

# Development of a reduced model for resonant fast ion transport in TRANSP

**M. Podestà**

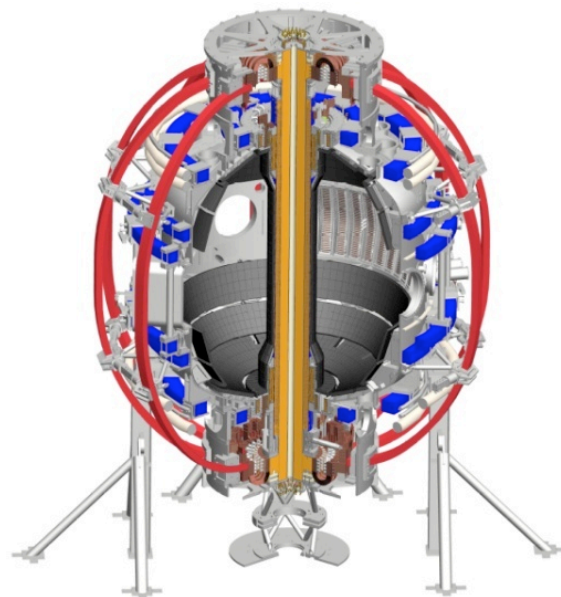
M. Gorelenkova, R. B. White

*and many other contributors*

**NSTX-U Monday Meeting**

**B238, PPPL**

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# Four models are presently implemented in TRANSP for fast ion diffusion/convection coefficients, $D_b$ & $v_b$

All have diffusive/convective nature in radial coordinate; little/no phase space selectivity:

[from <http://w3.pppl.gov/~pshare/help/transp.htm>]

$$1) D_b = k_{ADIFB} \times D_e$$

$k_{ADIFB}$  : multiplier

$D_e(x,t)$  : electron particle diffusivity

$$2) D_b = k_{ADIFB} \times D_e^{WP}$$

$D_e^{WP}(x,t)$  : electron particle diffusivity from Ware-pinch corrected flux

$$3) \Gamma_{fi} = -D_b \nabla n_b + v_b n_b$$

*diffusion/convection model*

$$4) D_b(E, t, x) = \sum_k \alpha_k D_k$$

$D_k(E,x,t)$ : diffusivity for deeply trapped, barely trapped, co-barely passing, ...

# Proposed model introduces selectivity in phase space, generalizes “diffusive transport” *ad-hoc* models

- Info on phase space dynamics is of paramount importance for Verification&Validation of codes, theory-experiment comparison
- “Fluid” (integral) quantities do not provide all information we need
  - Re-computed (“inverted”) solutions for  $F_{nb}$  based on measured integral quantities (f.i. density, neutrons,  $E_r$ /rotation, NB-driven  $J_{nb}$ , ...) **are not unique**
- ***Need details on fast ion energy, pitch and their (consistent) evolution***
  - E.g. \*AE bursts transiently (and **selectively**) modify phase space; effects propagate during slowing-down

*The new model must be “simple enough” to be included in TRANSP/NUBEAM for routine use —→ **reduced model***

# Outline

- New ‘kick model’ : basic ideas
- Implemented algorithm
- Initial validation against full ORBIT simulations
  - Example from NSTX case w/ TAE avalanche
- Additional remarks and summary

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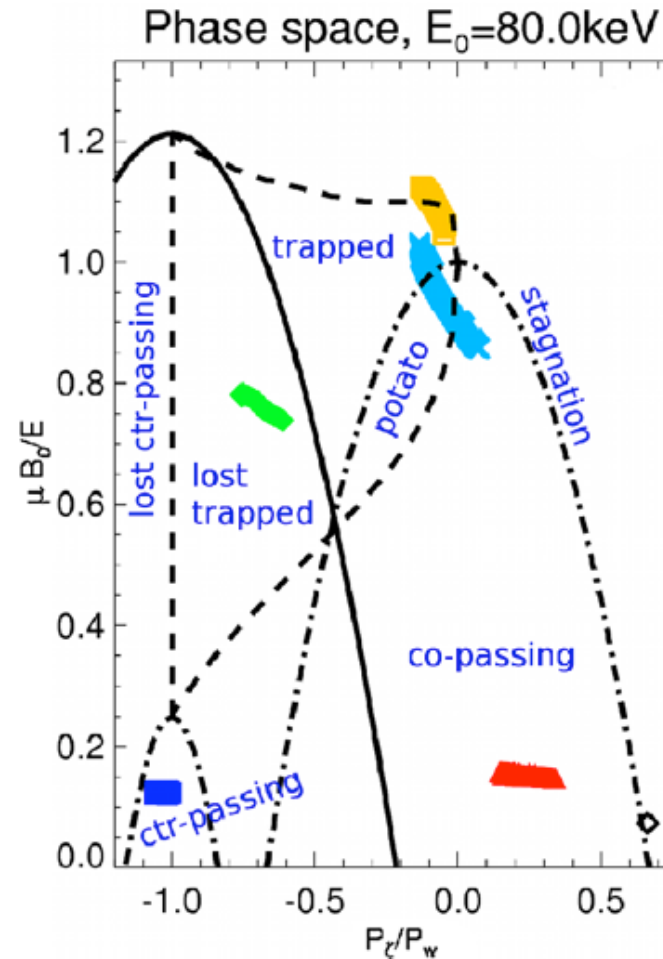
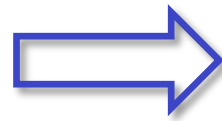
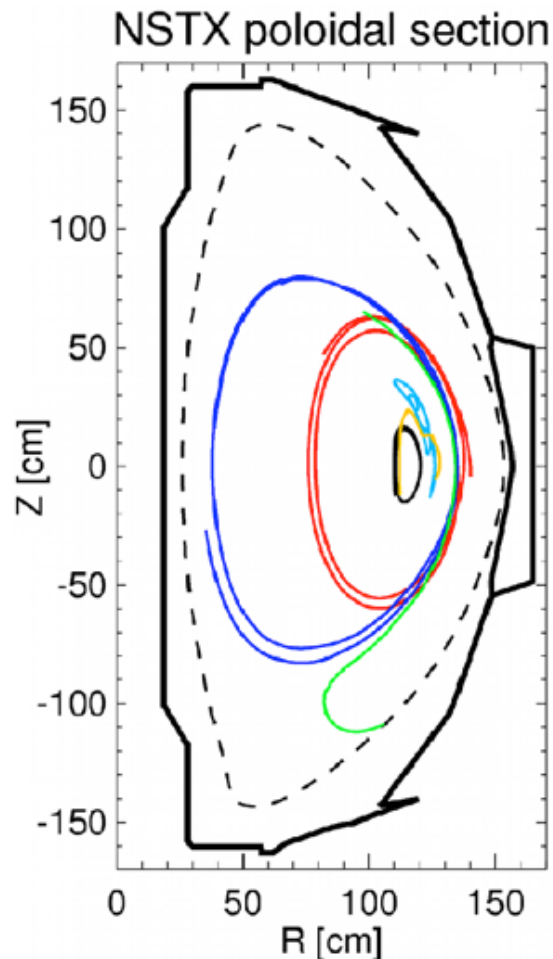
# 'Constants of motion' are the natural variables to describe resonant wave-particle interaction

Each orbit fully characterized by:

- $E$ , energy
- $P_\zeta \sim mRv_{\text{par}} - \Psi$ , canonical angular momentum
- $\mu \sim v_{\text{perp}}^2 / (2B)$ , magnetic moment

Energy exchange rate:

$$\gamma \propto \omega \frac{\partial F_{nb}}{\partial E} + n \frac{\partial F_{nb}}{\partial P_\zeta}$$



R. B. White, Theory of toroidally confined plasmas, Imperial College Press (2001)

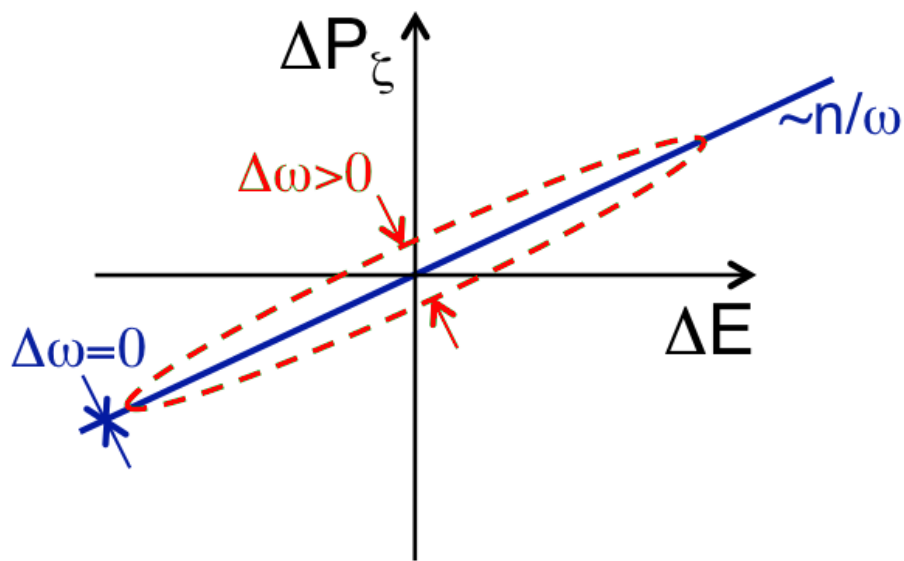
# 'Kick Model' is based on a *probability distribution function* for particle transport

- Resonances introduce fundamental constraints on particle's trajectory in  $(E, P_\zeta, \mu)$

- From Hamiltonian formulation – single resonance:

$$\omega P_\zeta - nE = \text{const.} \implies \Delta P_\zeta / \Delta E = n/\omega$$

$\omega = 2\pi f$ , mode frequency       $n$ , toroidal mode number



*For each bin in  $(E, P_\zeta, \mu)$ , steps in  $\Delta E, \Delta P_\zeta$  are described by*

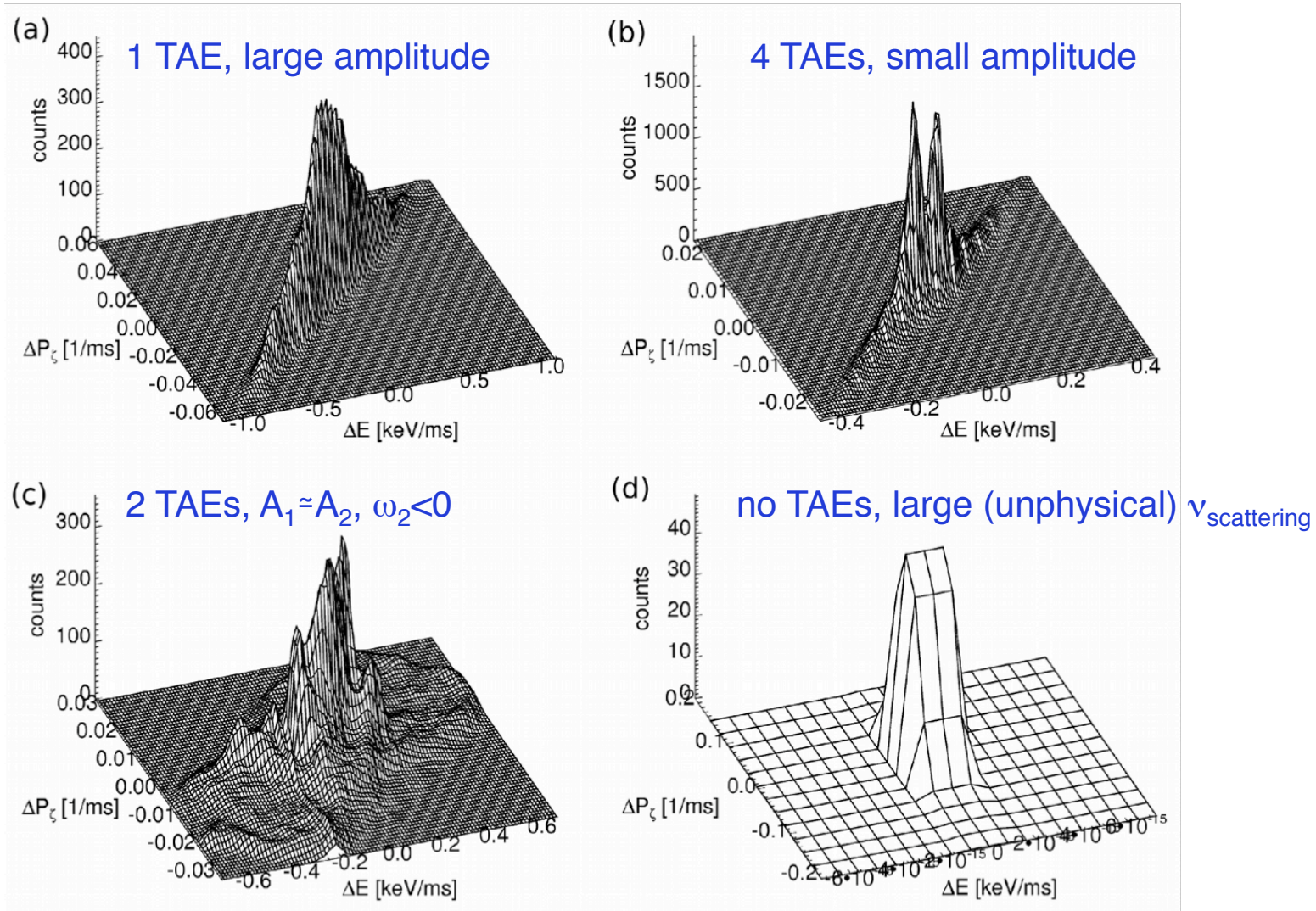
$$p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu, A)$$

*which can incorporate the effects of multiple modes & resonances:*

***Correlated random walk***

# An analytical formulation for $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu)$ could be developed – but it would be quite unpractical

E.g., use:  $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu) = \sum_{i=1}^N w_i p_i(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu)$



**In practice, will use the full 5D matrix for  $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu)$**



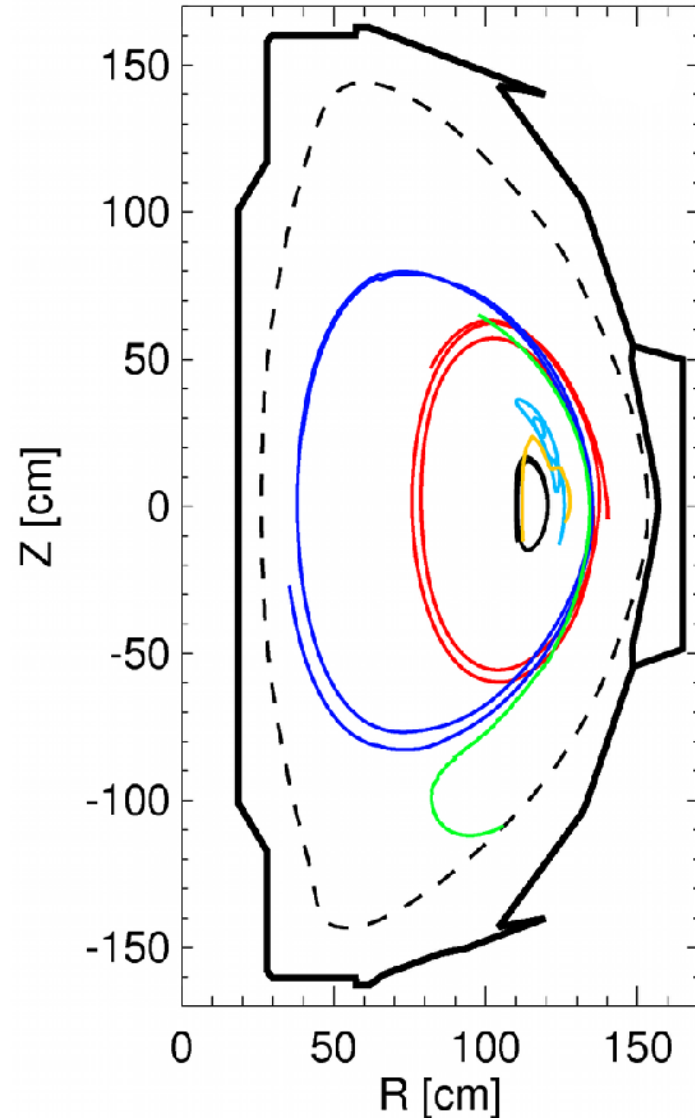
# A quick summary

Two main ingredients for new model:

- Probability distribution function for particle's "kicks" in  $E, P_\xi$
- Mode amplitude scaling factor  $A_{\text{mode}}(t)$

