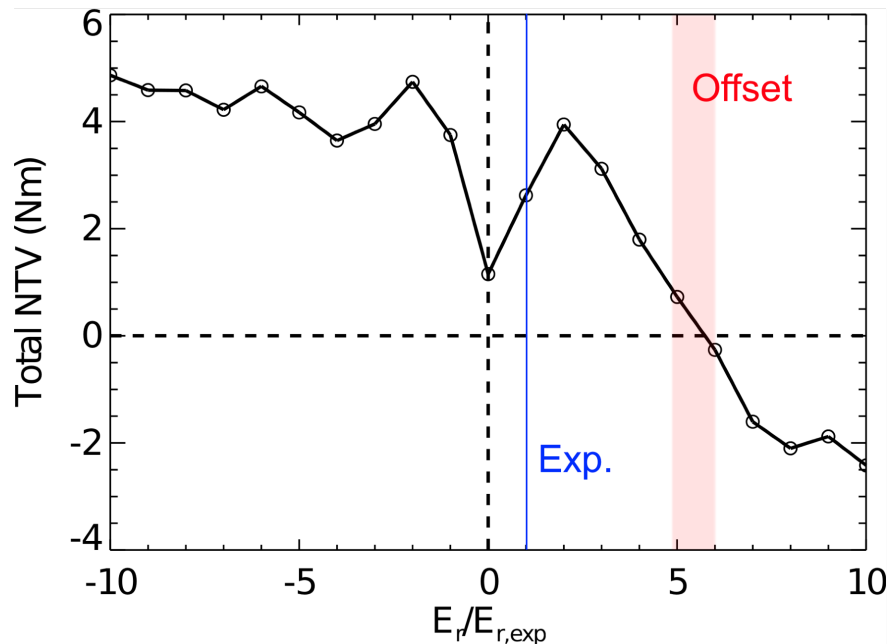
[K. Kim et al., Submitted to Nucl. Fusion (2012)]

Neoclassical offset can be identified at zero-torque

- Rotation scan predicts neoclassical offset rotation
 - Neoclassical offset can be found approximately, when NTV crosses zero in rotation scan ($E_r/E_{r,exp} \sim 5$)
 - Theoretically, neoclassical offset rotation is predicted by

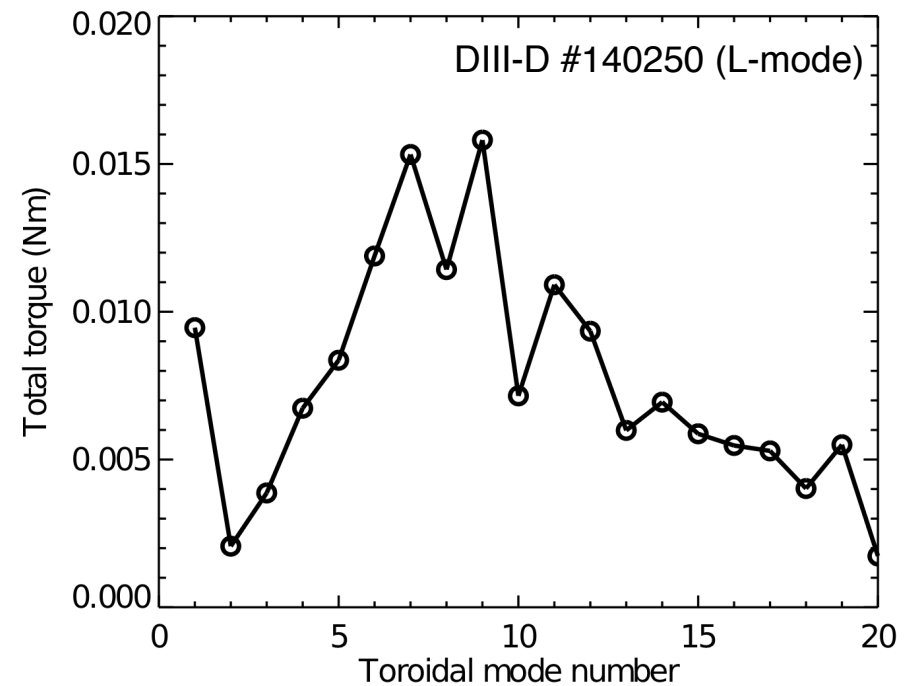
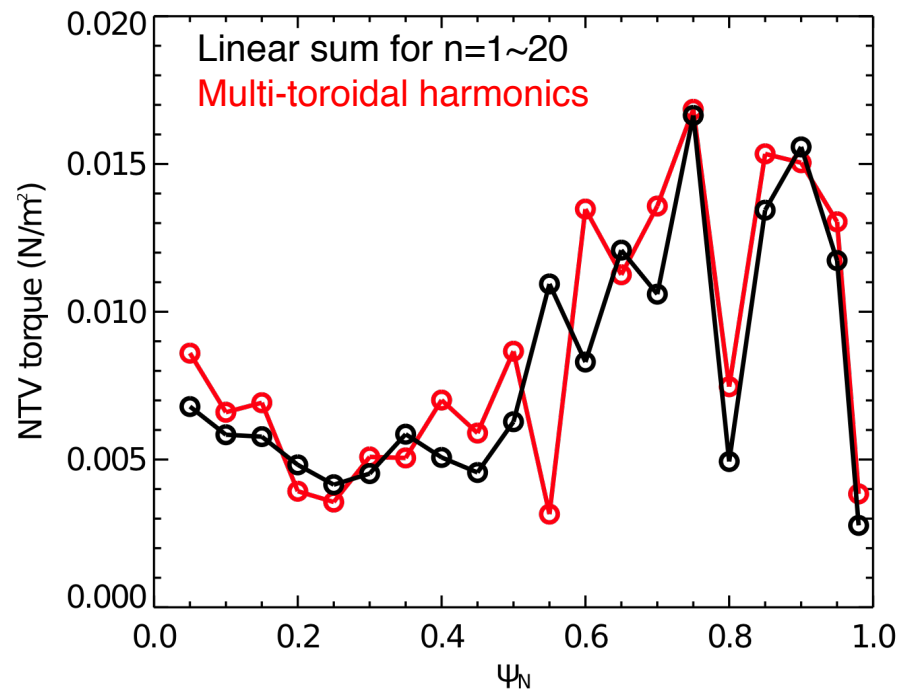
$$V_{\phi,offset} = K_c \frac{1}{Z_i e B_\theta} \frac{dT_i}{dr} \quad K_c = \frac{V_{\phi,offset}}{\frac{1}{Z_i e B_\theta} \frac{dT_i}{dr}} \approx 3.5 \text{ (1/\nu regime) or } 0.9 \text{ (\nu regime)}$$

- Numerically obtained K_c is consistent with theory prediction, indicating the plasma is in the connected regime



