
Coupled Ray-Tracing and Fokker-Planck EBW Modeling for Spherical Tokamaks

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A short intro on EBW

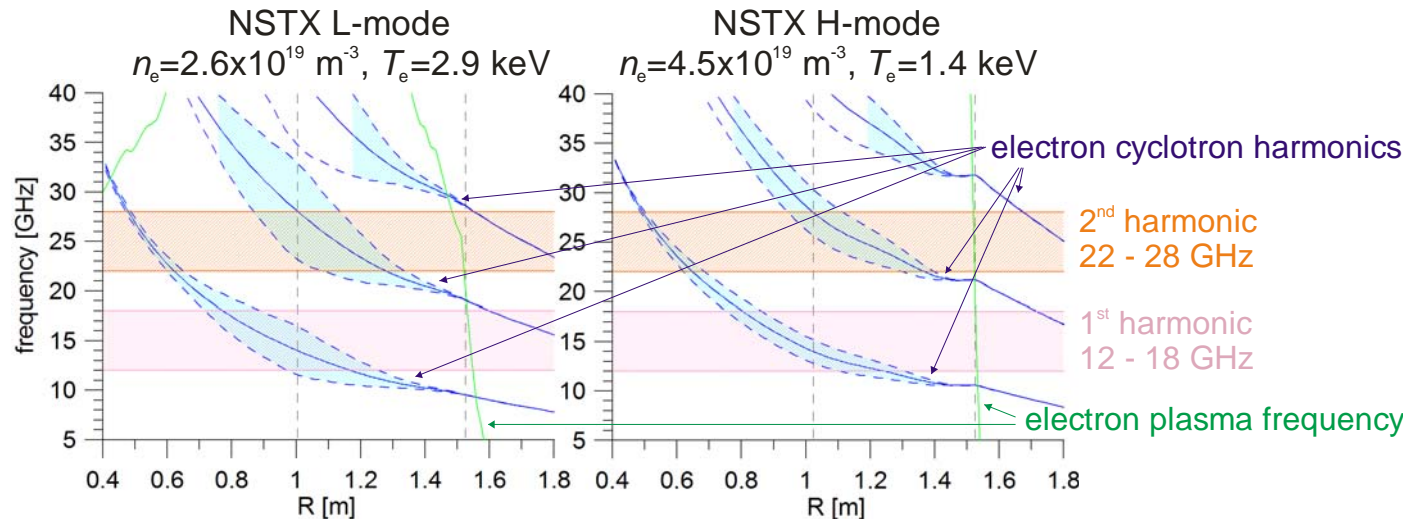
- **Electron cyclotron (EC) heating** and **current drive** is important for magnetic confinement fusion
 - “Standard” **EC waves are cut off** in spherical tokamaks (STs): $\omega_{pe} \gg \Omega_{ce}$
 - **Electron Bernstein wave (EBW)** must be used in STs
 - no upper density cut-offs
 - efficiently absorbed at any harmonics
 - efficient current drive because of interaction with suprathermals
 - must be excited by O- or X-modes
 - more difficult to control
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The EBW simulation

- **Two coupled codes**
 - **AMR** (Antenna—Mode-conversion—Ray-tracing)
 - **LUKE** (3D Fokker-Planck code)
 - (similar to GENRAY + CQL3D)
 - **Antenna aiming determined** by the O-X-EBW optimum
 - **EBW propagation** inside the plasma well described by **ray tracing**
 - **Quasi-linear damping and current drive** calculated by **LUKE**
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Injection and target plasma parameters

- **Only two EBW launching parameters** can be chosen **arbitrarily**
 - **frequency**
 - **vertical antenna position**
 - (two O-X-EBW optimums exists in opposite toroidal directions)
- Calculations carried out
 - 1 MW power
 - typical NSTX L-mode (and H-mode)
 - 1st and 2nd harmonic