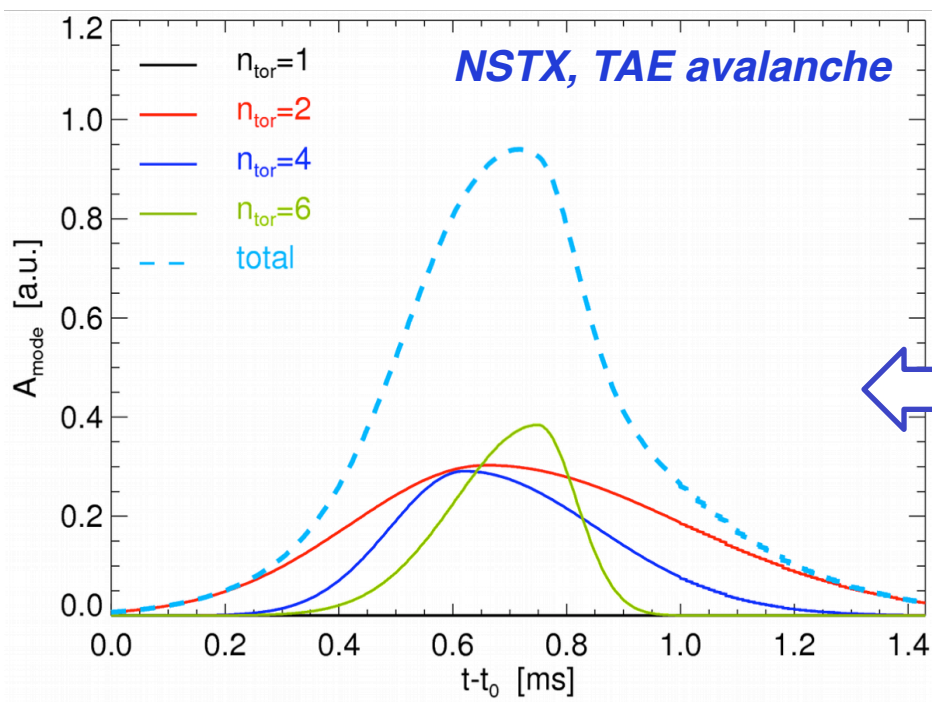
# 'Transport probability' $p(\Delta E, \Delta P_\xi | P_\xi, E, \mu)$ can be computed through numerical codes (e.g. ORBIT) or theory

NSTX poloidal section



- Run ORBIT with  $A_{\text{mode}}=1$ , constant
- Simulation time long enough ( $\sim 1$ ms, many toroidal transit times) to capture \*AE effects
- Track  $(P_\xi, E, \mu)$  in time for each particle, steps  $\delta t_{\text{sim}} \sim 25-50 \mu\text{s}$
- Compute  $\Delta E, \Delta P_\xi, \Delta \mu$
- Re-bin over  $(P_\xi, E, \mu)$  space
- > Get  $p(\Delta E, \Delta P_\xi | P_\xi, E, \mu)$  for each bin

# Mode amplitude can evolve on time-scales shorter than typical TRANSP/NUBEAM steps of ~5-10 ms

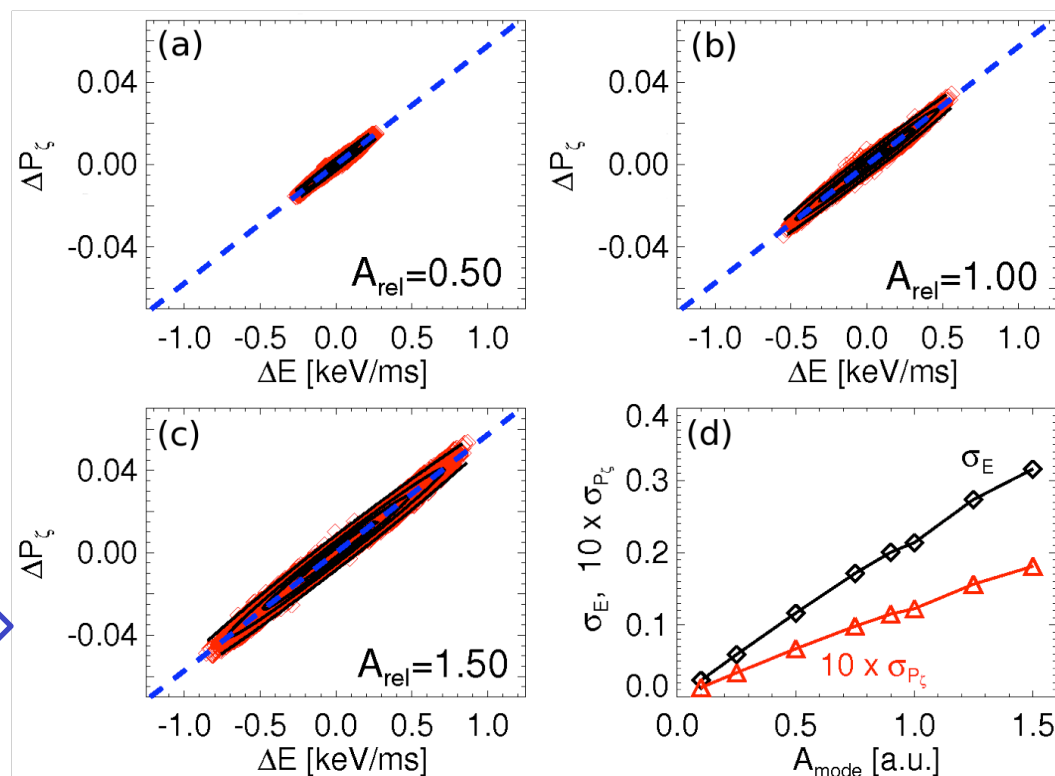


$F_{nb}$  evolution must be computed as a sequence of sub-steps

- Duration  $\delta t_{\text{step}}$  sufficiently shorter than time-scale of mode evolution
- Examples here have  $\delta t_{\text{step}} \sim 25\text{-}50 \mu\text{s}$

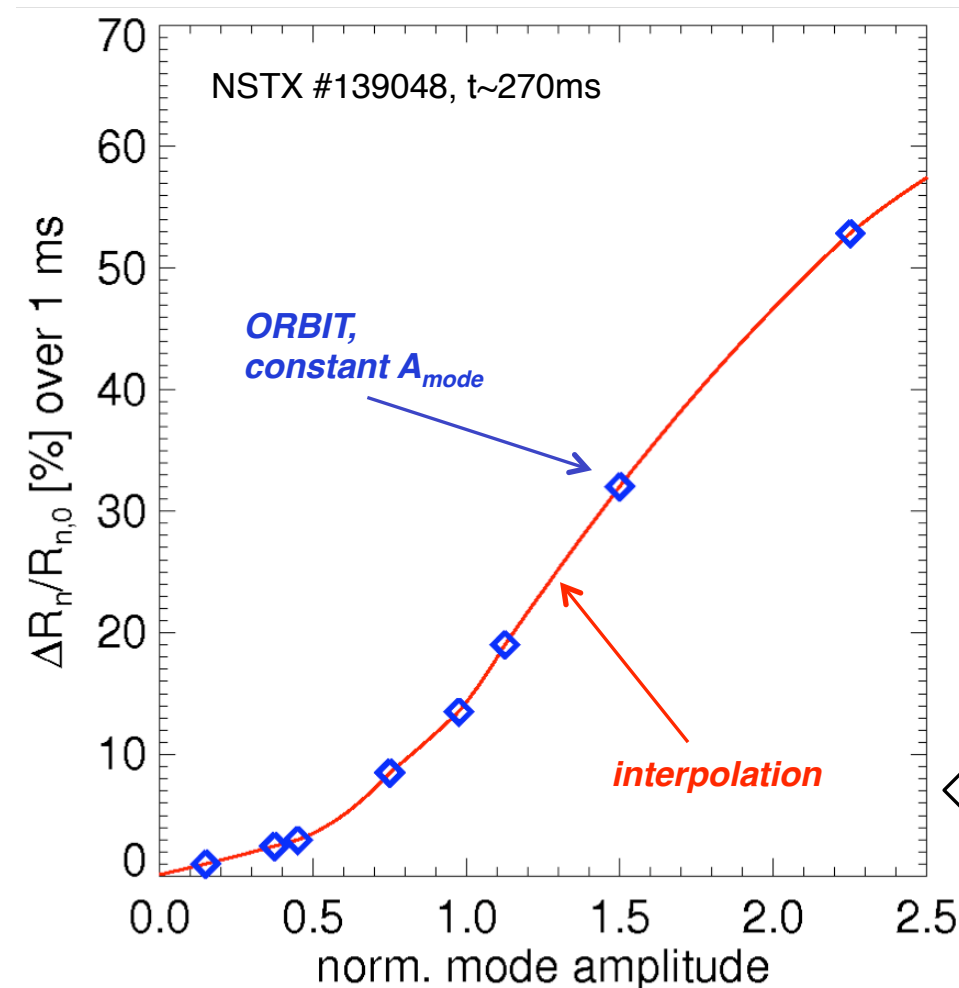
Energy and  $P_{\xi}$  steps assumed to scale linearly with mode amplitude

- Roughly consistent with ORBIT simulations



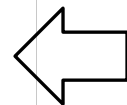
# Scaling factor $A_{mode}(t)$ is obtained from measurements, *observables* such as neutron rate + modeling

- Best option: use experimental data (e.g. reflectometers, ECE)
- If no mode data available,  $A_{mode}$  can be estimated based on other measured quantities



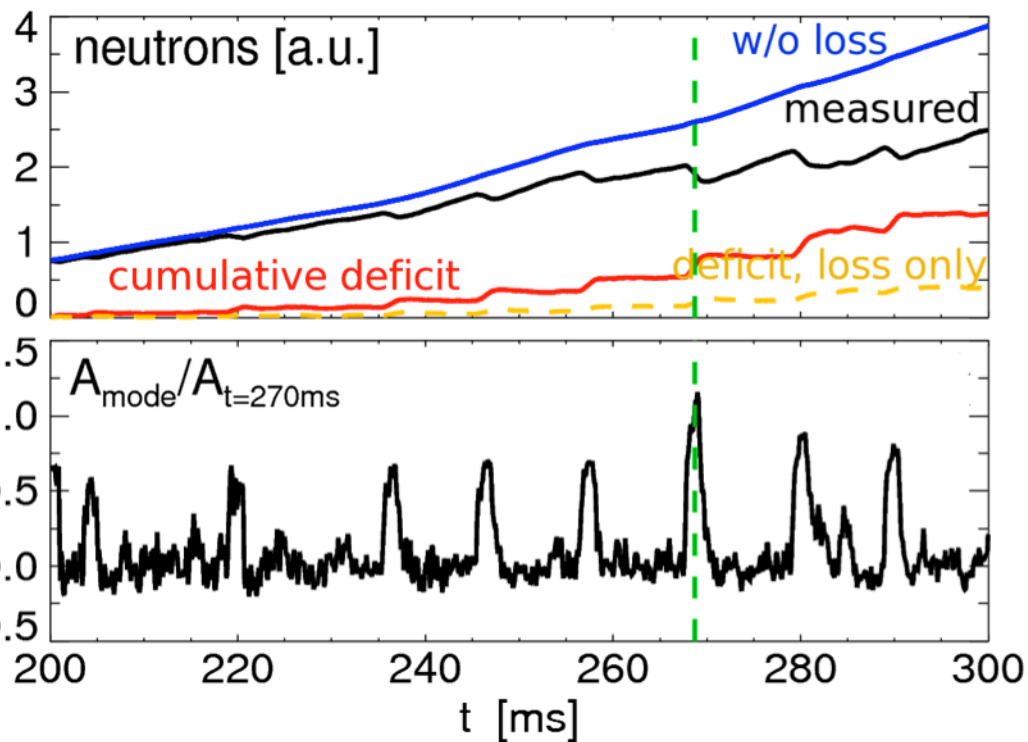
## - Example: use measured neutron rate

- Compute ideal modes through NOVA
- Rescale relative amplitudes from NOVA according to magnetics
- Rescale total amplitude based on computed neutron drop from ORBIT
- Scan mode amplitude w.r.t. experimental one,  $A_{mode}=1$ : get table
- Build  $A_{mode}(t)$  from neutrons vs. time, table look-up



# Example: $A_{mode}(t)$ computed from measured neutron rate or Mirnov coils' signal

Get  $A(t)$  from measured neutrons+table look-up:

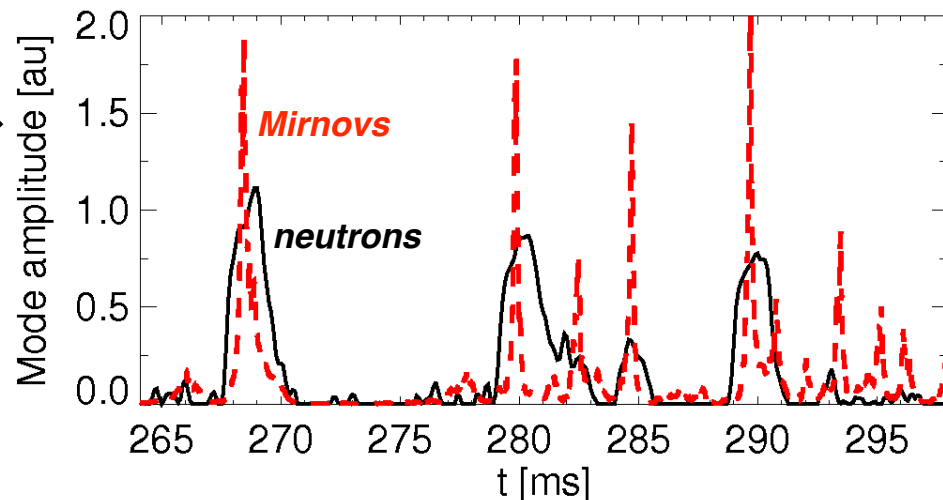


– Compute fractional  $R_n$  drops vs. time

– Use table  $R_n$  vs  $A_{mode}$  to find corresponding (normalized) mode amplitude

– Comparison w/  $A_{mode}(t)$  from Mirnov coils

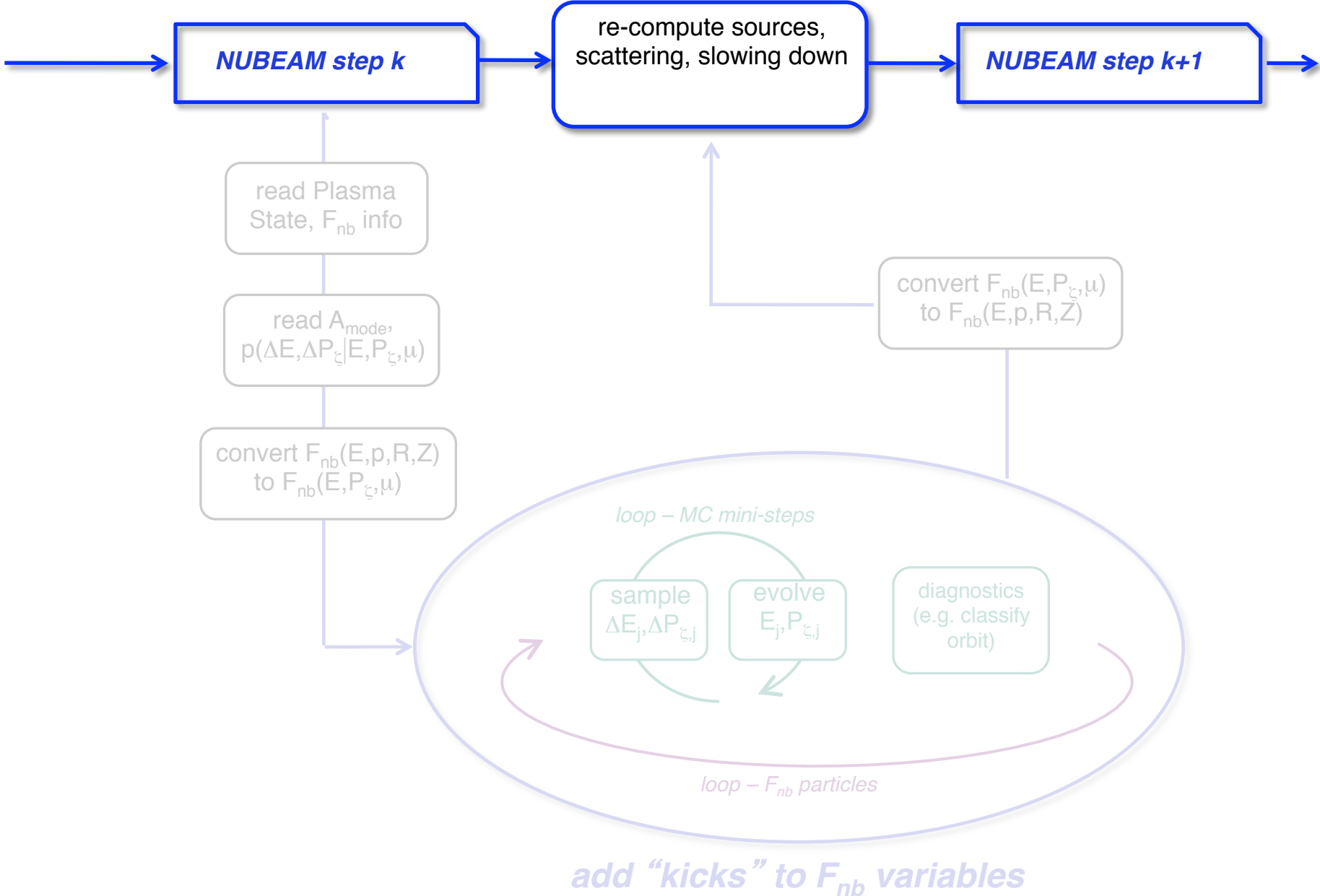
– Do different waveforms lead to differences in fast ion evolution?  
*Not on relevant time scales > 1ms*