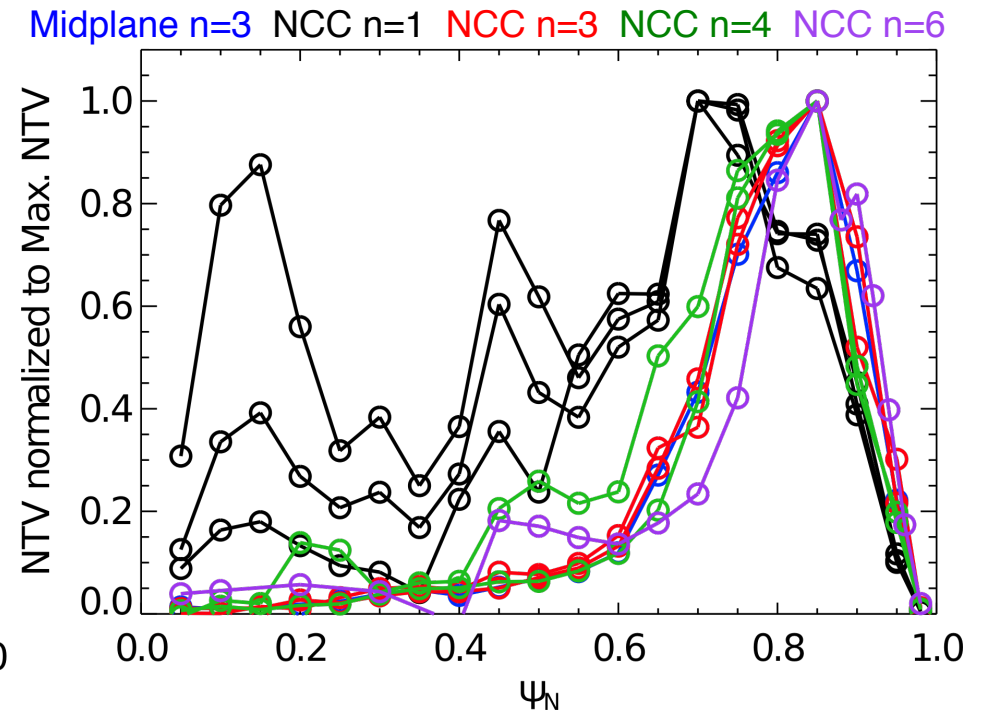
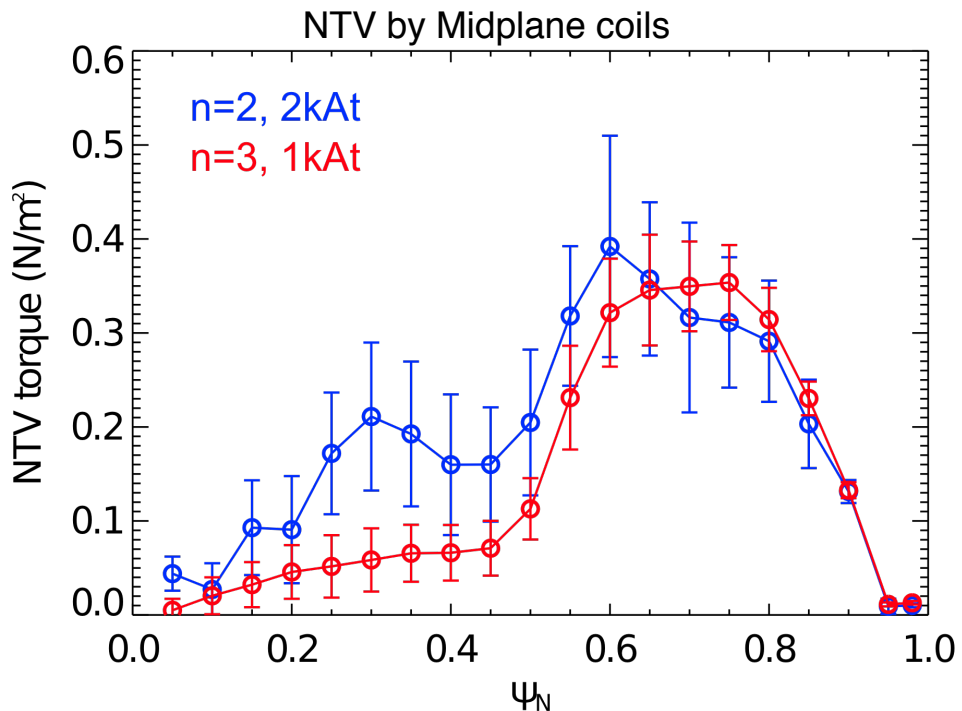
Toroidal mode coupling was tested for TBM mock-up experiment in DIII-D

- No (or weak) toroidal coupling was found in TBM case
 - POCA can take multi-toroidal harmonics simultaneously
 - Obtained NTV by multi-modes comparable to linear NTV sum by single modes (~ 0.2 Nm with $10\times\delta B$ in L-mode), indicating no (weak) toroidal mode coupling
 - Consistent trend with theory for NTV vs. toroidal mode number, but smaller NTV
→ Weak (no) braking in experiment?
 - High β plasma with visible braking effects will be the next target



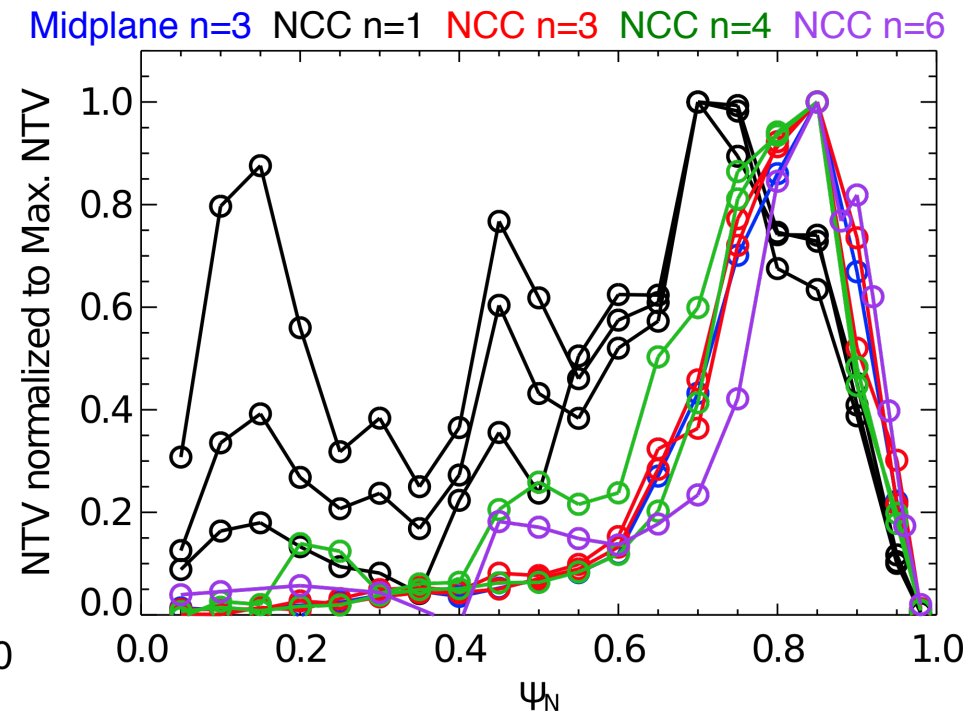
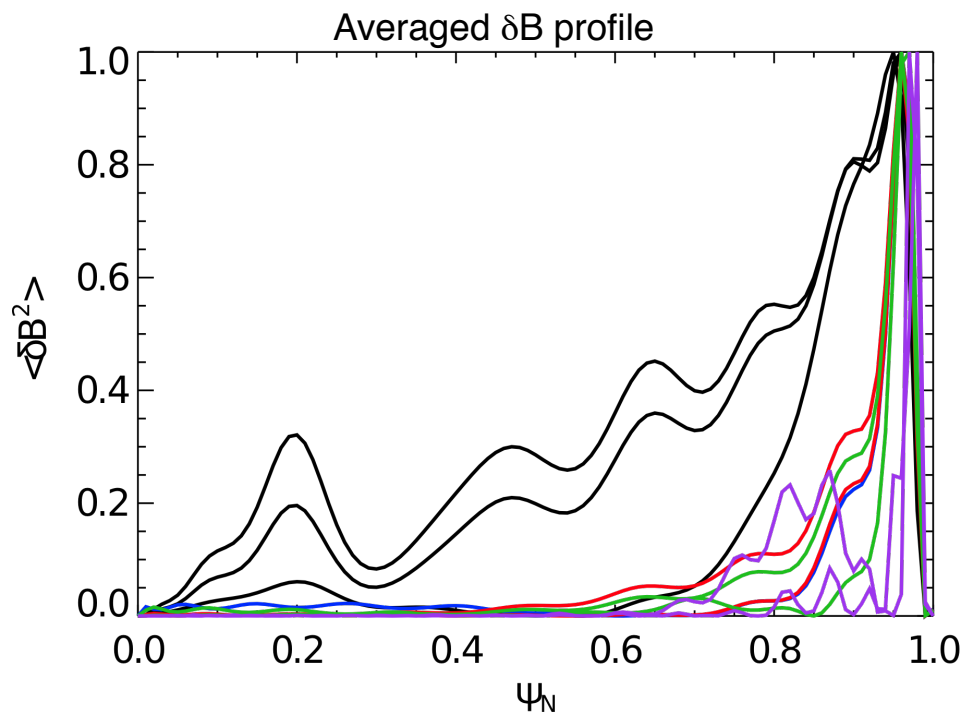
Various Braking Profiles in NSTX and NSTX-U