Typical EBW behavior

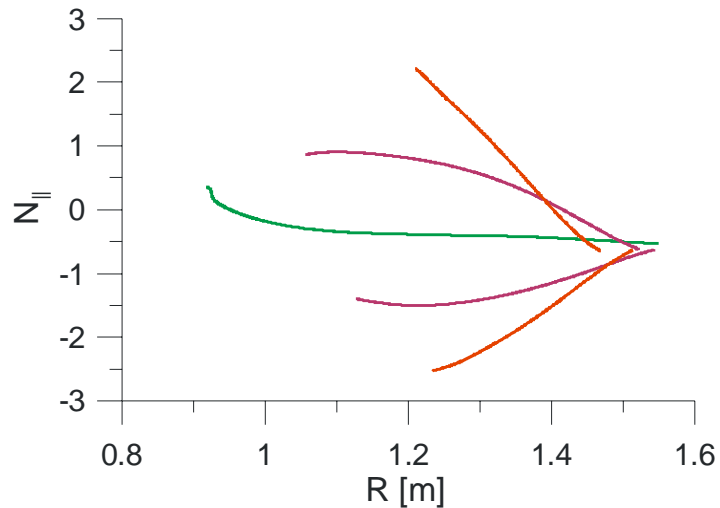
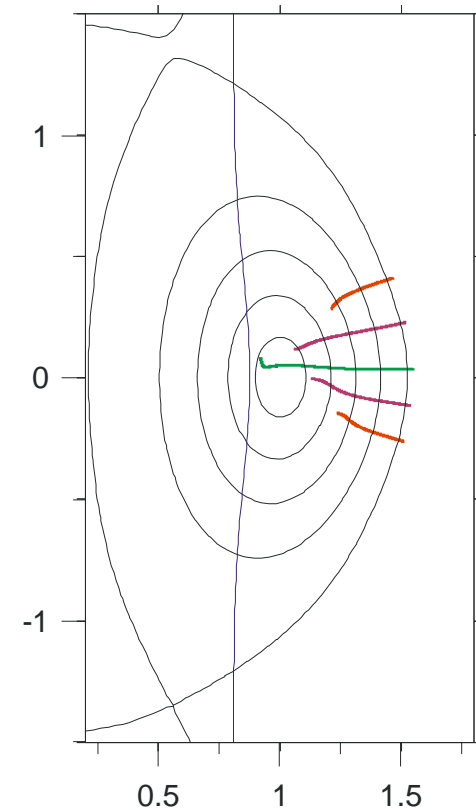
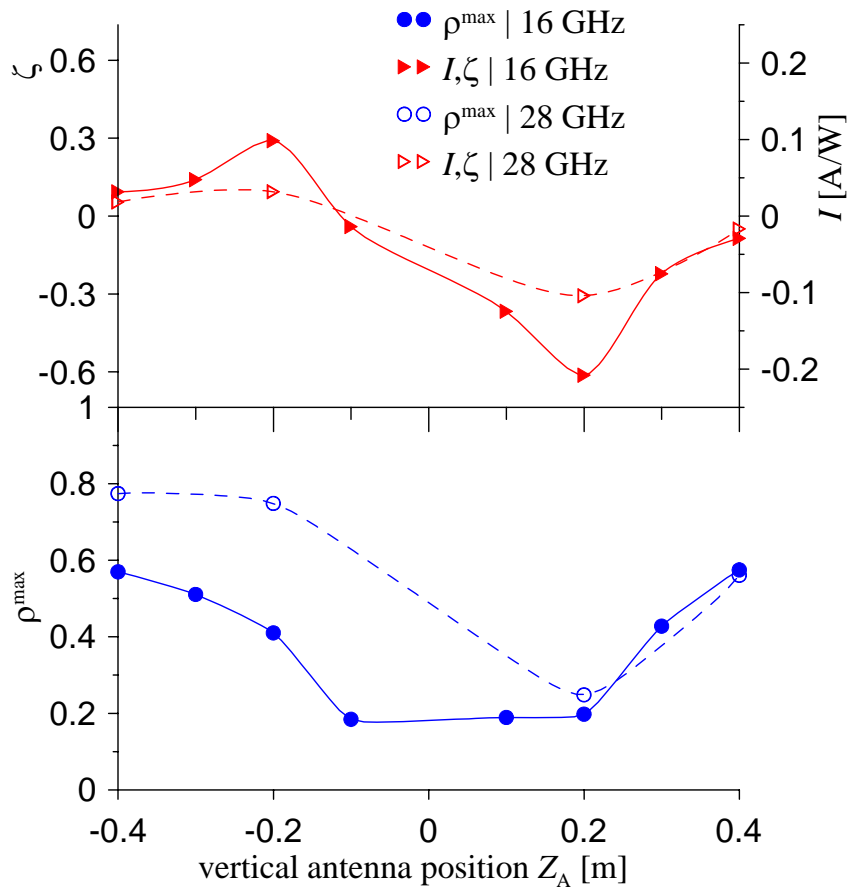


Fig. 2. Ray N_{\parallel} (top) and trajectories (right) of 16 GHz EBWS in L-mode plasma for $Z_A=0, \pm 0.2, \pm 0.4$ m.