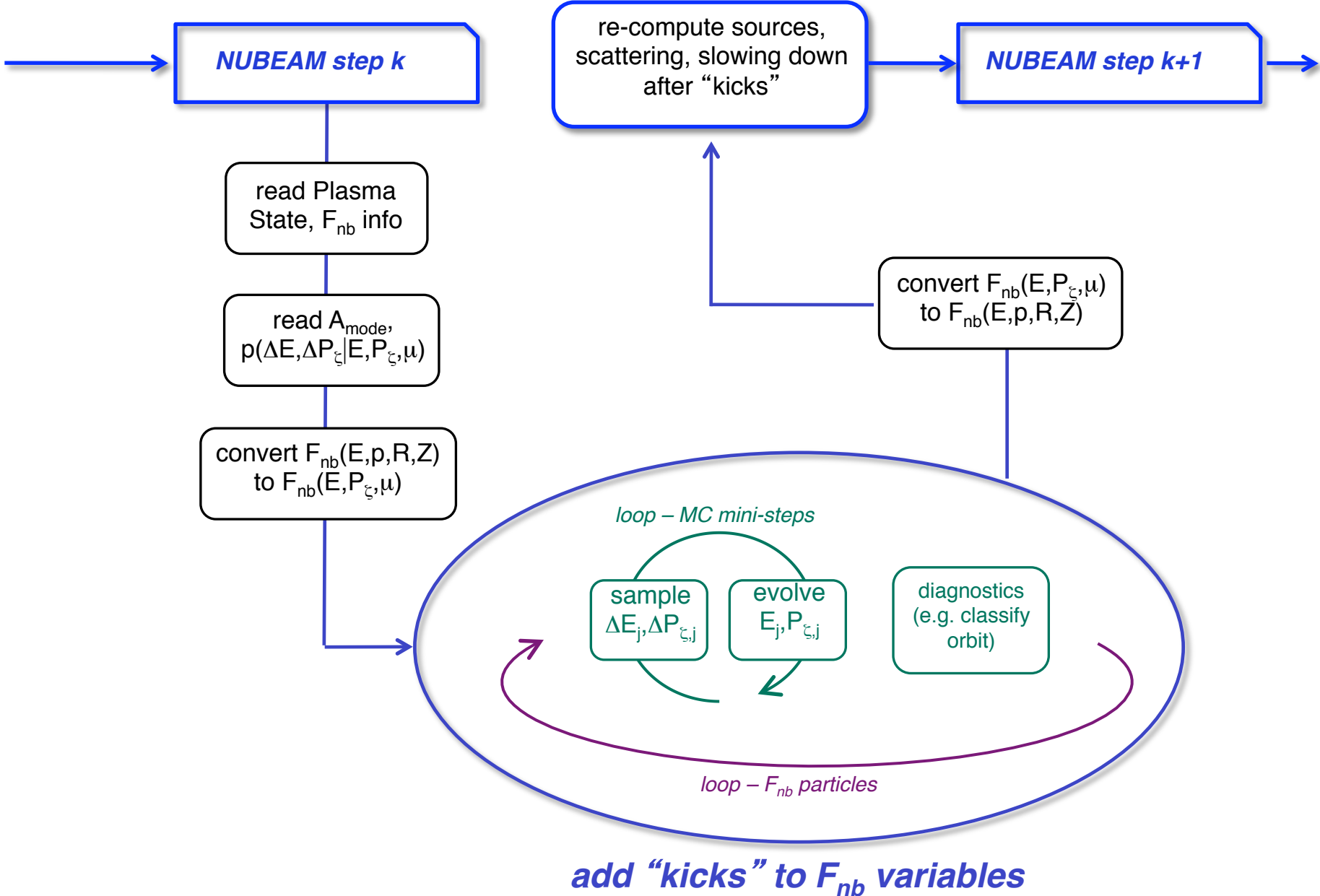
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- New ‘kick model’ : basic ideas
- **Implemented algorithm**
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# Scheme to advance fast ion variables according to transport probability in NUBEAM module of TRANSP



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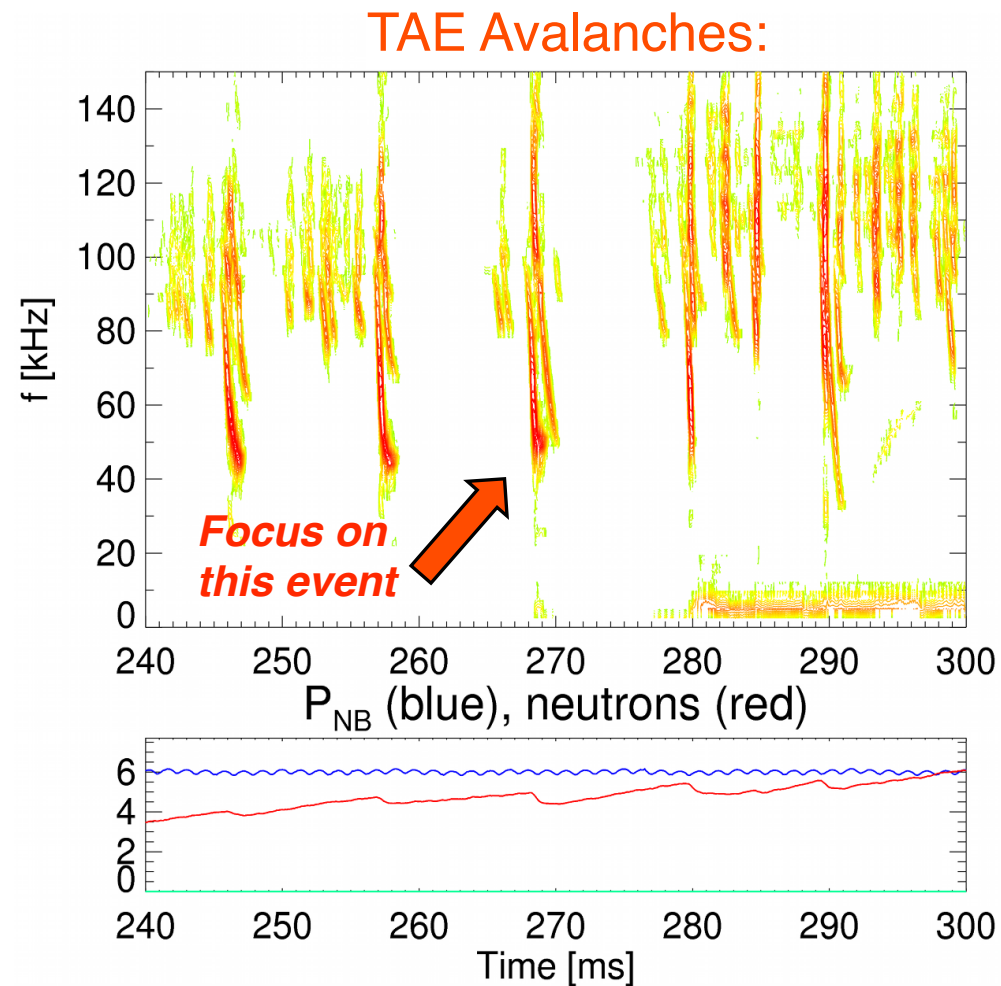
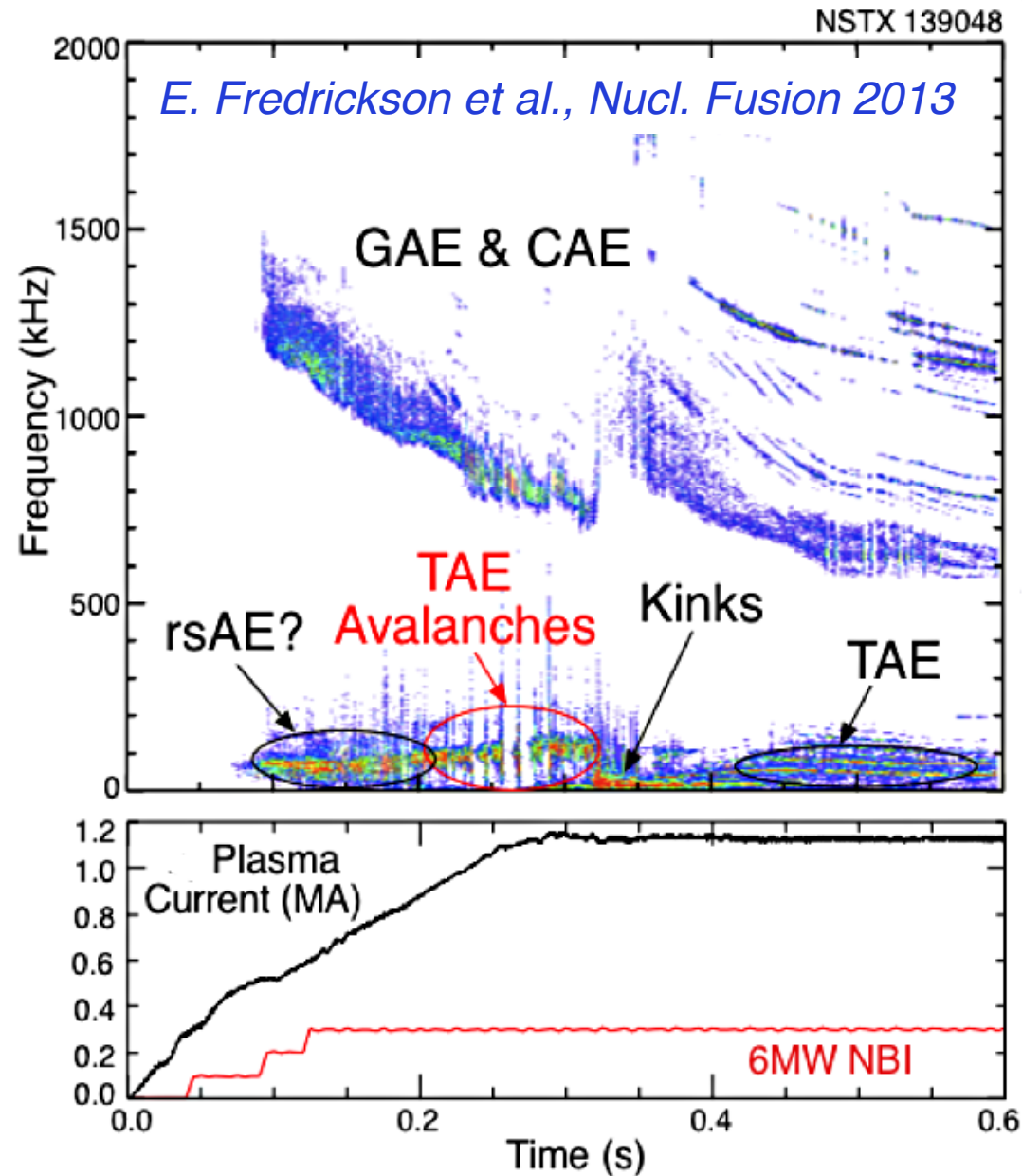




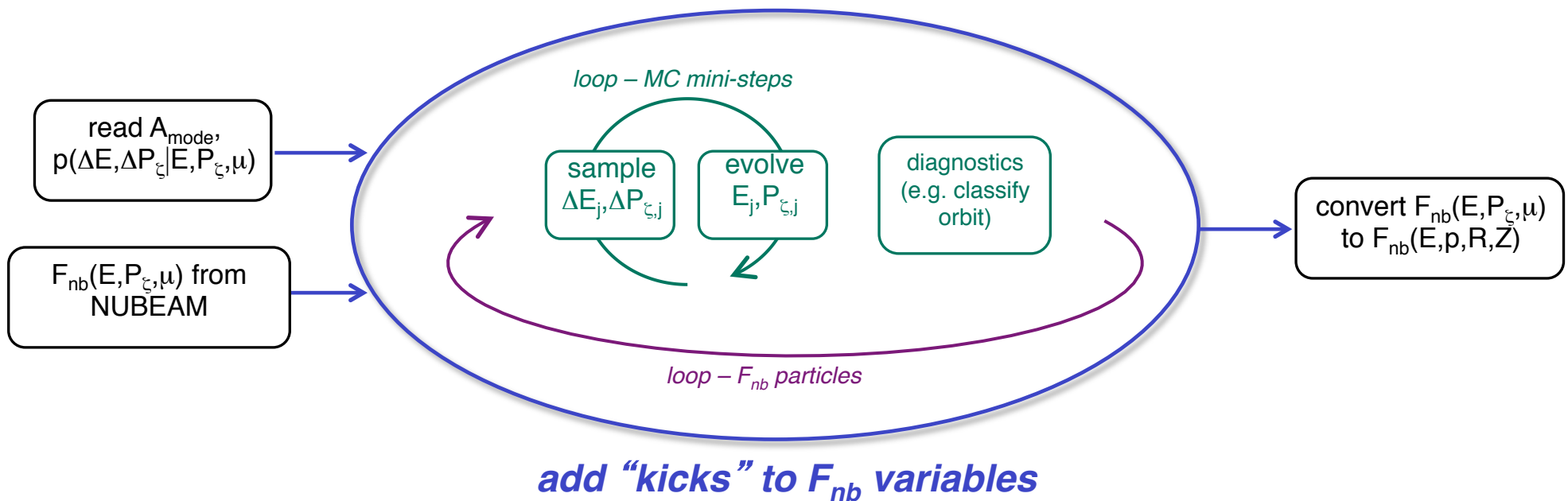
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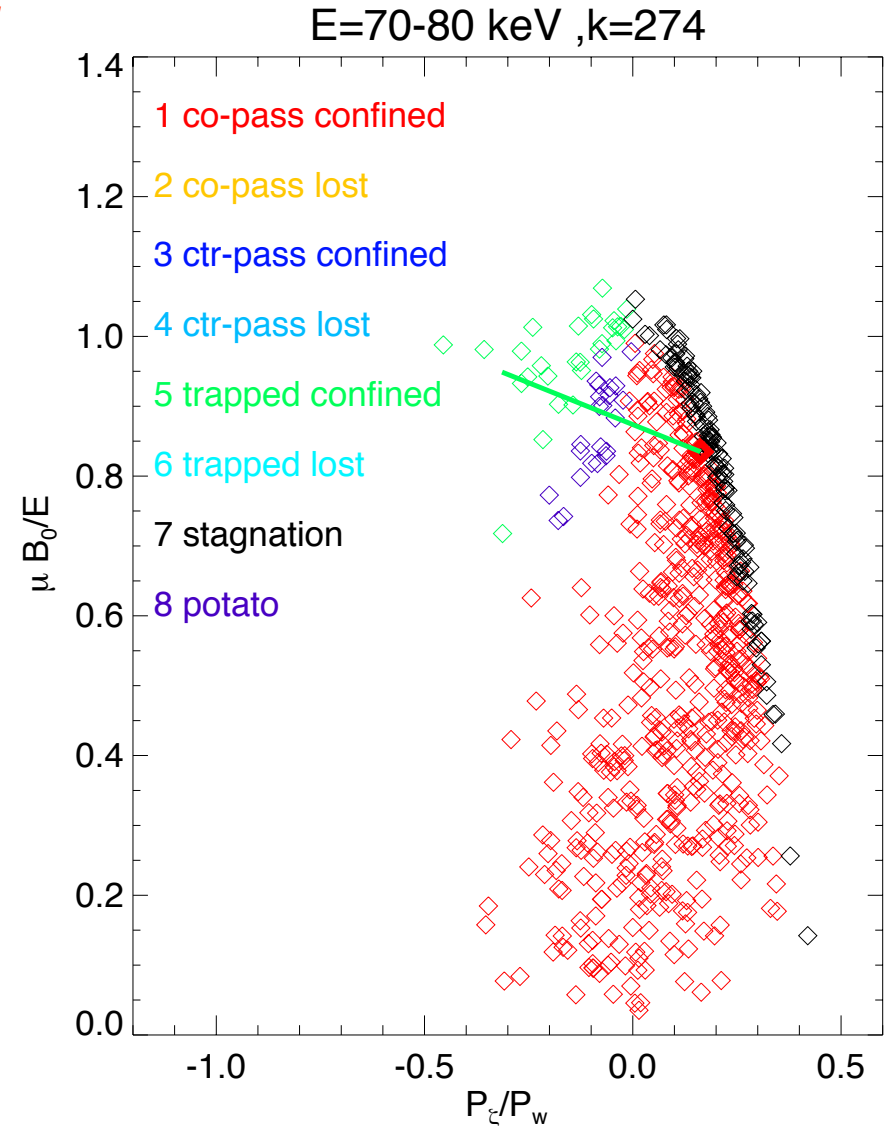
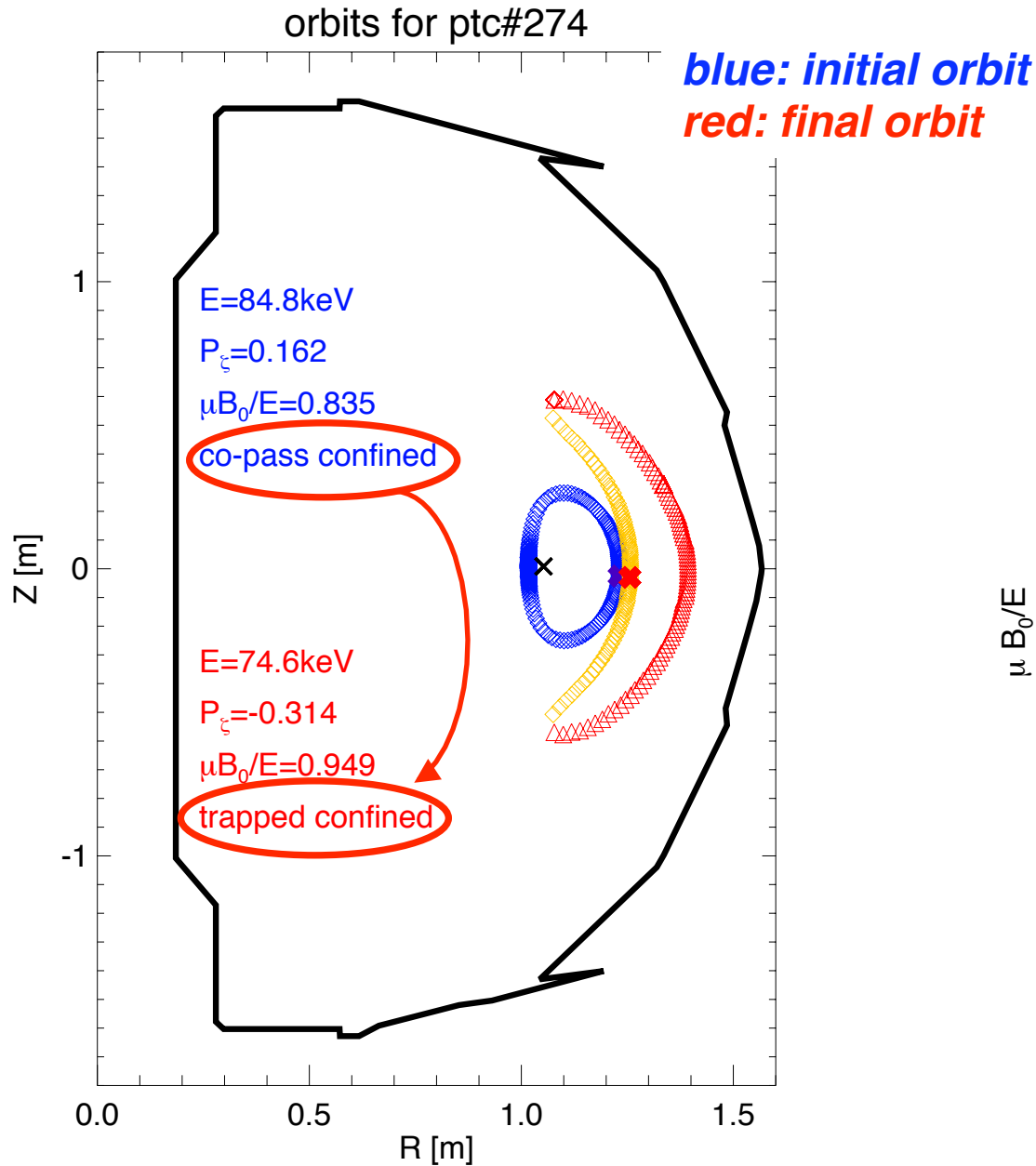
# Experimental scenario for initial validation: NSTX H-mode plasma with bursts of TAE activity