- POCA predicts variability of braking profiles by Midplane and NCC
 - Midplane coils can drive broad ($n=2$) or peaked ($n=3$) NTV profiles
 - (Full and/or partial) NCC can provide various braking profiles by $n=1 \sim 6$, depending on phases



Various Braking Profiles in NSTX and NSTX-U

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 - Midplane coils can drive broad ($n=2$) or peaked ($n=3$) NTV profiles
 - (Full and/or partial) NCC can provide various braking profiles by $n=1 \sim 6$, depending on phases: **Consistent to δB profiles**
 - Polynomial degree for fitting δB should be carefully selected in high- n ($=4,6$) cases for better radial resolution in NTV calculations



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New Capability for Field Line Tracing

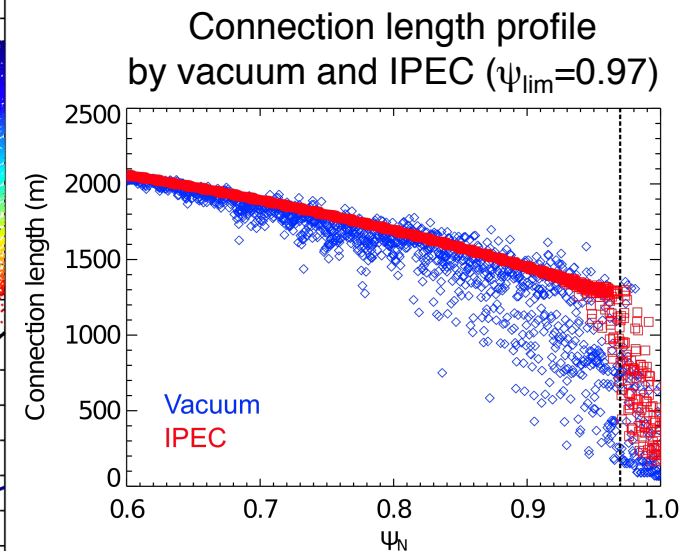
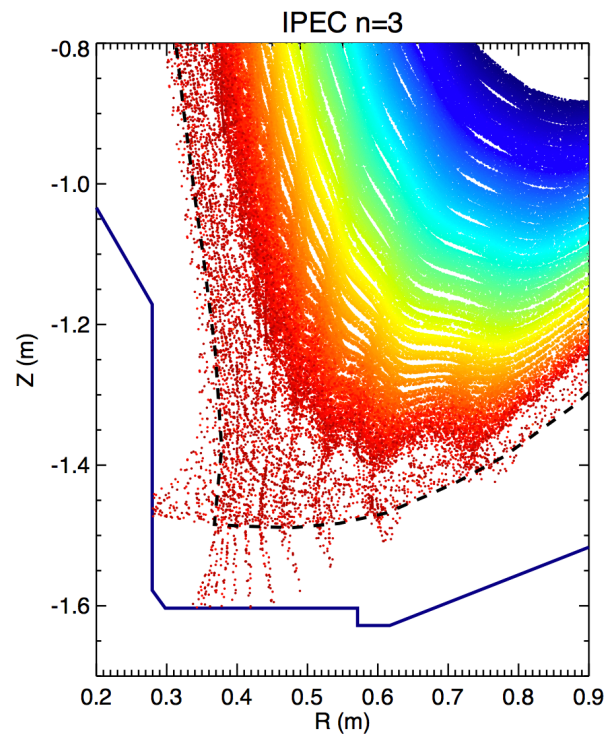
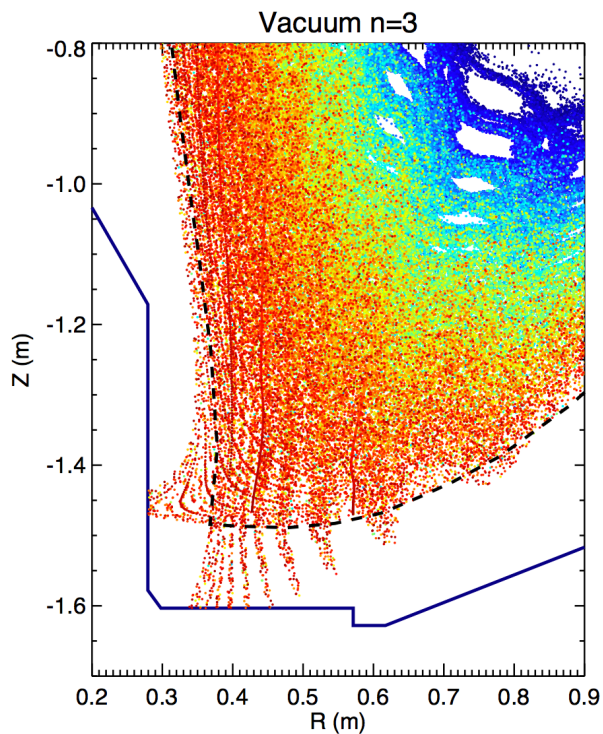
- Field line tracing (FLT) routine is now implemented in POCA
 - Use cylindrical coordinates (R, Z, ϕ) including divertor/limiter/wall structures
 - Solves a set of magnetic differential equations

$$\frac{\partial R}{\partial \phi} = \frac{RB_R}{B_\phi} \qquad \frac{\partial Z}{\partial \phi} = \frac{RB_Z}{B_\phi}$$

- Provide Poincare plot and divertor footprint modified by 3D field
 - Calculate field line connection length and field line loss fraction
 - Benchmarked with NSTX and DIII-D cases (i.e. lobe, divertor footprint, heat flux)
 - Applicable to NSTX-U NCC coil design study
- POCA-FLT has a unique capability via coupling with IPEC
 - Vacuum δB or IPEC δB with ideal plasma response can be taken, which reveals the shielding of resonant field and δB amplification by ideal response
 - Every possible field component can be included for NSTX(-U) with or without plasma response (EFC, PF5EF, TFEF)
 - Multi-toroidal mode can be calculated simultaneously
 - Parallelized for computational efficiency