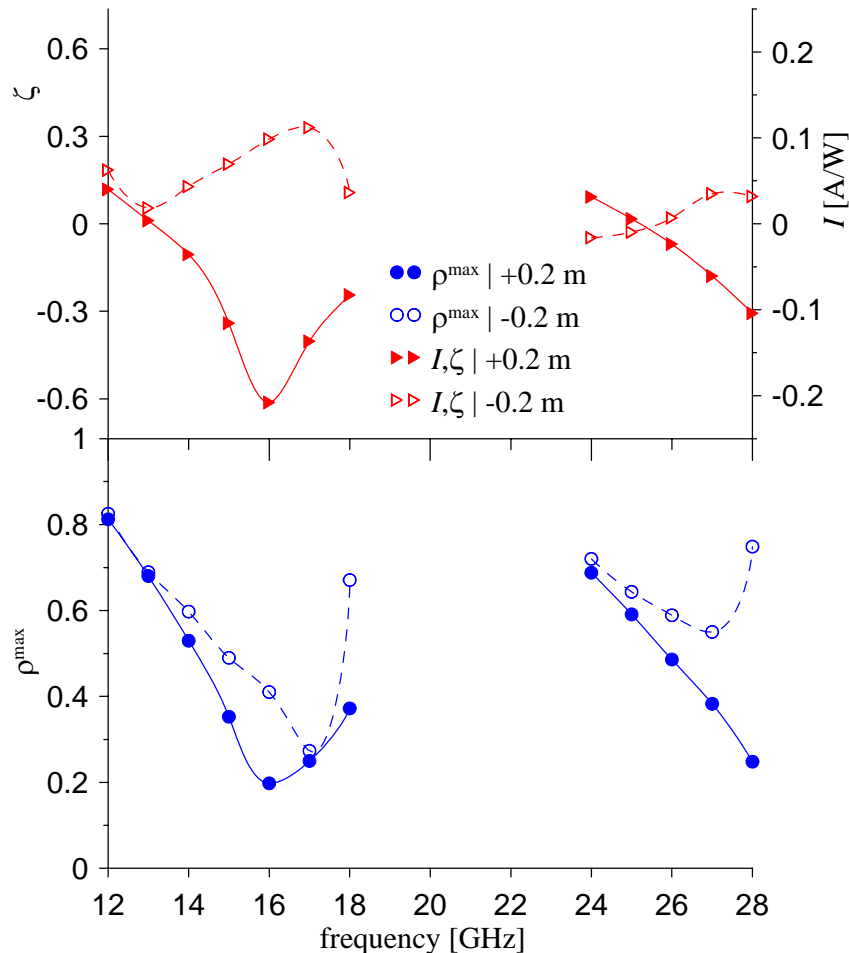
Mid-plane launch can be favorable for **central heating**, but is very **unfavorable for current drive** because of the small (and oscillating) N_{\parallel} .

Antenna vertical position scan



- **The current direction is determined**
 - by the vertical launch position, which determines the $N_{||}$
 - by the damping location, which determines the current drive mechanism: Fisch-Boozer at smaller radii, Ohkawa at the edge

Frequency scan



- Frequency dependence qualitatively similar to the antenna position dependence.
- Current drive typically more effective at the 1st harmonic.

Conclusions and outlook

- AMR + LUKE is capable of complex and realistic EBW heating and current drive simulations.
 - The excited EBWs can efficiently be absorbed and drive current at certain radii. The current direction is opposite for above / below mid-plane launch.
 - The deposition location can be controlled by changing either Z_A or the frequency. The frequency-based control provides larger radial range.
 - More simulations have to be done and analyzed
 - NSTX H-mode
 - NHTX
 - other devices (MAST, MAST Upgrade)
 - EBW H&CD robustness and sensitivity will be addressed.
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