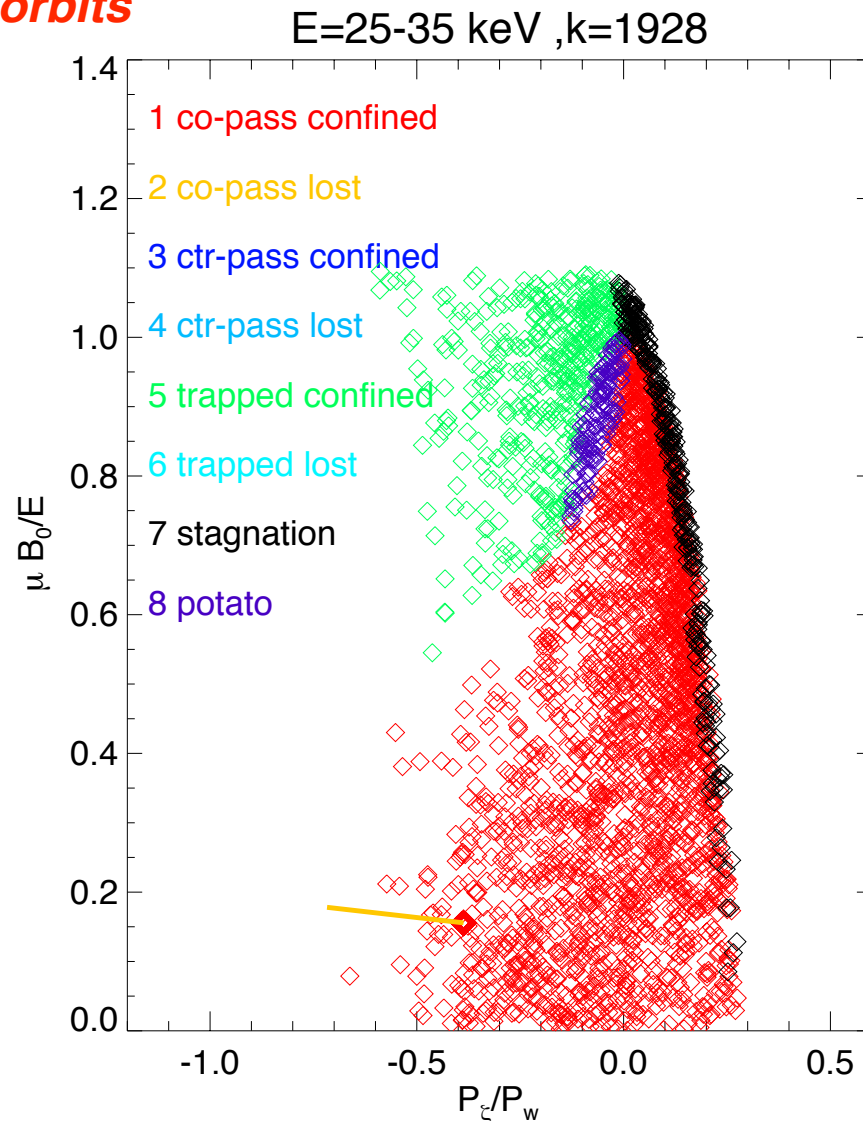
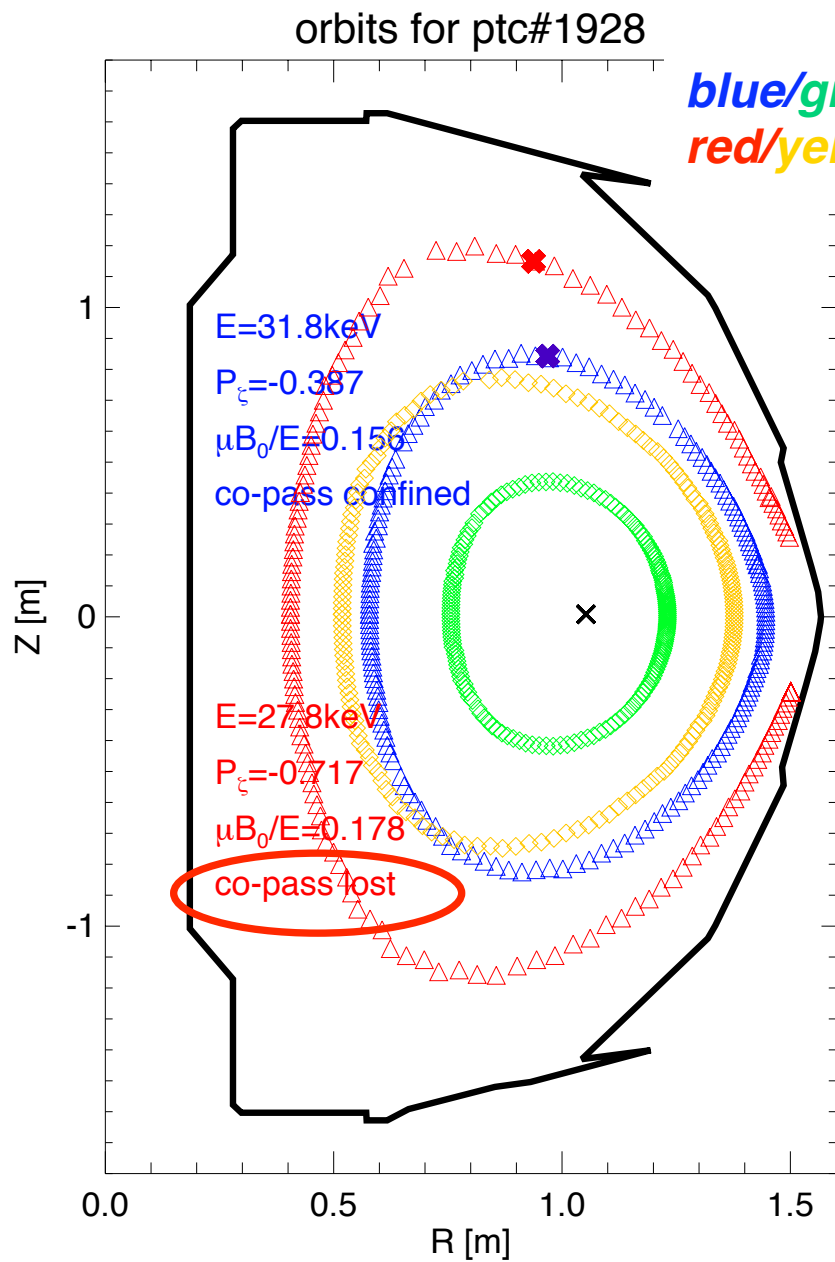
# First tests: evolve $F_{nb}$ from NUBEAM assuming fixed background, compare with ORBIT



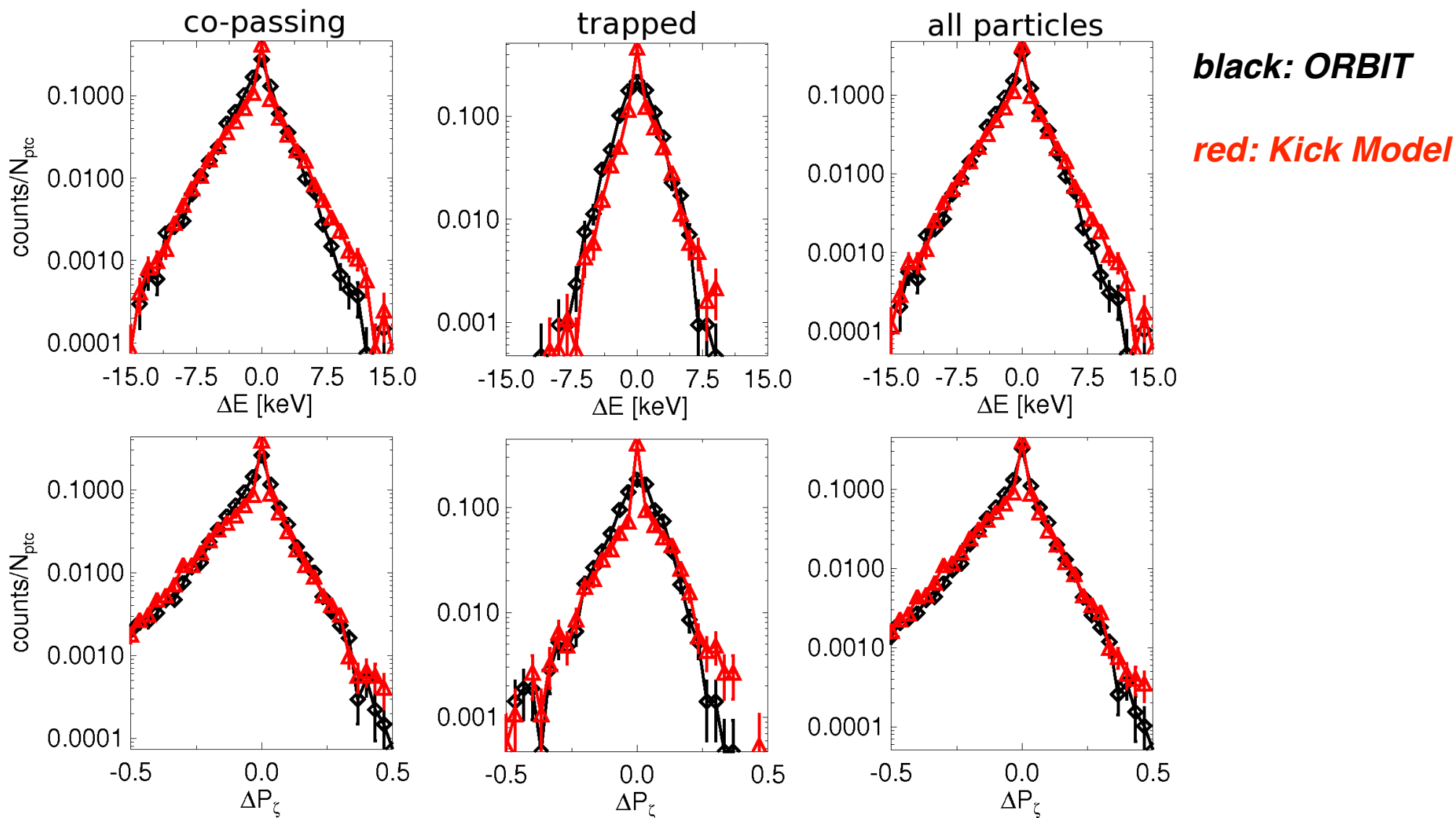
# Change in orbit type observed for some particles: fast ions are kicked around in phase space by TAEs



# Only a few particles (~0.1%) are actually lost for the cases examined here; redistribution dominates



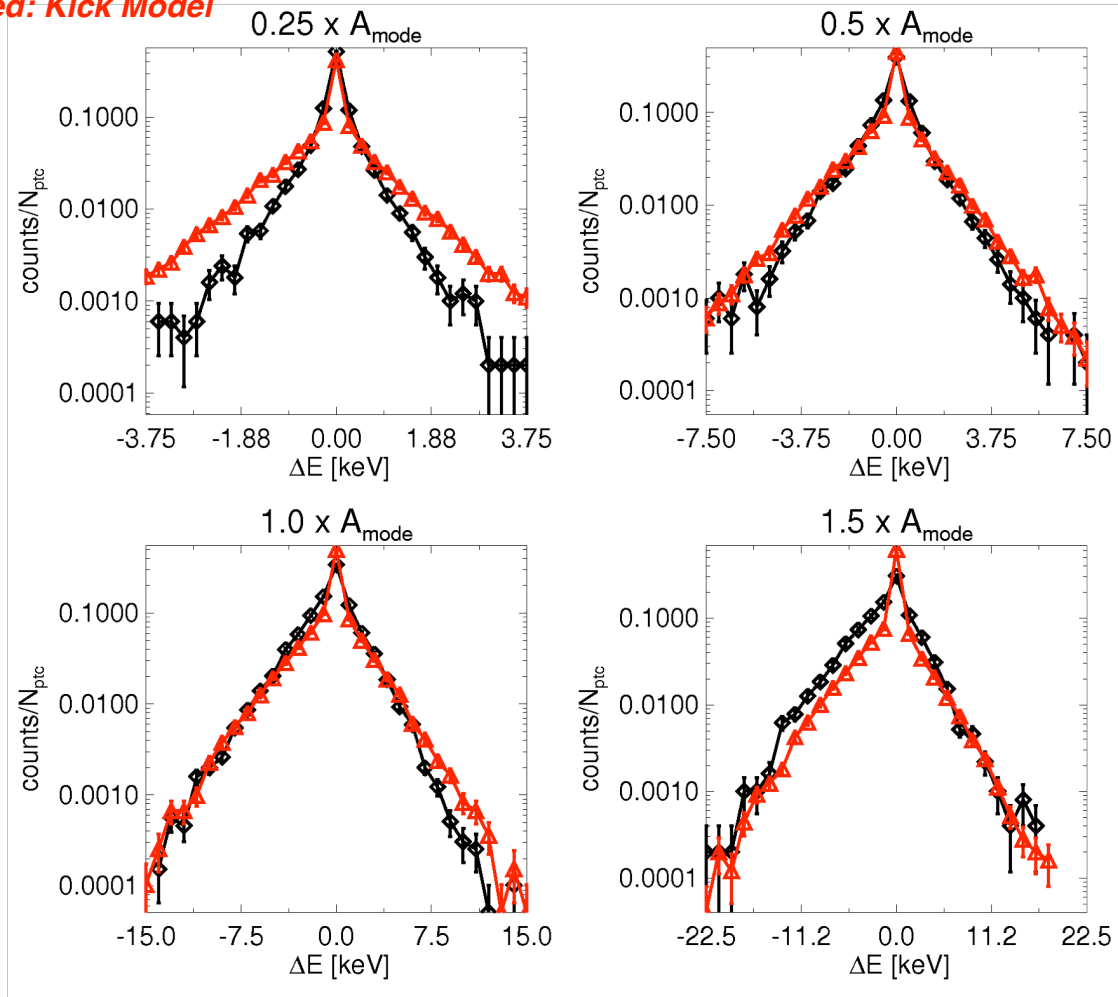
# Good agreement with ORBIT is preserved when evolving $F_{nb}$ over 5 ms, typical macro-step of NUBEAM