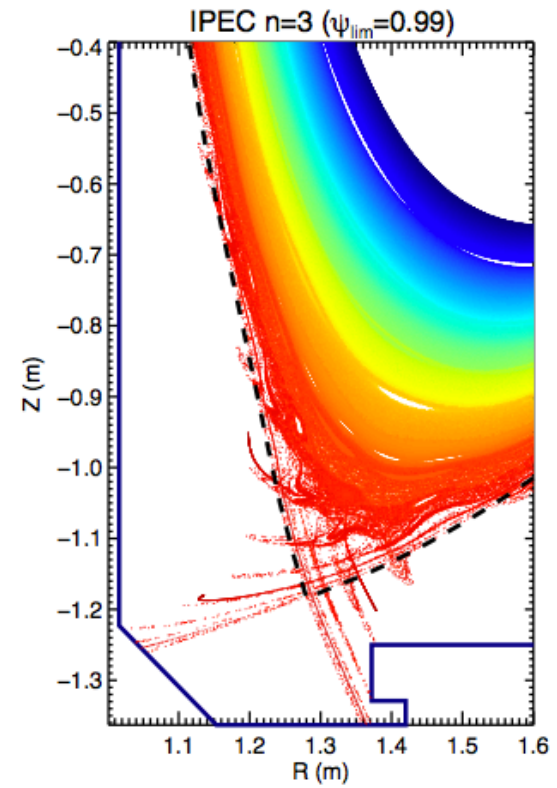
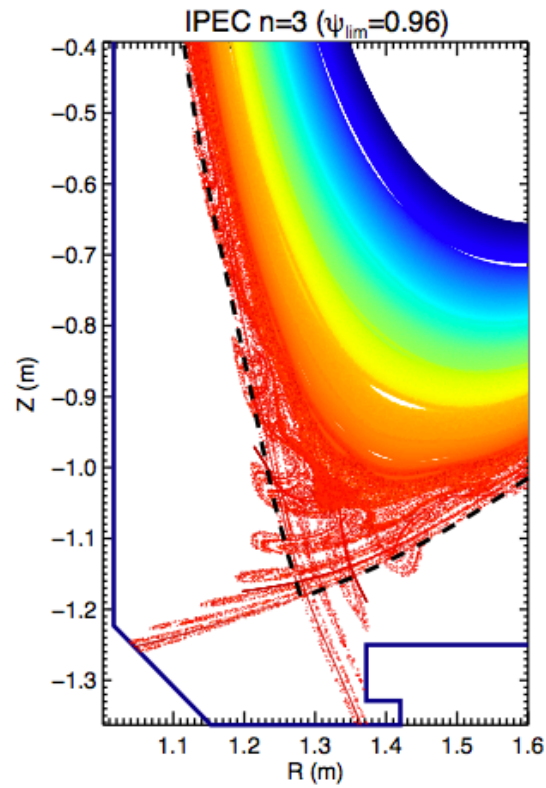
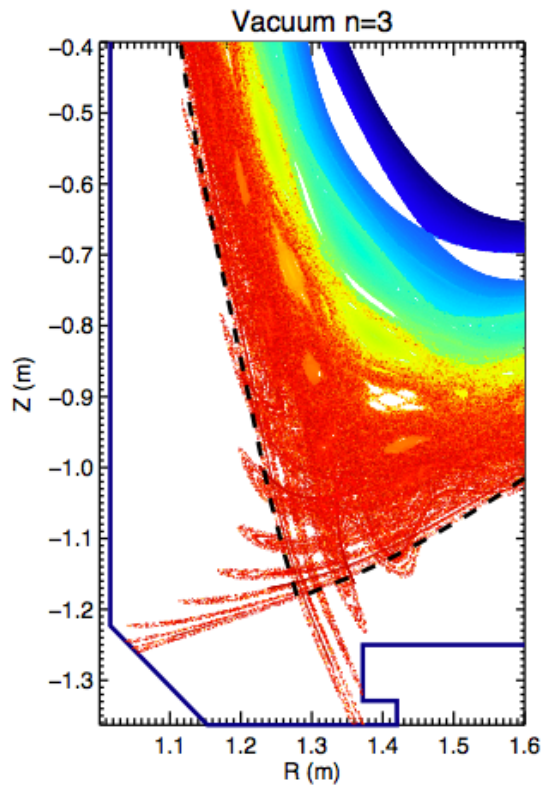
Example 1: NSTX n=3 mode

- Clear lobe structures are found from Poincare plot
 - Similar color indicates the similar initial flux surfaces (core: BLUE \rightarrow edge: RED)
 - Show field line splitting on the divertor plate
 - Indicates ideal plasma responses shield resonant components, preventing opening of islands and field line loss from inner region ($< \psi_n=0.97$)
 - Connection length indicates the fundamental feature of stochastic fields



Example 2: DIII-D n=3 mode

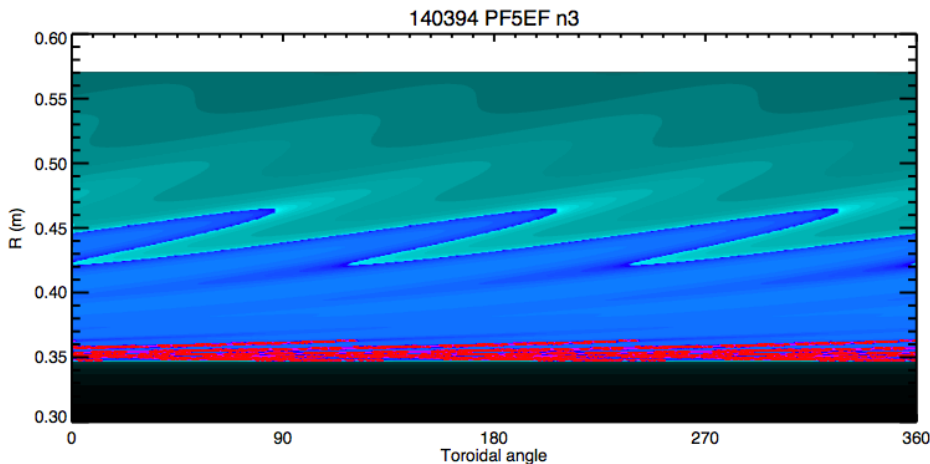
- Similar features by vacuum and plasma response are found
 - Stochastic region changes depending on the boundary for ideal plasma response ($\psi_{lim}=0.96, 0.99$)
 - Typical lobe structures in DIII-D n=3 are clearly obtained
 - Microscopic structure inside lobe is also captured



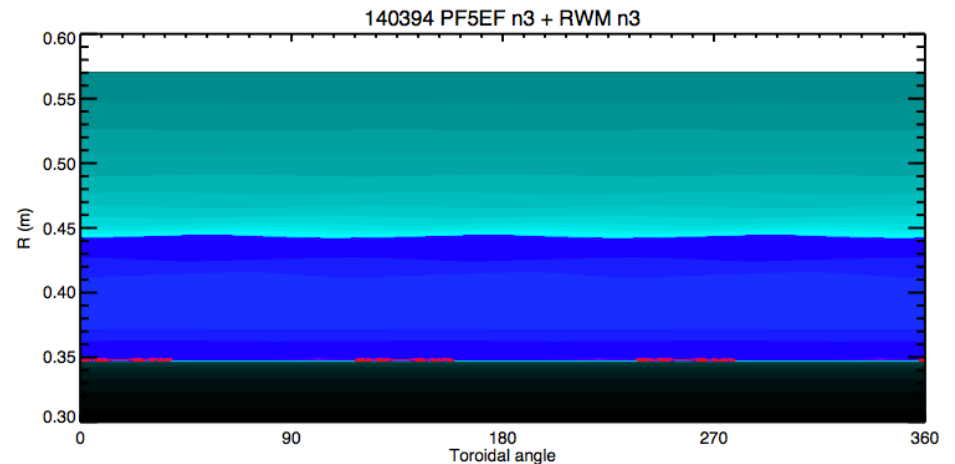
Example 3: $n=3$ Error Field Correction in NSTX

- Divertor footprint in NSTX EFC discharge shows a clear correction of $n=3$ error fields from PF5 coil
 - POCA-FLT takes multi-toroidal harmonics using IPEC inputs (i.e. $n=3$ by midplane, PF5 error field, $n=1$ by midplane, PF5/TF error fields)
 - Intrinsic $n=3$ error field by PF5 makes clear $n=3$ pattern
 - When error fields are corrected by midplane coils, split field lines almost disappear, indicating perfect error field correction

Intrinsic error field by PF5 ($n=3$)

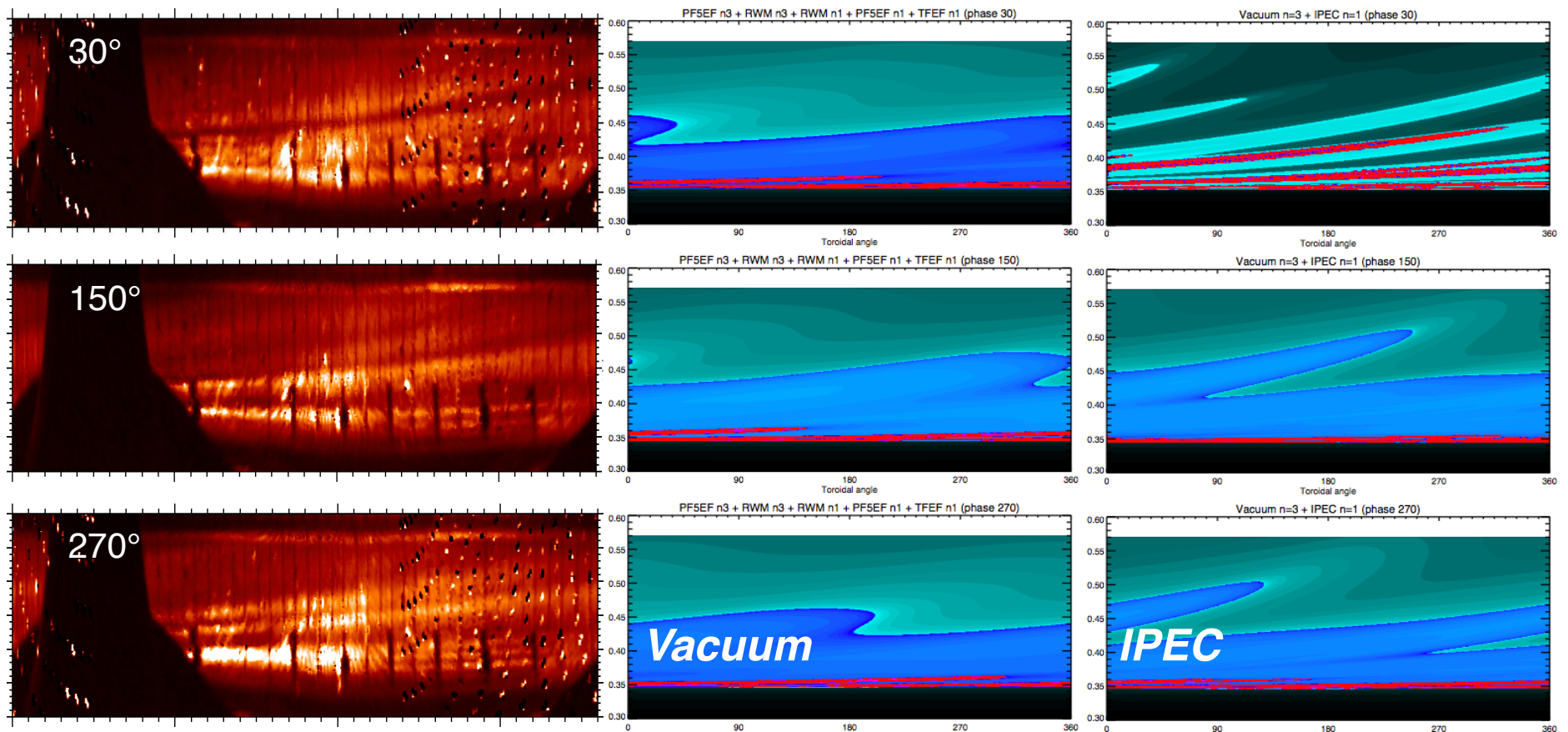


$n=3$ error field correction

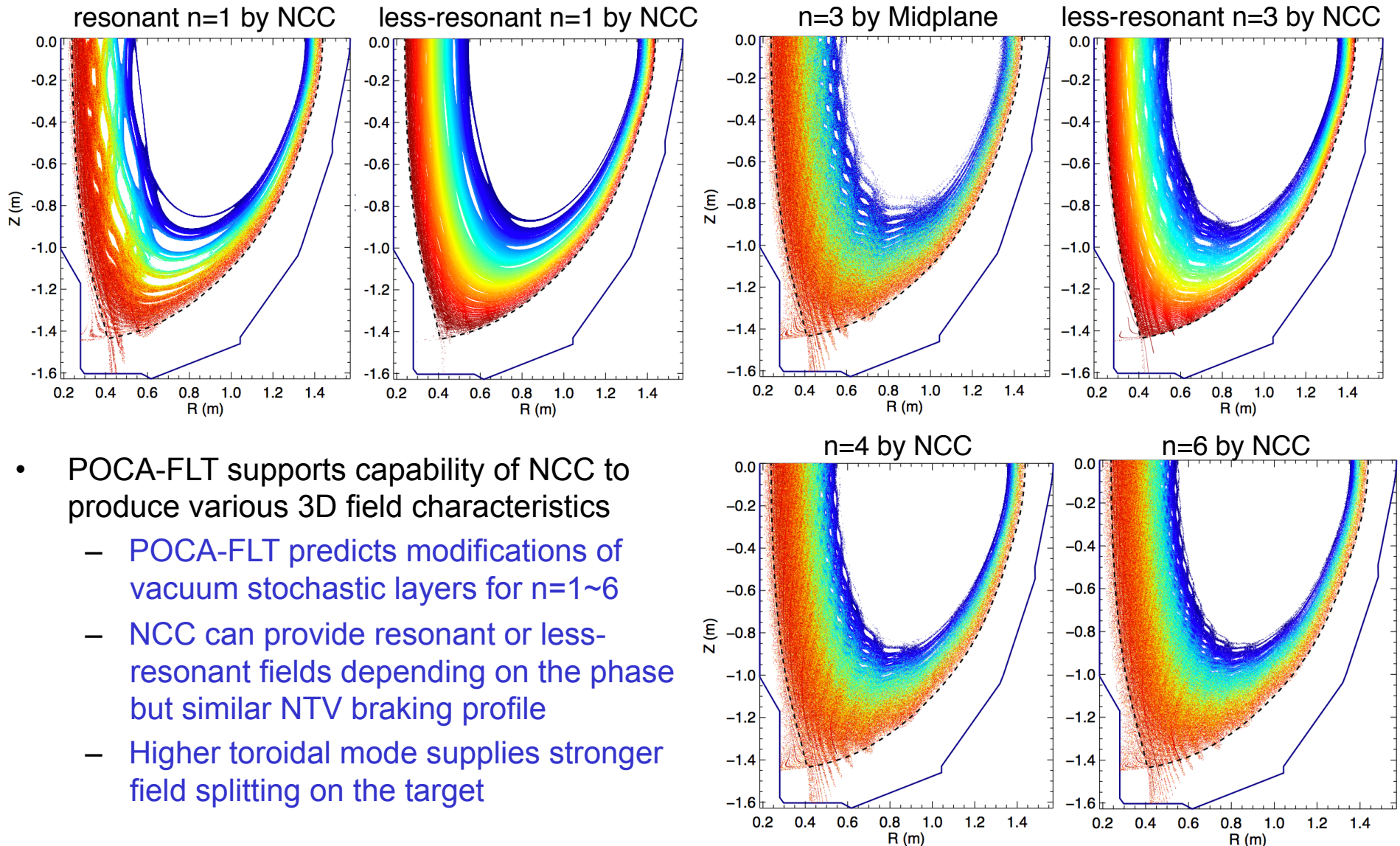


Example 4: POCA-FLT indicates the amplification of $n=1$ is responsible for the divertor footprint