Reconstruction is satisfactory for different classes: co-, trapped (, counter- not present in *this* input  $F_{nb}$ )

# Tests assuming different mode amplitudes are satisfactory, though not perfect...

**black:** ORBIT  
**red:** Kick Model

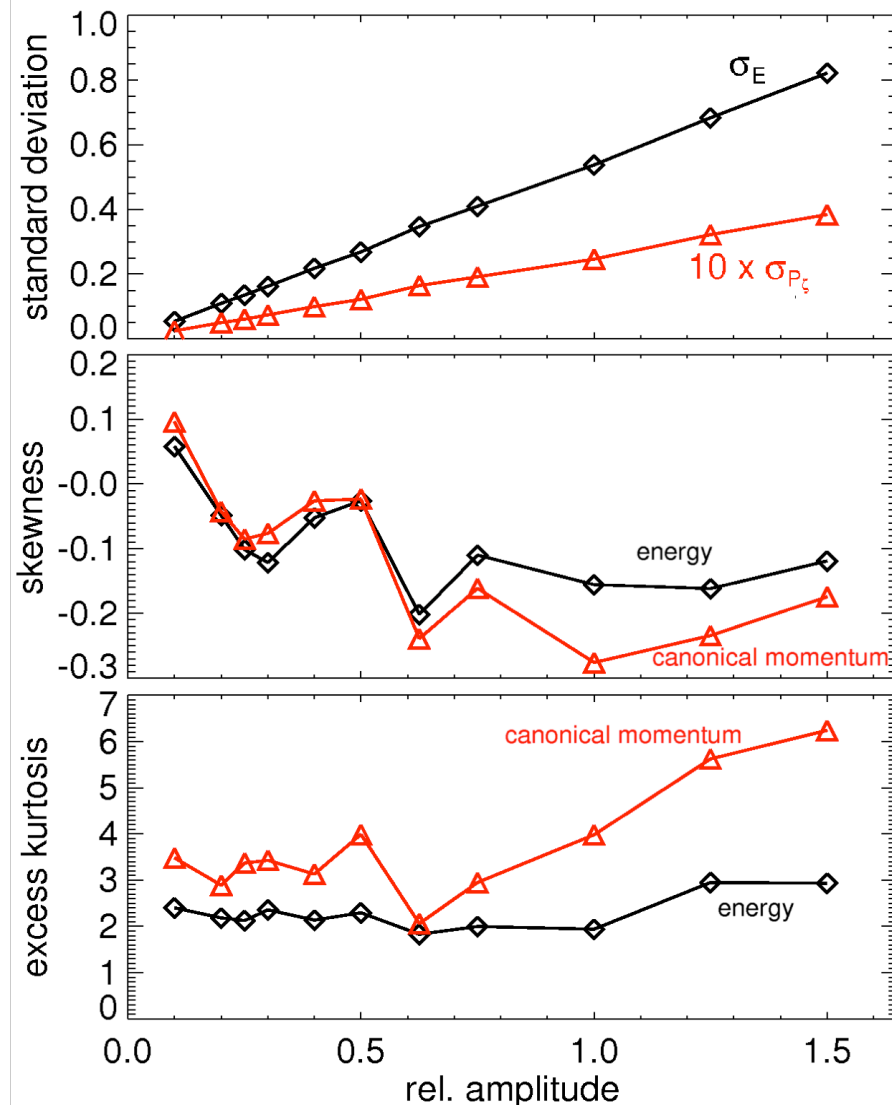
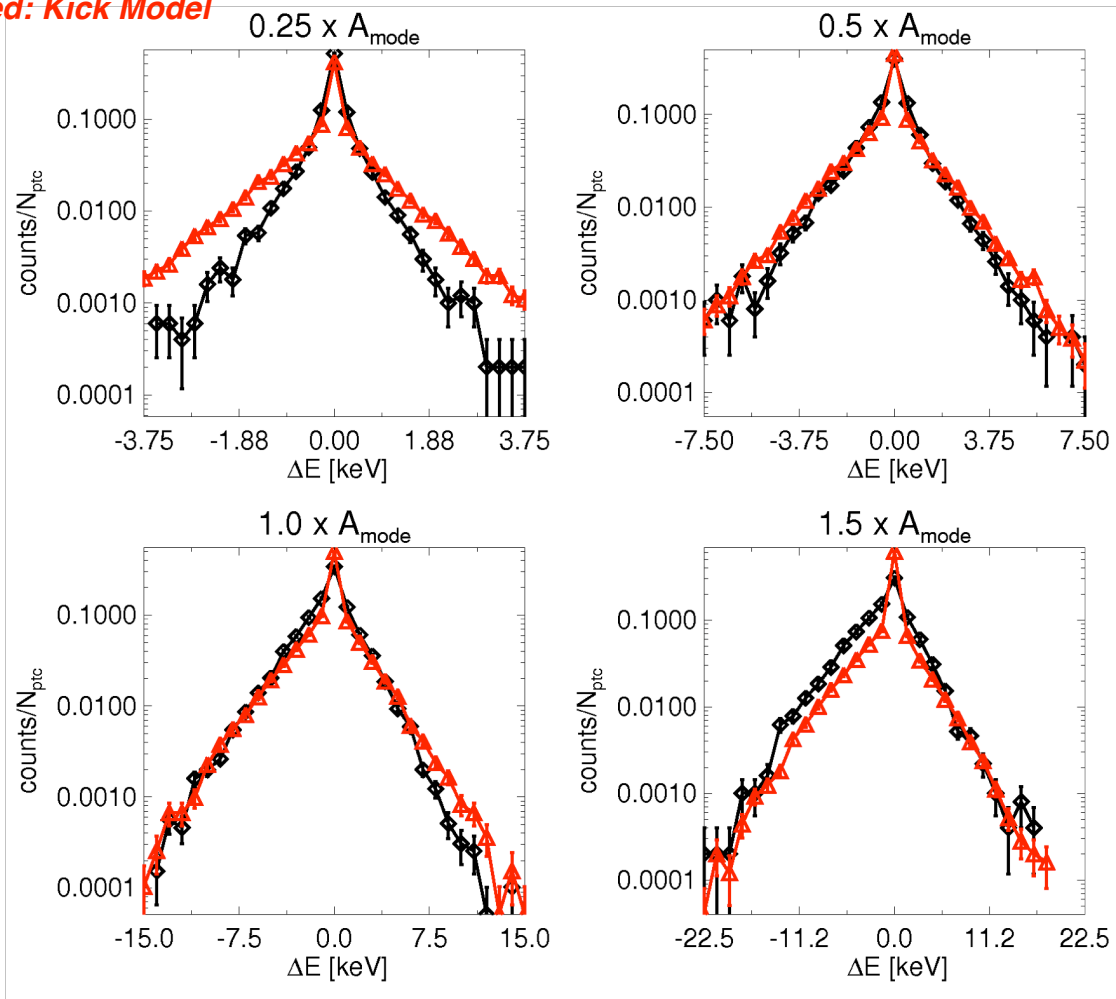


- Reconstruction vs. amplitude is satisfactory

# Tests assuming different mode amplitudes are satisfactory, though not perfect...

black: ORBIT

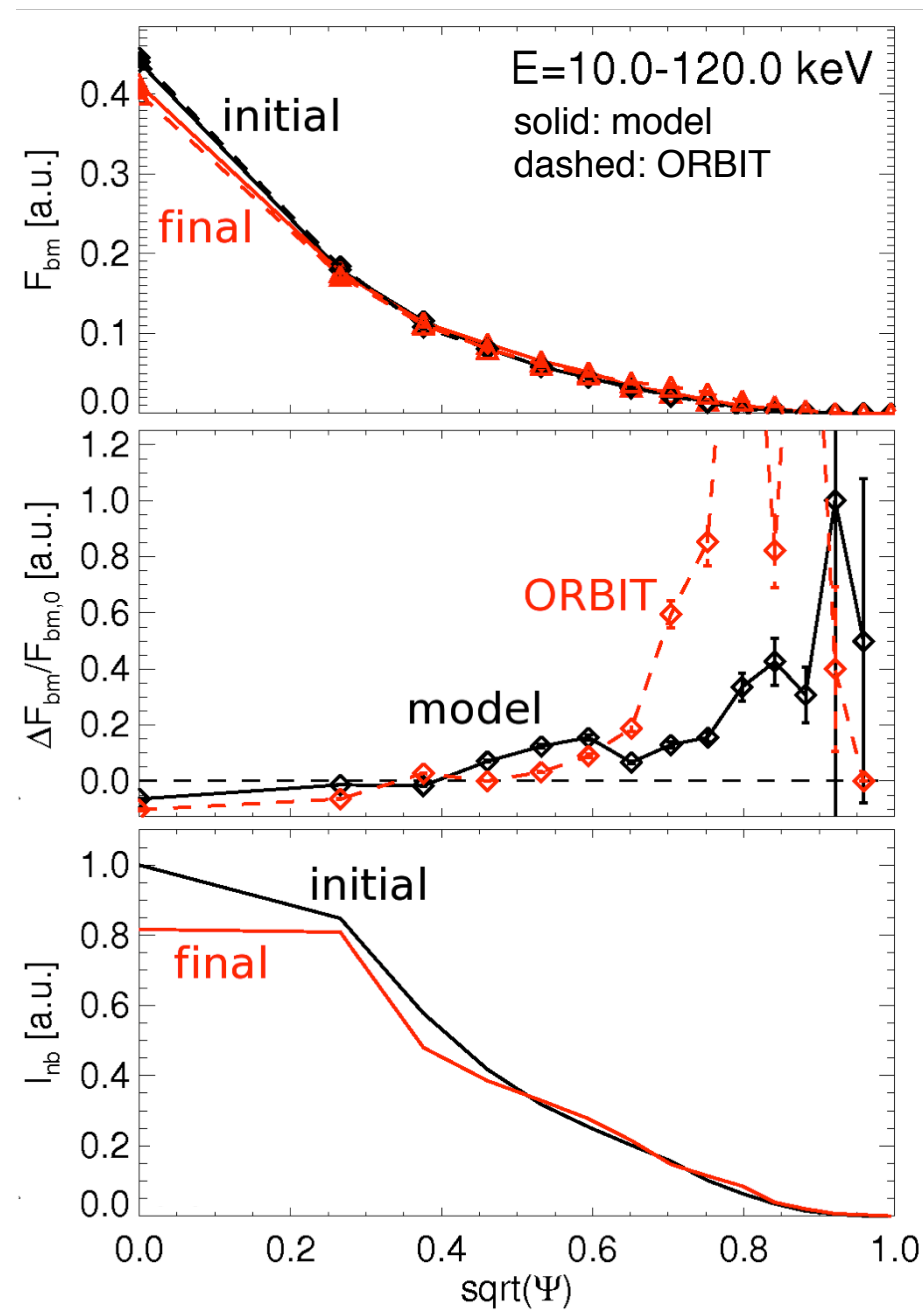
red: Kick Model



- Reconstruction vs. amplitude is satisfactory
- Differences ascribed to simplifications in  $p(\Delta E, \Delta P_\zeta)$  scaling with  $A_{mode}$ : shape *does* change

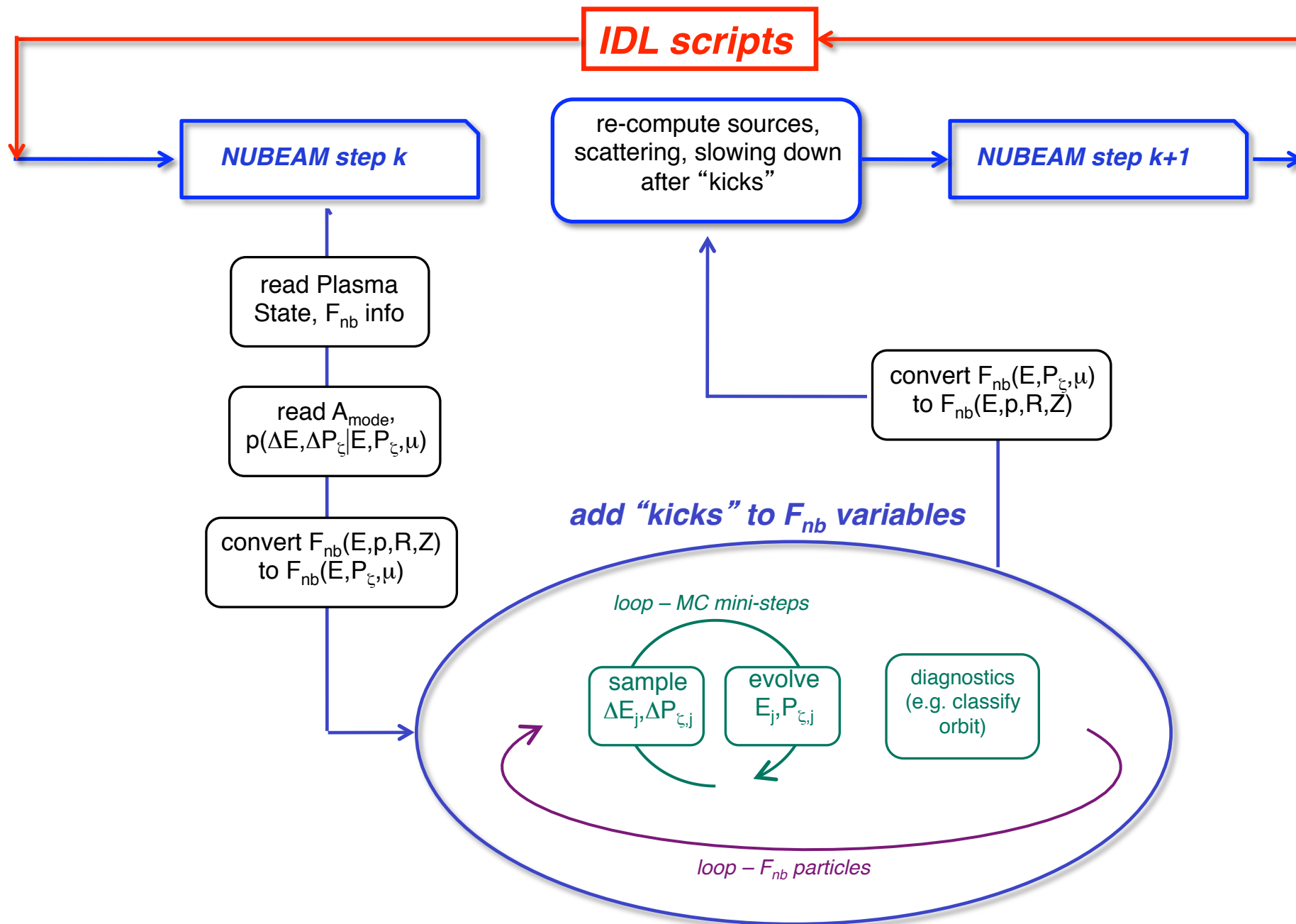


# $F_{nb}$ after 5ms shows fast ion redistribution, only modest losses; NB driven “current” also affected.



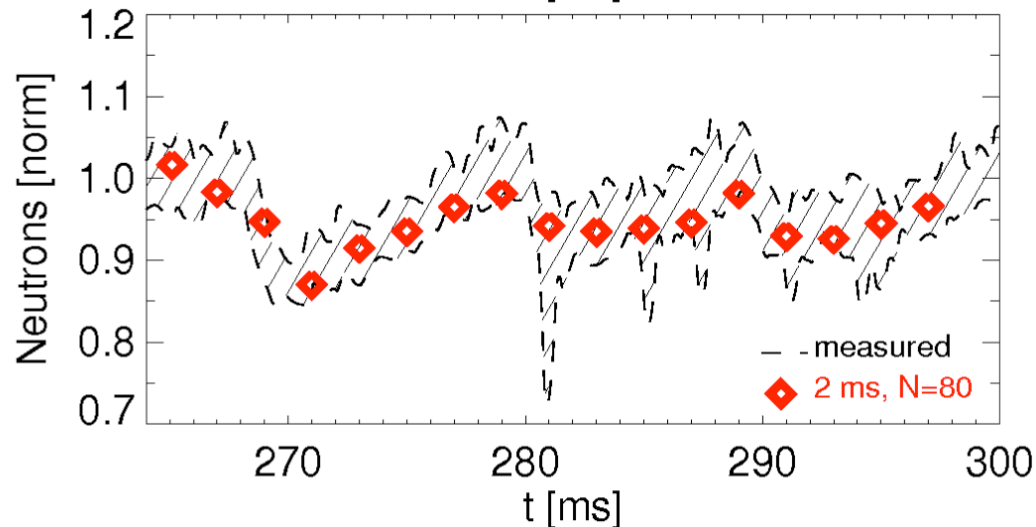
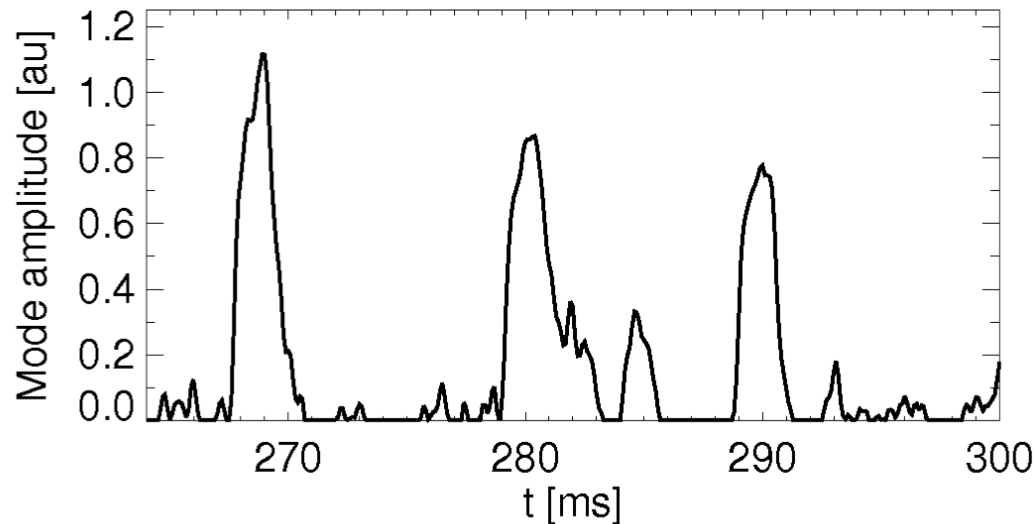
- Compare full ORBIT simulation with reconstructions from reduced model
- $F_{nb}$  drops in the core, fast ions redistributed to larger radii
- Define rough proxy for NB-driven (parallel) current,  $I_{nb} \sim F_{nb} \rho v$ 
  - Larger (relative) variations for  $I_{nb}$  than for  $F_{nb}$  in the core
  - > **Need NUBEAM/TRANSP for more quantitative calculations of  $I_{nb}$**