- Additional splitting is generated by amplification of $n=1$ perturbations
 - Vacuum field cannot explain observed divertor footprint, but $n=1$ amplifications can do
 - Field amplification by ideal plasma response is well-known, and revealed by FLT
 - Accurate equilibrium required due to sensitivity of $n=1$ plasma response



Example 5: Field Line Tracing for NCC



- POCA-FLT supports capability of NCC to produce various 3D field characteristics
 - POCA-FLT predicts modifications of vacuum stochastic layers for $n=1\sim 6$
 - NCC can provide resonant or less-resonant fields depending on the phase but similar NTV braking profile
 - Higher toroidal mode supplies stronger field splitting on the target

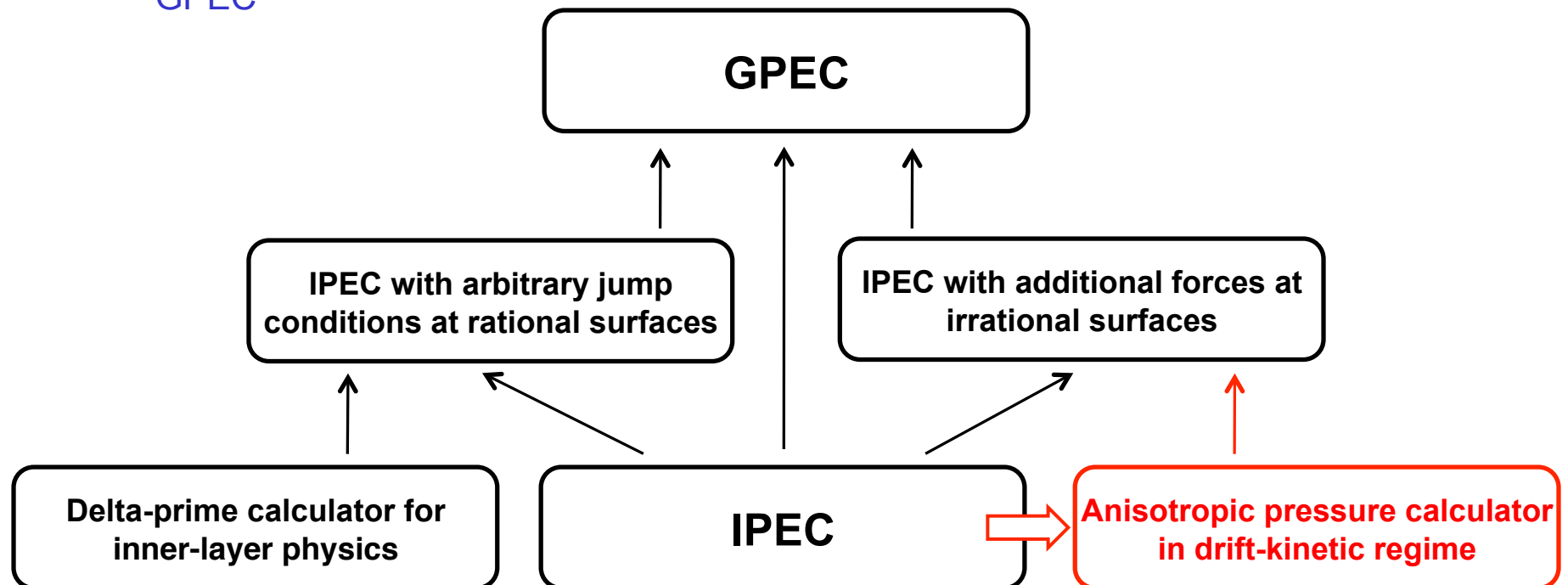
Summary

- POCA is being extensively applied to study NTV physics
 - Verify and validate theories, analyze and predict experiments, investigate new physics
- POCA studies 3D field structures via FLT
 - Well-established method, but unique capability to take 3D field of multi-toroidal harmonics, with or without plasma response
- There are many ongoing and future research topics
 - NTV prediction, analysis and field line tracing for NSTX-U NCC
 - TBM mock-up experiment analysis with high β plasmas in DIII-D
 - Ion/electron orbit tracing to study particle loss by 3D field
 - Calculate δW_k (equivalence of NTV- δW_k)
 - 3D field effects on (ion) bootstrap current

Back up

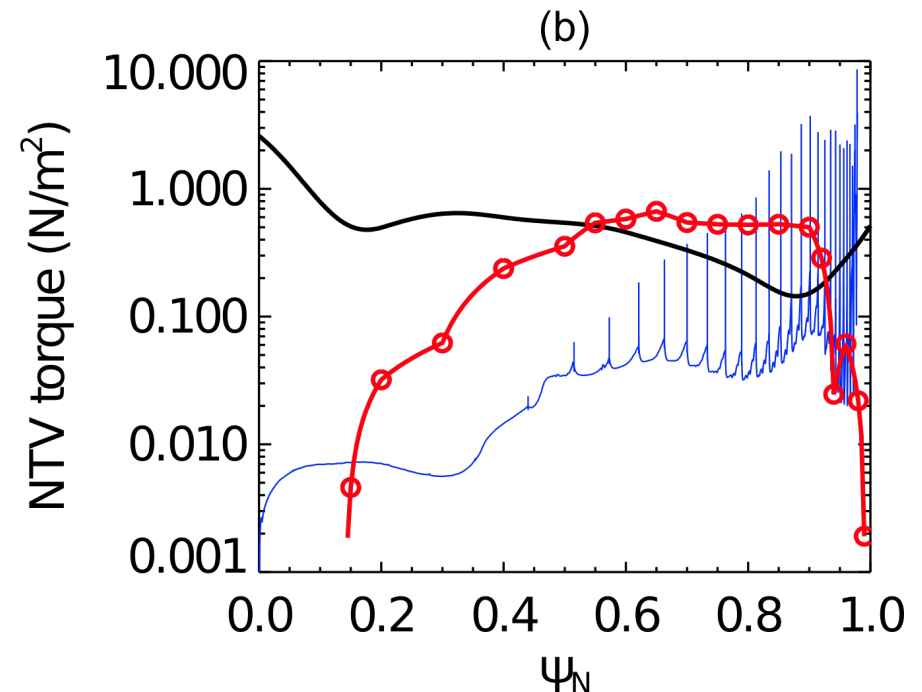
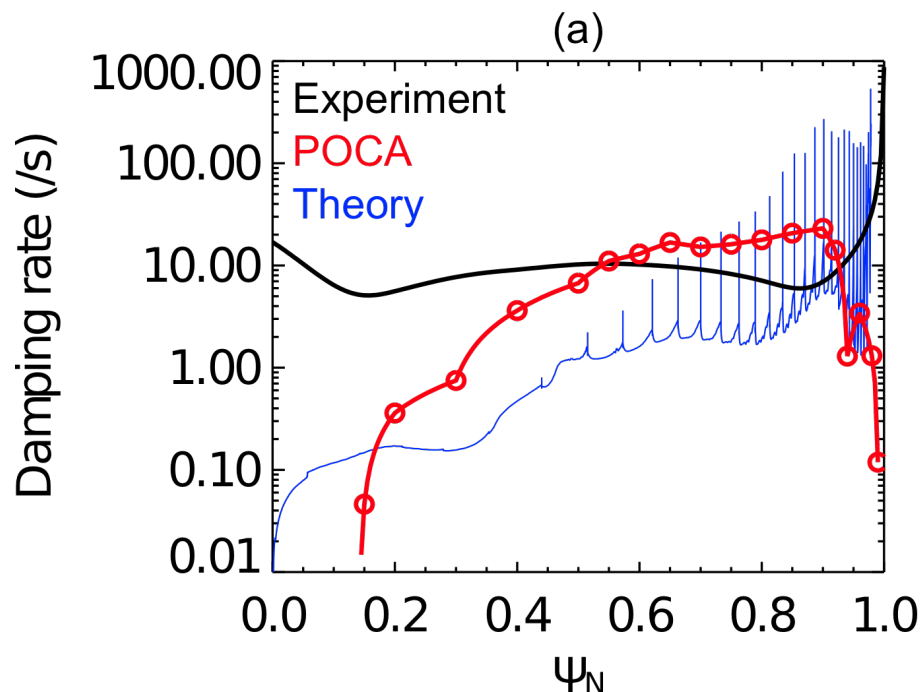
General Perturbed Equilibrium Code (GPEC) is on progress