# Next step: use stand-alone NUBEAM and IDL scripts to simulate $F_{nb}$ evolution for $\gg 5$ ms



# Reduced model reproduces neutron evolution for nominal $A_{mode}(t)$ from neutrons, Mirnovs

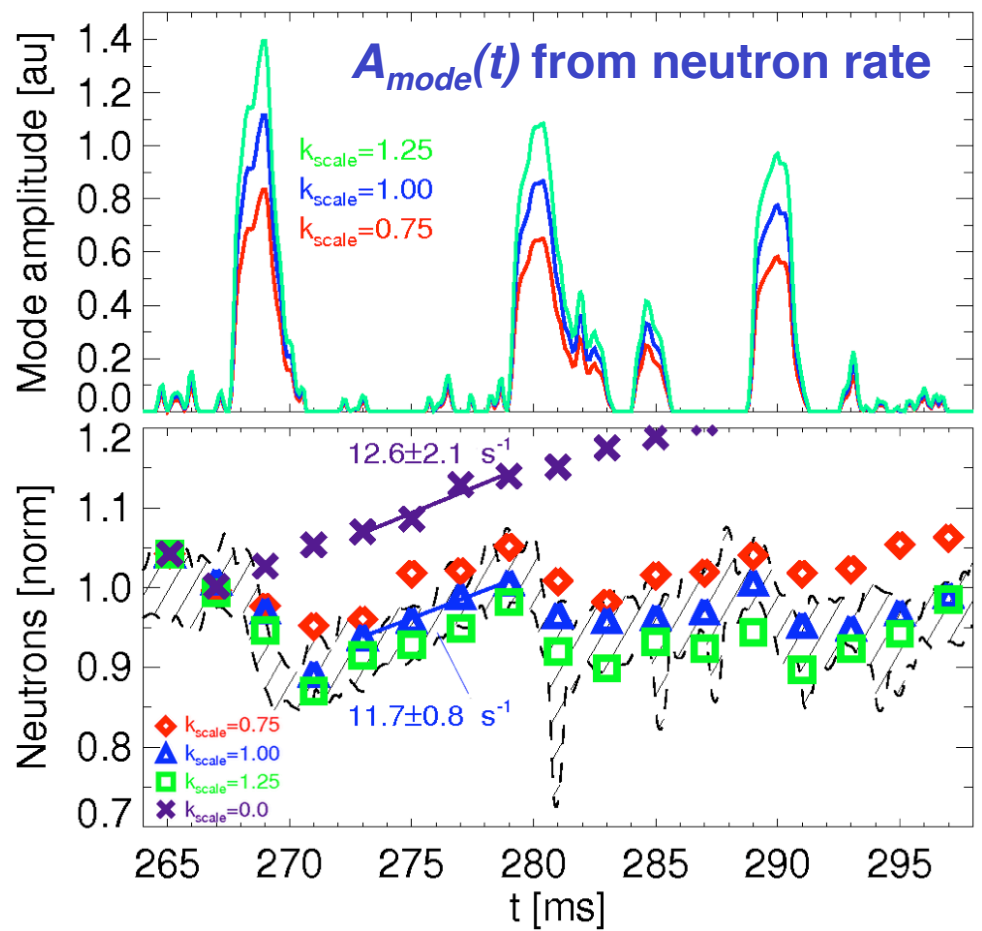
NSTX #139048



- Initial comparison of model predictions w/ experimental neutron rate
- Use stand-alone version of NUBEAM, iterate with reduced model (IDL scripts)
- Background plasma is fixed
  - Normalize exp. neutron rate to central ion density vs. time
  - Normalize all neutron rates at  $t=267$  ms (before first burst)

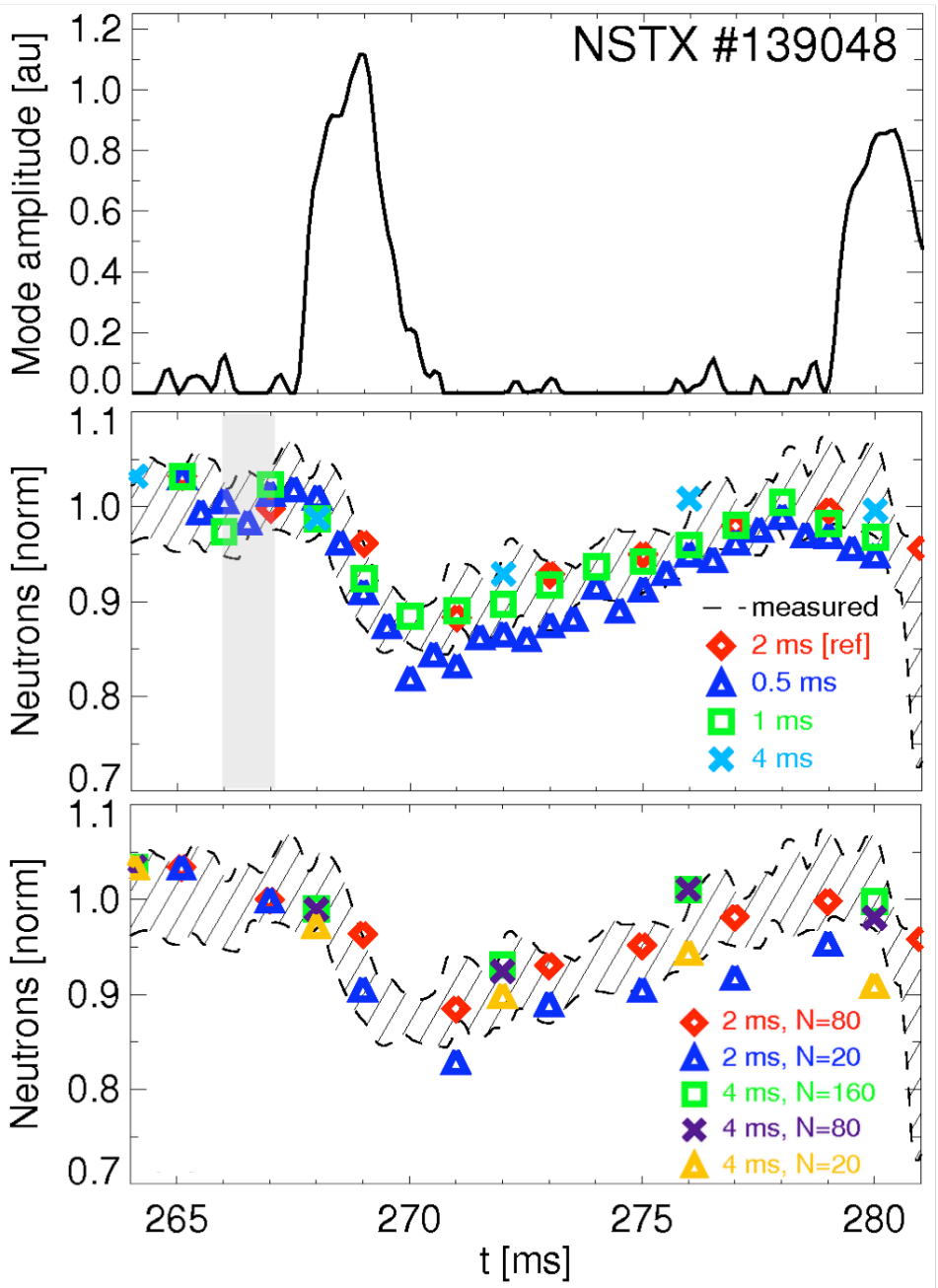
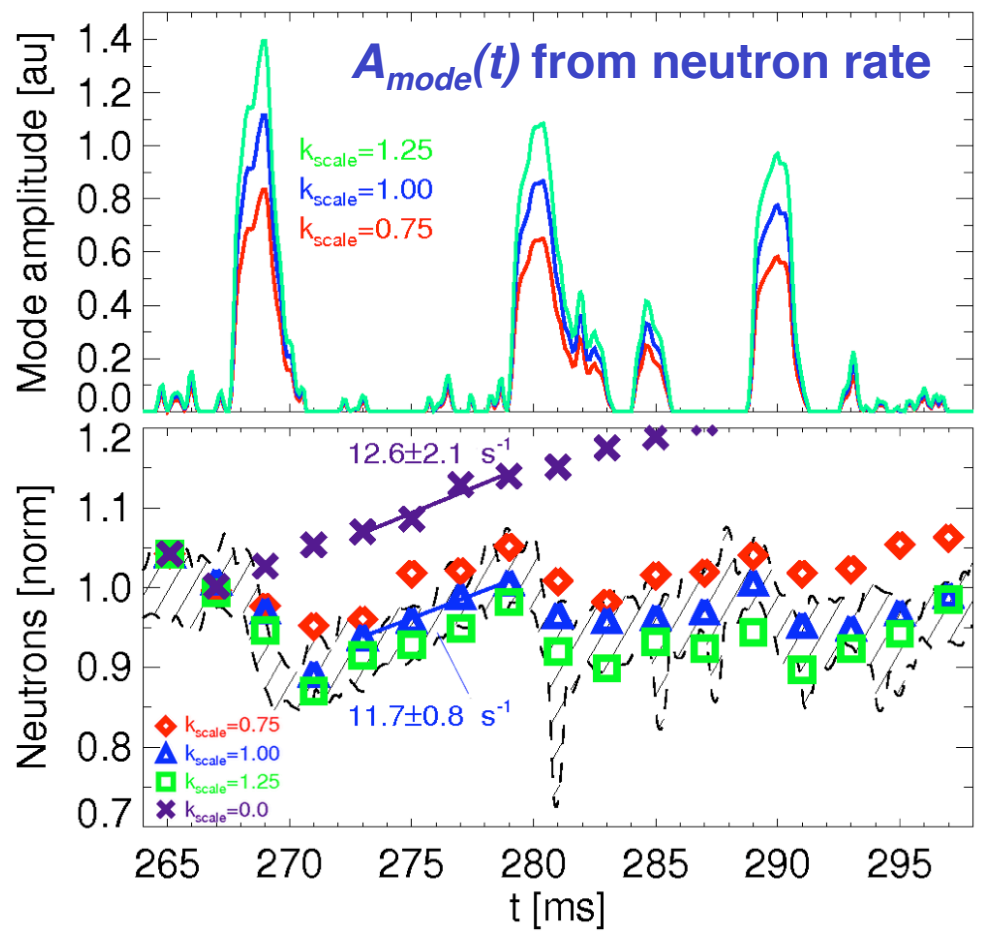
*Satisfactory agreement for  $A_{mode}(t)$  from neutron rate, Mirnovs*

# Results are sensitive to input mode amplitude; time steps must be chosen to satisfy “statistical” approach



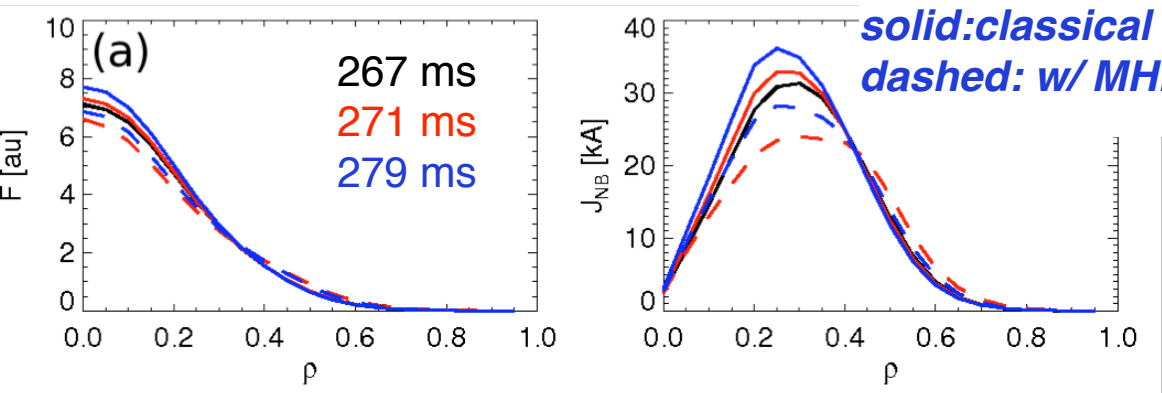
- Typical time scales:
  - Slowing down: 15-30ms
  - Collisions: 2-5ms
  - AEs: 0.1-2ms

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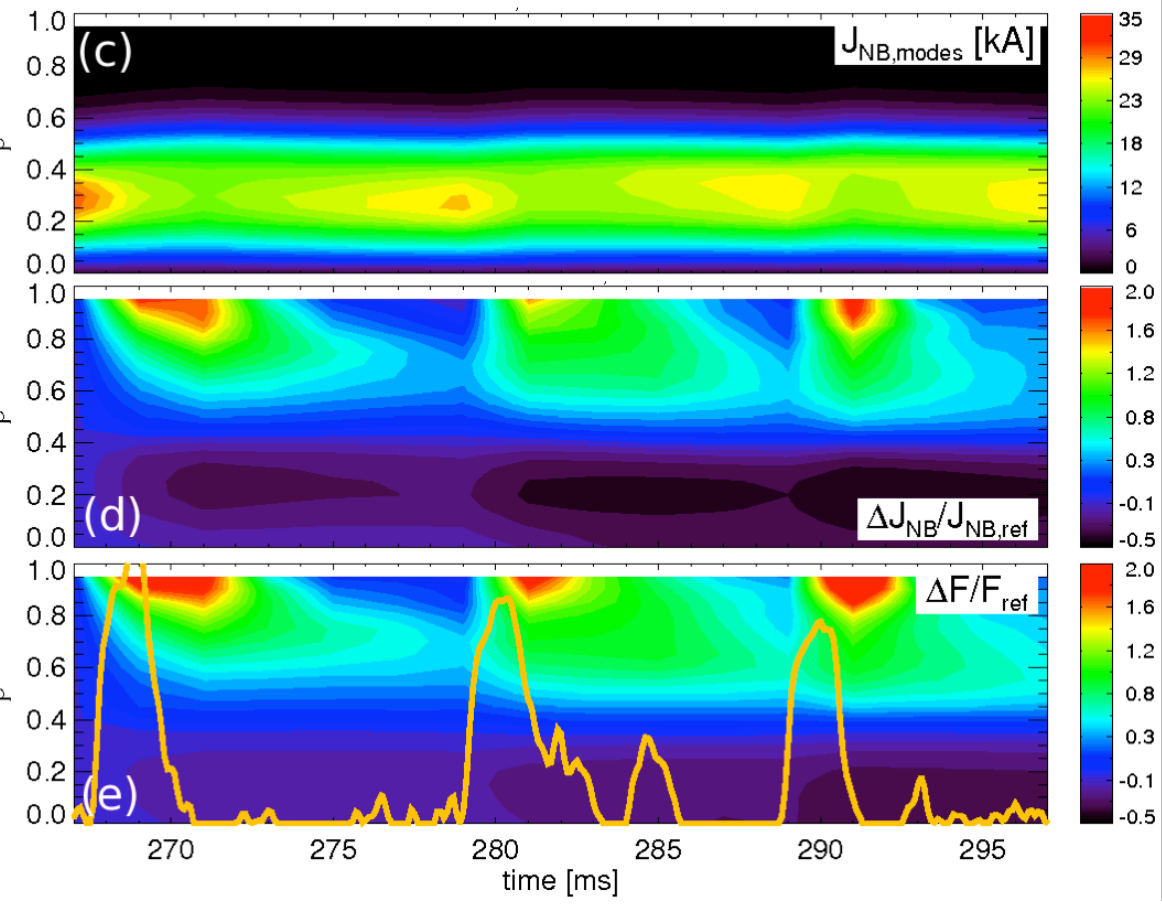


- Typical time scales:
  - Slowing down: 15-30ms
  - Collisions: 2-5ms
  - AEs: 0.1-2ms
- N: number of micro-steps for MC particle evolution, duration  $25\mu\text{s}$

# Reduced model + NUBEAM computes “measurable” fast ion redistribution; NB-driven current strongly affected



- Initial tests show fast ion redistribution induced by each TAE avalanche
- Clear effects on NB-driven current,  $J_{nb}$ 
  - Stronger effect than on  $F_{nb}$
- AE effects persist on slowing down time scales
- Constant NB injection counteracts  $F_{nb}$ ,  $J_{nb}$  depletion

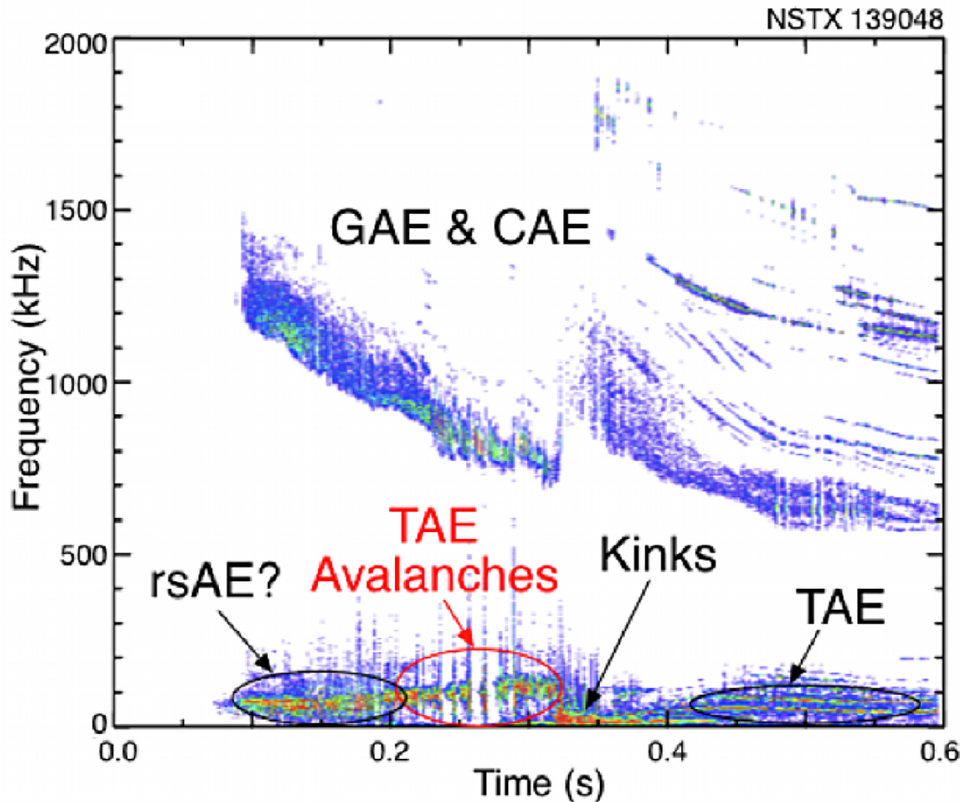


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# The model has been recently improved to account for multiple “classes” of instabilities at a given time

Scenarios with more than one type of modes are quite common:



- This general case is of great practical interest
  - More realistic  $F_{nb}$  evolution when modes have comparable amplitude
  - Can account for different equilibria at different stages of the discharge
  - Can mock-up scenarios such as ‘fast ion channeling’

- Each “class” is modeled by its  $p_k(\Delta E, \Delta P_\xi)$ , weighted by  $A_{mode,k}$
- Instantaneous  $p(\Delta E, \Delta P_\xi) = \sum_k A_{mode,k} \times p_k(\Delta E, \Delta P_\xi)$



# Can the model be used in ‘predictive’ mode?

- As it is, the model is OK to analyze ‘real’ discharges
  - Need mode structure to calculate  $p(\Delta E, \Delta P_{\xi})$ , e.g. from NOVA-K and reflectometers’ data + ORBIT
  - Need data (Mirnovs, neutrons, etc.) to infer  $A_{\text{modes}}(t)$
- Two possibilities for ‘predictive’ runs:
  - Find reasonable guess for unstable modes (e.g. NOVA-K)
  - Explore different scenarios w/ scan of  $A_{\text{modes}}(t)$ : weak \*AE activity, bursts/ avalanches, etc.

or:

- Have an additional module to compute  $A_{\text{modes}}(t)$  self-consistently?
- Still requires mode structure, probably estimates for  $\gamma_{\text{drive}}$ ,  $\gamma_{\text{damp}}$
- Let  $A_{\text{modes}}(t)$  evolve according to  $F_{\text{nb}}$  evolution – but how?
- *Couple to other reduced models, e.g. 1.5D Quasi-Linear?*

# Summary

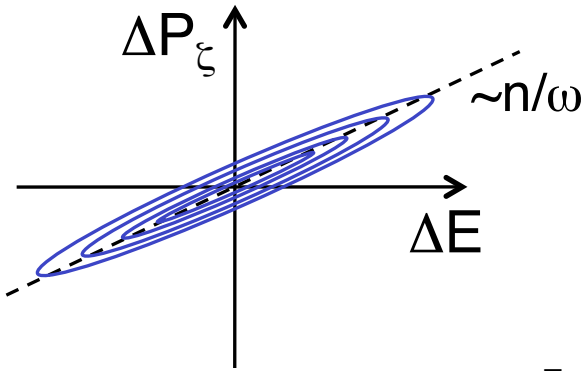
- Test algorithm for reduced fast ion transport model developed, being verified
  - Results compared with full ORBIT runs
  - Confirmed validity of approach for practical case
- Successful preliminary tests for NSTX case with TAE modes (*avalanches*)
  - Shot#139048,  $t \sim 265\text{-}300$  ms: H-mode avalanches
  - Strong redistribution of fast ions observed
  - Modest losses, consistent with previous detailed modeling
  - NB-driven current  $J_{nb}$  is affected, too
- > **Implementation in NUBEAM under way**
  - > Extensive Verification&Validation planned (multi-machine)
  - > Identify issues, possible improvements to the model

# Backup

# For low-frequency \*AEs with $\omega \ll \omega_{ci}$ such as TAEs, magnetic moment $\mu$ is conserved (...but maybe it's not)

- *In this presentation, it is assumed that  $\Delta\mu=0$*
- However:  $\Delta\mu=0$  hypothesis can break down if
  - $\rho_f \sim$  radial width of the modes
  - $\rho_f \sim$  scale-length of equilibrium profiles
  - ⇒ *Both conditions are likely to be met in spherical tokamaks (e.g. NSTX)*
- ⇒ **Proposed model can be generalized to cases where  $\mu$  is *not* conserved**
  - Also important for  $\omega_{ci}$ -range instabilities: GAE/CAEs

# Example for single-resonance case: analytical *probability distribution function*



For each *bin* in  $(E, P_\zeta, \mu)$ , steps in  $\Delta E, \Delta P_\zeta$  can be approximated by a bivariate  $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu, A)$

$$p = p_0 e^{-\frac{1}{2(1-\rho)} \left[ \frac{(\Delta E - \Delta E_0)^2}{\sigma_E^2} + \frac{(\Delta P_\zeta - \Delta P_{\zeta 0})^2}{\sigma_{P_\zeta}^2} - 2\rho \frac{(\Delta E - \Delta E_0)(\Delta P_\zeta - \Delta P_{\zeta 0})}{\sigma_E \sigma_{P_\zeta}} \right]}$$

$$p_0 = \frac{1}{2\pi \sigma_E \sigma_{P_\zeta} \sqrt{1 - \rho^2}} \quad \text{normalization}$$

$$\rho = \frac{\langle (\Delta E - \Delta E_0)(\Delta P_\zeta - \Delta P_{\zeta 0}) \rangle}{\sigma_E \sigma_{P_\zeta}} \quad \text{correlation parameter,}$$

*such that*  $\Delta P_\zeta(\Delta E) = \Delta P_{\zeta 0} + \text{sign}(\rho) \times \frac{\sigma_{P_\zeta}}{\sigma_E} (\Delta E - \Delta E_0)$

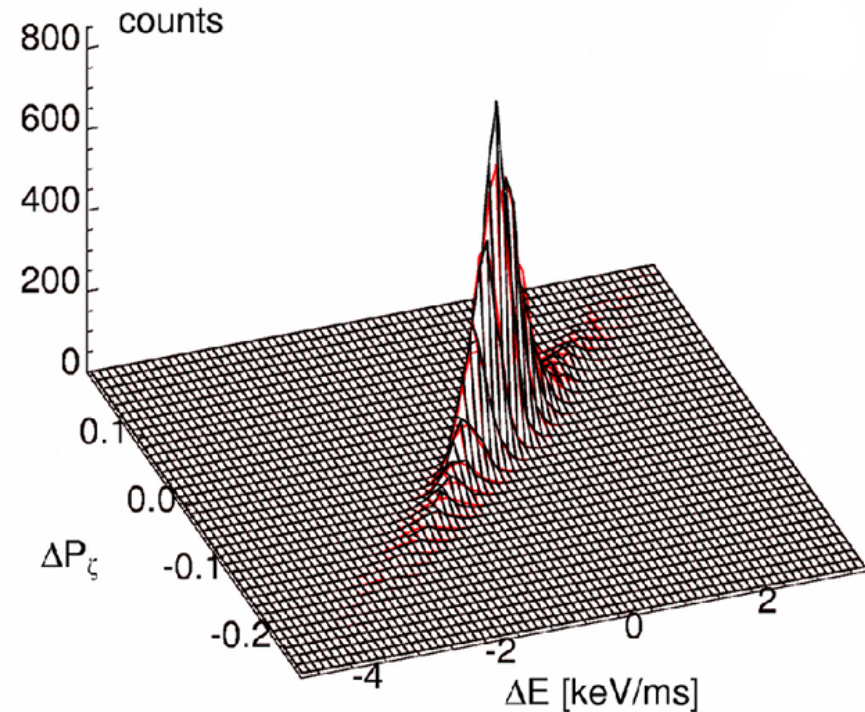
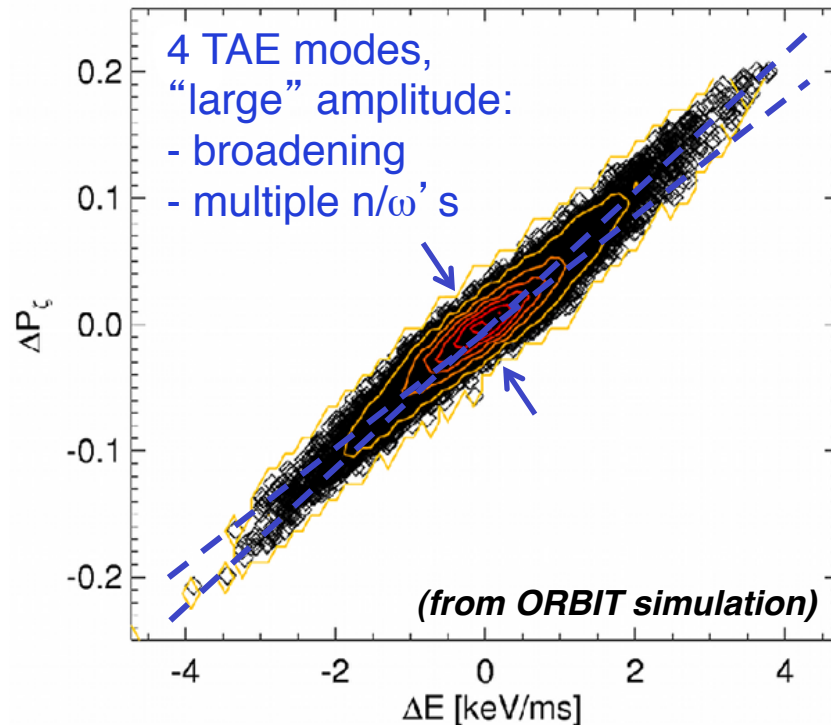
...and all parameters depend, in principle, on the mode amplitude  $A=A(t)$

# Single (isolated) resonances introduce fundamental constraints on particle's trajectory in $(E, P_\zeta, \mu)$

- From Hamiltonian formulation:

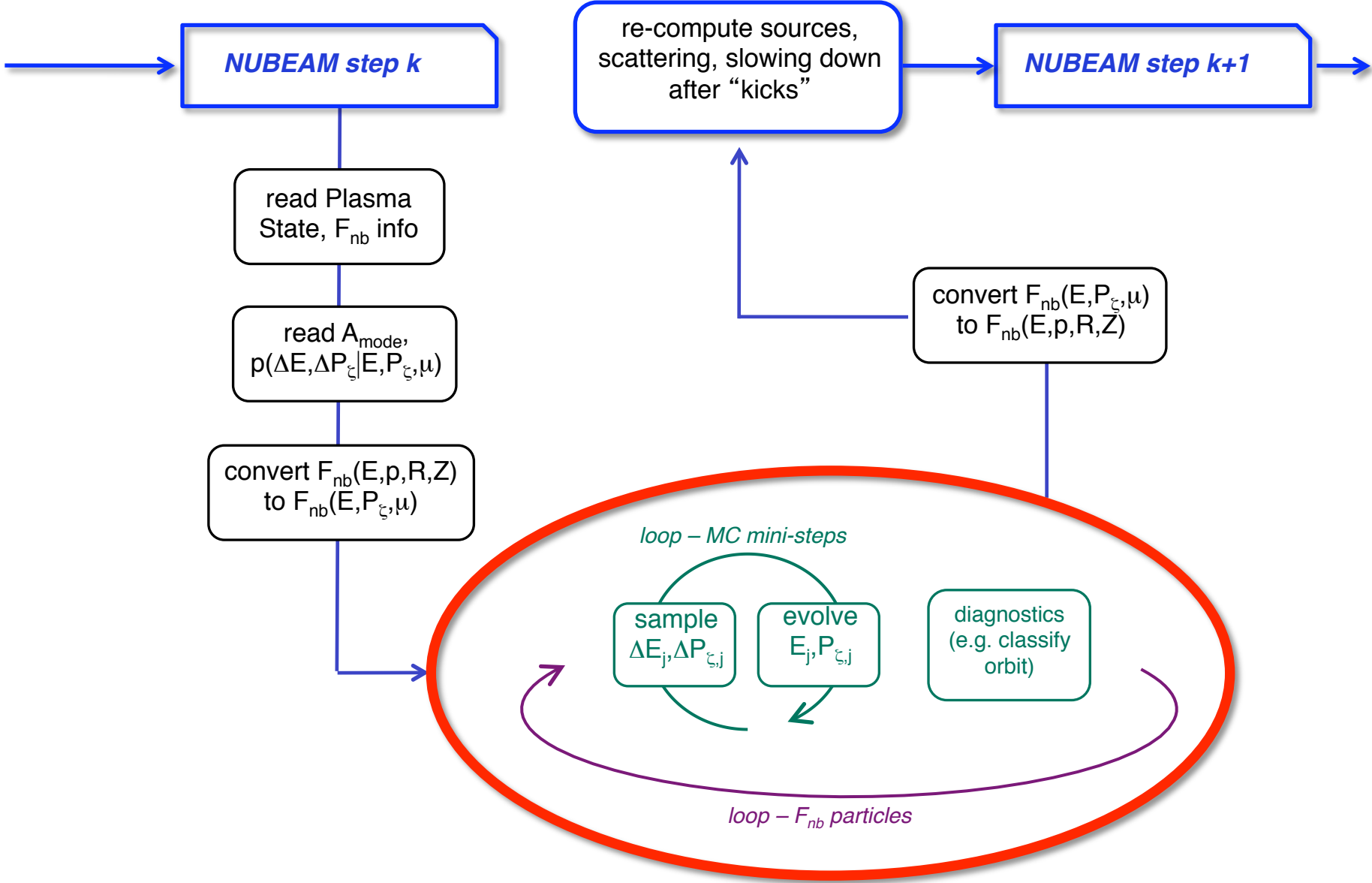
$$\omega P_\zeta - nE = \text{const.} \implies \Delta P_\zeta / \Delta E = n / \omega$$