- Perturbed equilibrium codes are efficient to study 3D field physics in tokamaks with non-axisymmetric perturbations
 - IPEC solves ideal force balance with ideal constraints
 - GPEC will solve non-ideal force balance with arbitrary jump conditions, which will be matched with inner-layer solver
 - POCA will use 3D perturbations from IPEC, and provide anisotropic pressure tensor to GPEC



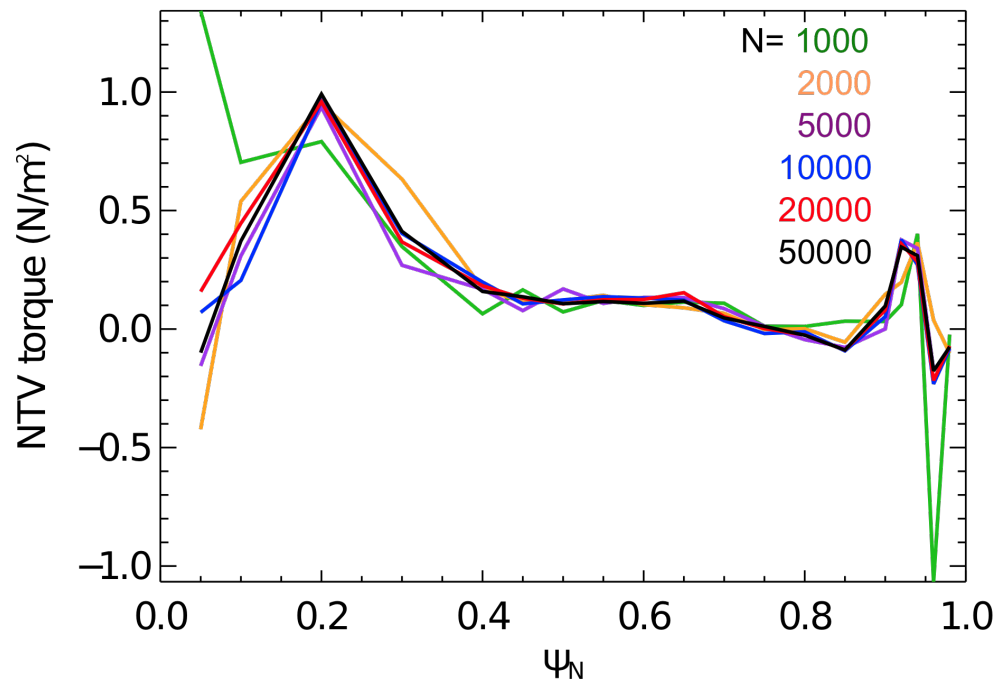
POCA improved a prediction on NTV in NSTX -2

- NTV is calculated for NSTX error field correction experiment
 - Selected discharge 132729 is a case of $I_{EFC}=750A$, which produced a strong magnetic braking ($I_p=1.1MA$, $B_{T0}=0.55 T$) [S.P. Gerhardt, PPCF (2010)]
 - Discrepancies are found in damping and NTV profiles: POCA predicts weaker NTV at inner and edge region and stronger NTV elsewhere
 - Total NTV torque still agrees well: **Experiment 5.1 Nm / POCA 3.5 Nm (#132729)**
 - Even though large discrepancies with measurement, POCA and theory show a similar profile shape; Check would be necessary for interpretation of measurements



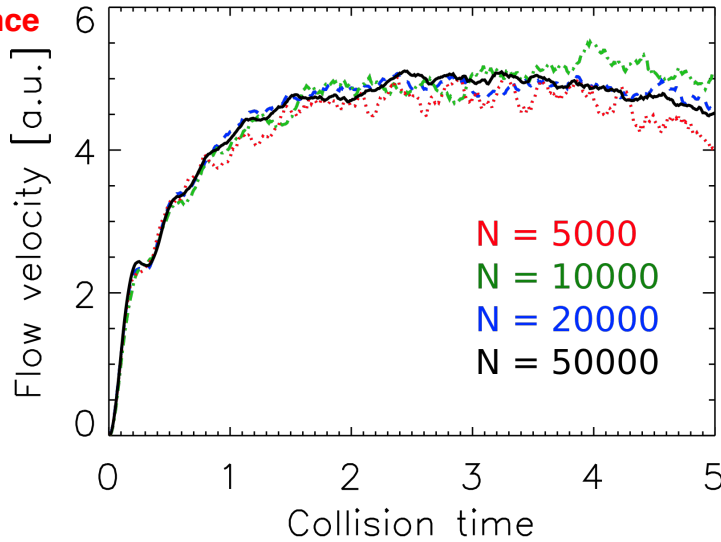
POCA shows a good convergence even in experiment

- Convergence has been tested by scanning test particle number
 - δf of particle simulation is acceptable, but still needs improvement for efficiency
 - $N=10^3 \sim 5 \times 10^5$ have been tested for DIII-D #145117
 - Good agreements with N over $\sim 5 \times 10^3$ have been found, indicating the computational efficiency can be improved using small number of test particles
 - Most time-consuming part is an interpolation of magnetic field, which will be considered for future upgrade