$\omega = 2\pi f$ , mode frequency       $n$ , toroidal mode number



*Presence of multiple modes/resonances distorts the 'ideal' (linear) relationship*

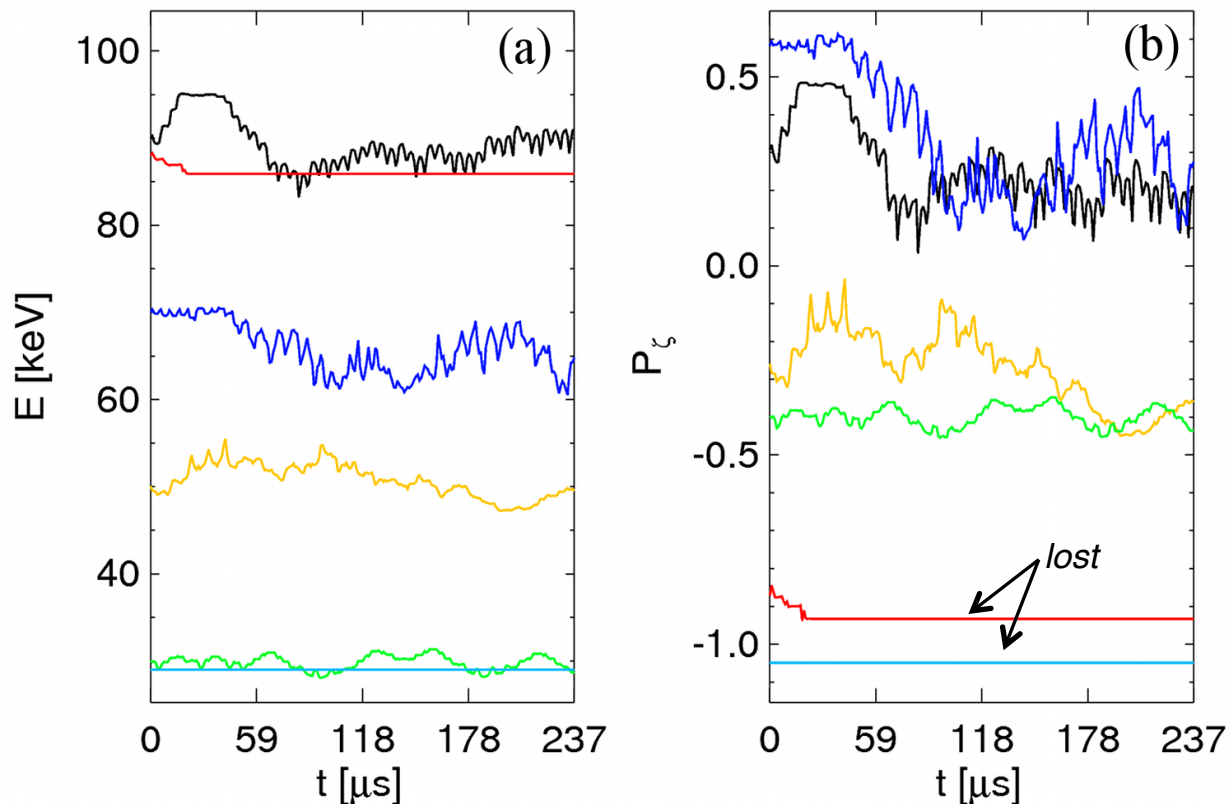
# Scheme to advance fast ion variables according to transport probability in NUBEAM module of TRANSP



**add “kicks” to  $F_{nb}$  variables**

# Scheme to evolve $F_{nb}$ takes into account (in a semi-empirical way) constraints of resonant interaction

- Particle's motion is characterized by different time-scales:
  - Fast oscillation in wave field – *neglected*
  - 'Jumps'  $\Delta E, \Delta P_\zeta$  around instantaneous energy,  $P_\zeta$
  - Slow (secular) drift from initial energy,  $P_\zeta$





# Putting all together

At each ‘macroscopic’ NUBEAM step:

- I. Re-normalize bins  $(P_\xi, E, \mu)$  based on q-profile, fields, ...
- II. Identify ‘bin’ in  $(P_\xi, E, \mu)$  for current ‘particle’ (i.e. orbit)
- III. Extract steps  $\sigma_E, \sigma_{P_\xi}$  ( $\sigma_\mu$ ) from multivariate  $p(\Delta E, \Delta P_\xi, \Delta \mu)$
- IV. Compute sign  $S_{r,i}$  from  $p(\Delta E, \Delta P_\xi)$ : positive or negative kicks
- V. Rescale steps based on  $A_{modes}(t)$

Loop over particles

VI. Advance  $E, P_\xi$  ( $\mu$ ):

$$\begin{cases} \overline{\Delta E}_i = S_{r,i} \times A_{mode}(t = \bar{t}) \times \sigma_{E,i} \\ \implies E_i = E_i + \overline{\Delta E}_i \end{cases}$$

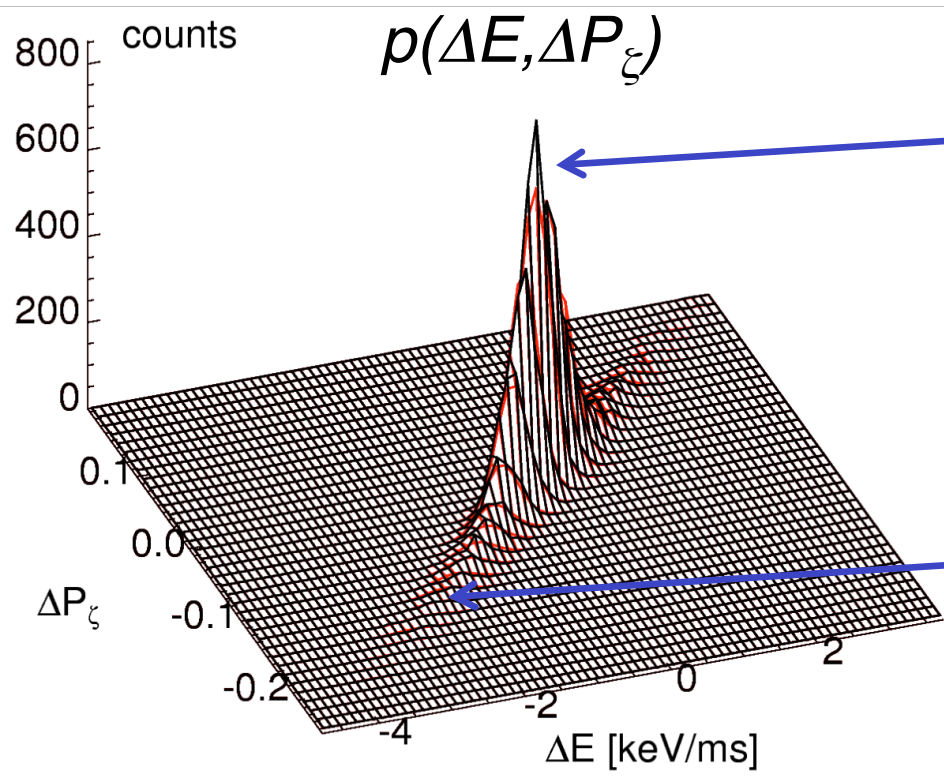
- VII. Advance particle’s trajectory in phase space for next step
- VIII. Compute slowing down, scattering (NUBEAM)

*where steps II-VII are divided in sub-steps for each particle*

**Required input, e.g. through ‘Ufiles’:**

**scaling factor (“mode amplitude”)  $A_{mode}$ , probability  $p(\Delta E, \Delta P_\xi)$**

# Discrete bins in $(P_\xi, E, \mu)$ can contain both *resonant* and *non-resonant* particles



- ‘Non-resonant’ particles have small fluctuations around initial  $(E, P_\xi)$
- ‘Resonant’ particles can experience large  $\Delta E, \Delta P_\xi$  variations

- To keep track of particle’s class:
  - Sample steps  $\sigma_E, \sigma_{P_\xi}$  at first step only
  - This mimic the “correlated random walk” experienced by the particles
  - Exception: particle move to a different bin -> re-sample

$p(\Delta E, \Delta P_\xi | P_\xi, E, \mu)$  can be skewed to positive/negative  $\Delta E, \Delta P_\xi$ , causing overall “drift” of  $F_{nb}(P_\xi, E, \mu)$

- Introduce ‘random sign’ for  $i$ -th step in MC procedure,  $S_{r,i}$
- For each particle (e.g. pair of correlated steps  $\sigma_E, \sigma_{P_\xi}$ ), calculate  $S_{r,i}$  from probability of positive vs. negative steps

- From  $p(\Delta E, \Delta P_\xi)$  compute

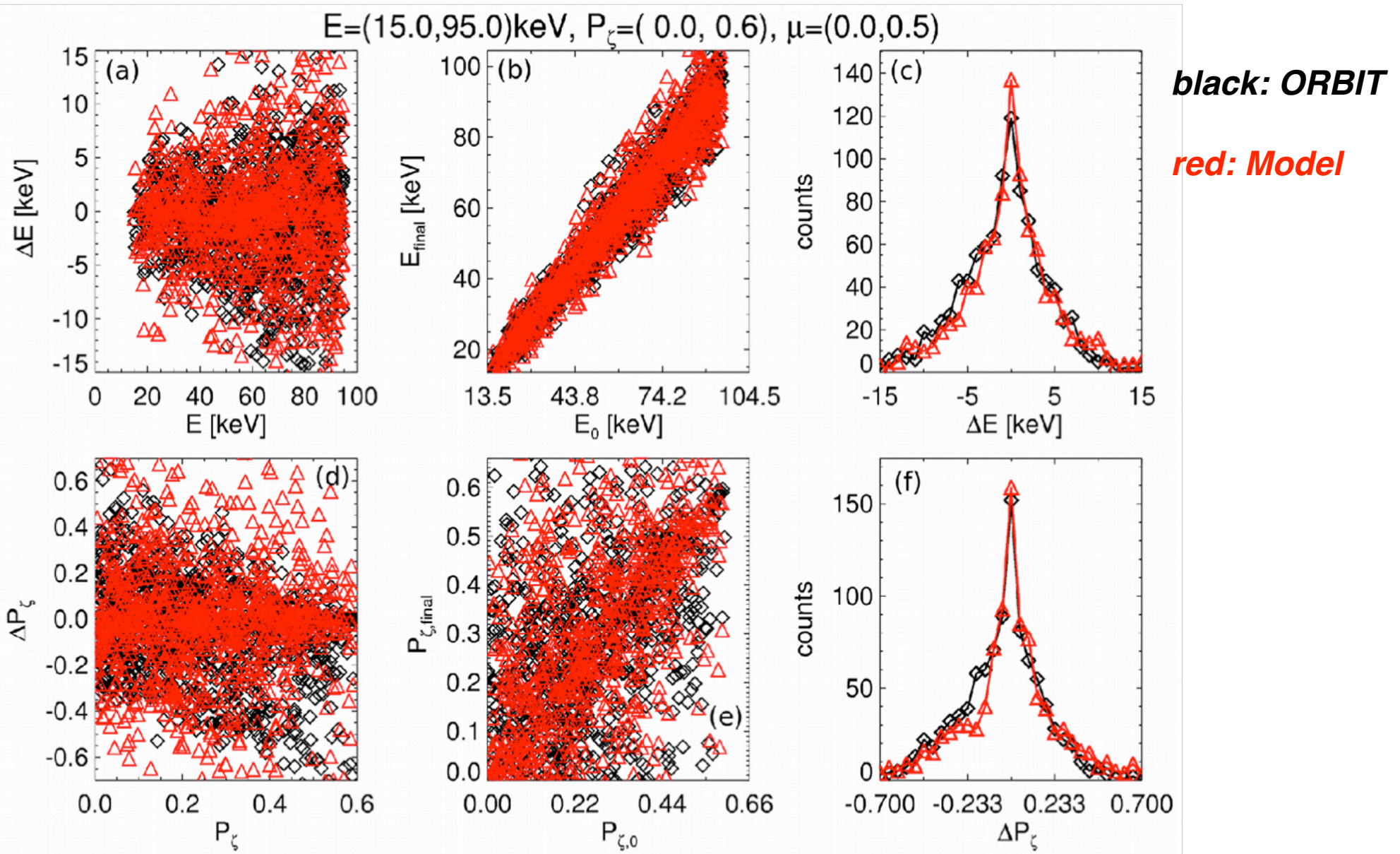
$$p_+ \doteq p(\sigma_{E,i}, \sigma_{P_\xi,i}) \quad ; \quad p_- \doteq p(-\sigma_{E,i}, -\sigma_{P_\xi,i})$$

- Then define  $f_{\text{sign}}$ :

$$f_{\text{sign}} = \frac{p_+}{p_+ + p_-}$$

- Finally, use  $0 < f_{\text{sign}} < 1$  to bias random extraction of  $S_{r,i} = +1, -1$

# Example: evolving $F_{nb}$ over 270 $\mu\text{s}$ in 5 sub-steps

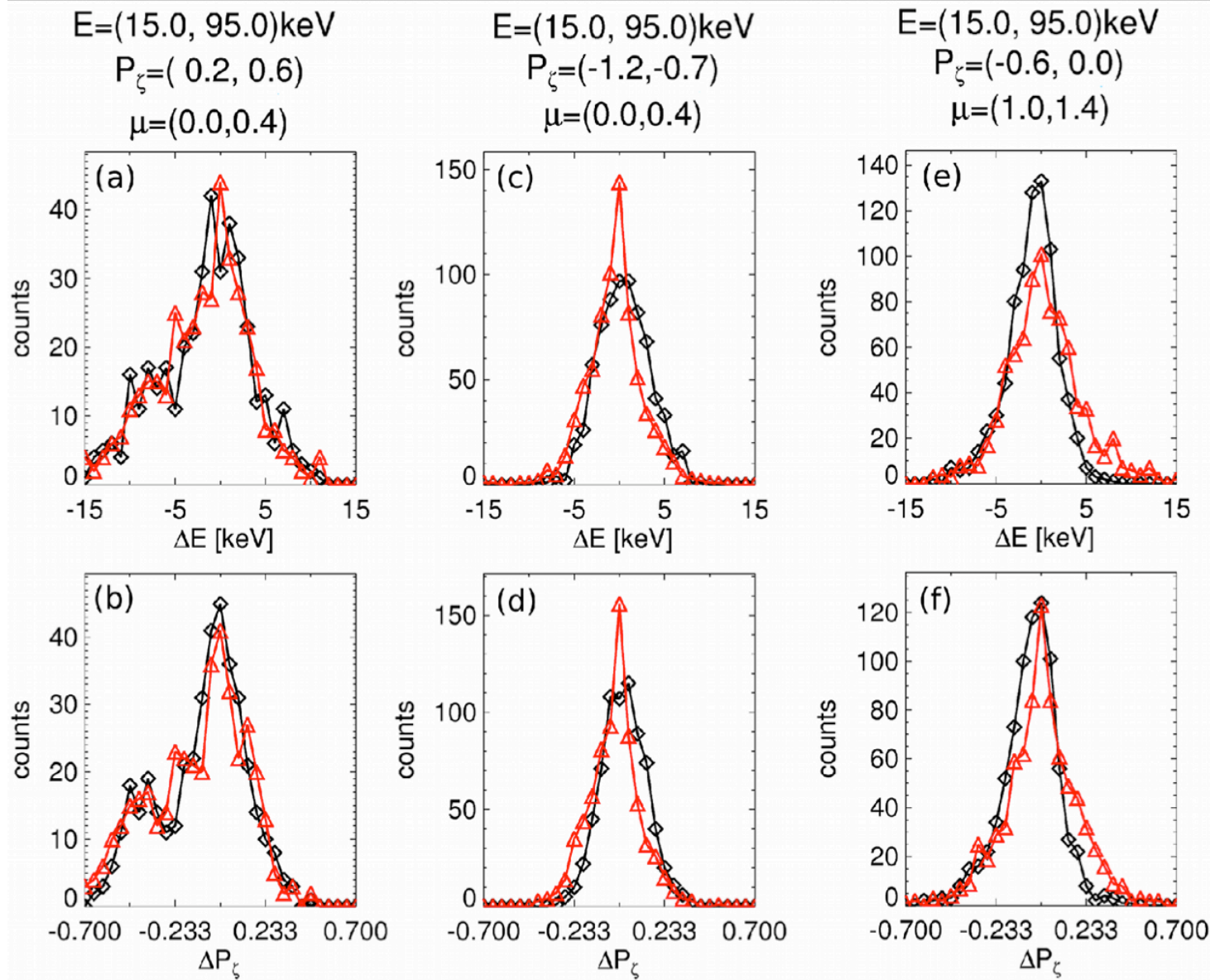


# Reconstruction works for different classes: co-, counter-, trapped

*co-passing*

*counter-passing*

*trapped*



**black: ORBIT**

**red: Model**

# Reconstruction works at different sub-steps

**black: ORBIT**    **red: Model**