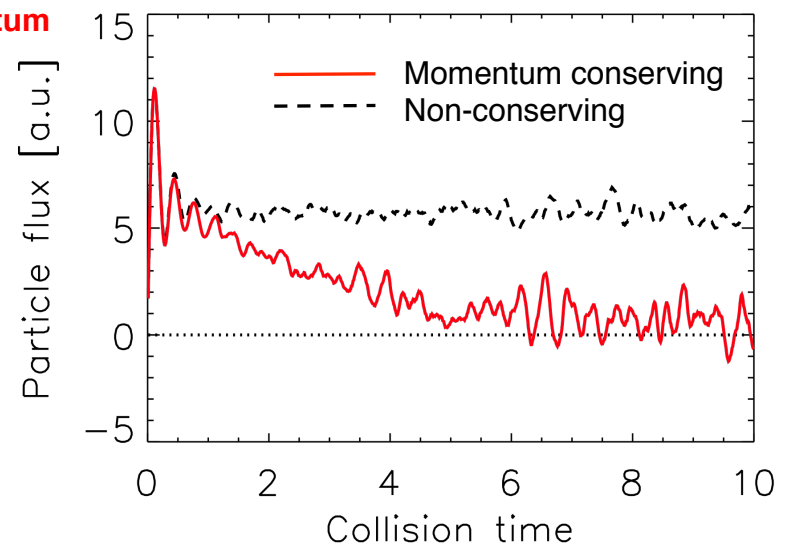


POCA has been successfully benchmarked in Axisymmetry

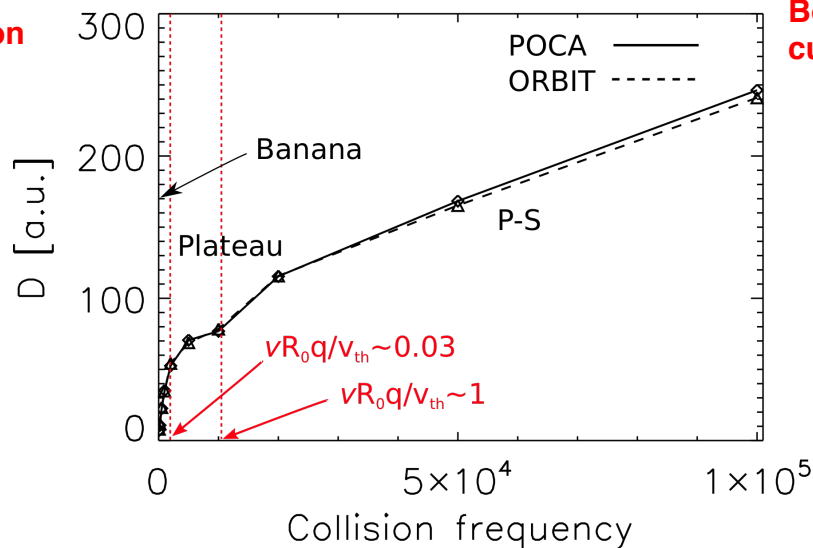
Convergence



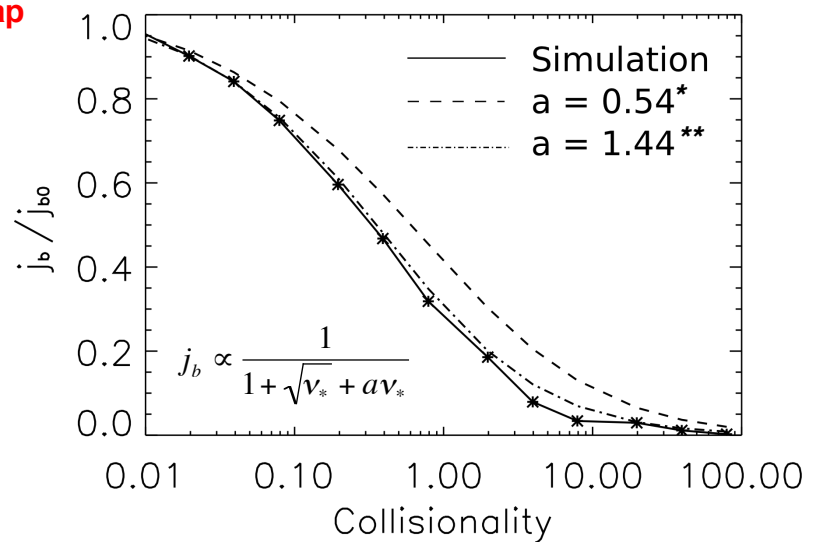
Momentum



Diffusion



Bootstrap current



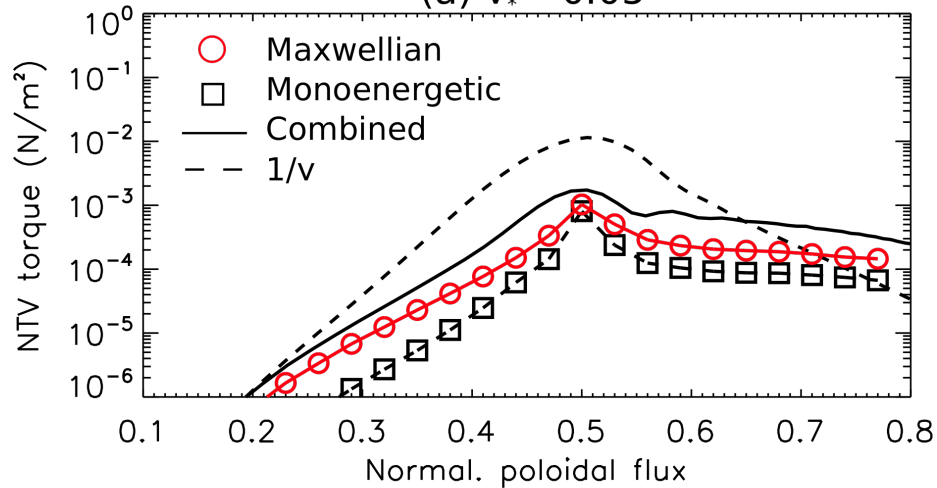
[R.B. White and M.S. Chance, Phys. Fluids B 27, 2455 (1984)]

*[F.L. Hinton and M.N. Rosenbluth, Phys. Fluids 16, 836 (1973)]

**[M. Sasinowski and A.H. Boozer, Phys. Plasmas 2, 610 (1995)]

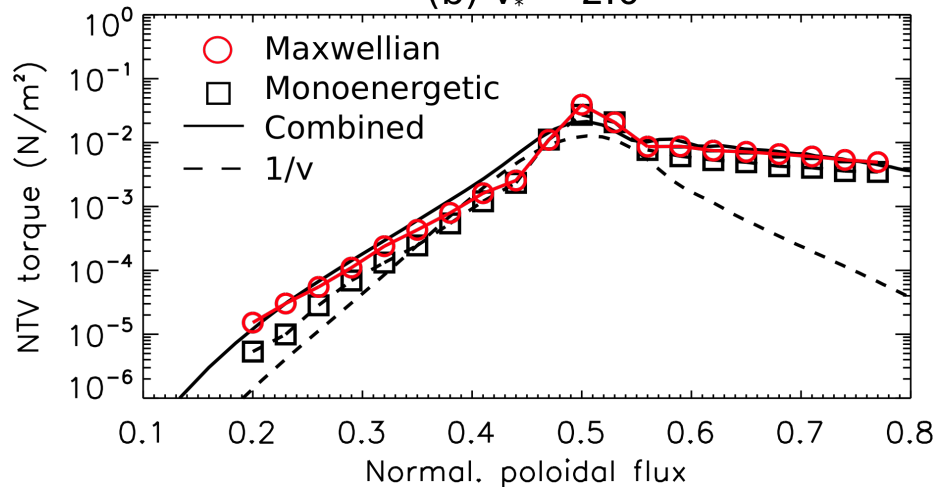
POCA was compared with $1/\nu$ theory

(a) $\nu_* \sim 0.05$



- NTV approaches $1/\nu$ regime as collisionality increases
 - $1/\nu$ formula indicates stronger resonance but weaker non-resonance
 - Magnetic precession and regime overlapping by Maxwellian energy distribution in POCA and combined formula cause broader NTV profiles than $1/\nu$ formula

(b) $\nu_* \sim 2.0$



- High energy particle impacts on NTV
 - High energy particles in the Maxwellian tails strongly impact at the non-resonant flux surfaces
 - In the high collisionality, collisions are found to become more dominant than the high energy particle effects

Bounce-harmonic resonance has been found for KSTAR

- Strong NTV was predicted in a particular level of high rotation, due to the 1st bounce-harmonic resonance, and found in KSTAR

