

Measuring electron gyro-scale fluctuations in NSTX plasmas using a collective scattering system

David R. Smith Princeton Plasma Physics Lab

In collaboration with:

E. Mazzucato and H.K. Park *Princeton Plasma Physics Lab*

W. Lee Pohang University

C.W. Domier and N.C. Luhmann, Jr. *University of California at Davis*

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Outline



- Motivation
 - Does electron temperature gradient (ETG) turbulence produce meaningful electron thermal transport?
- Principles of scattering measurements
- The NSTX high-k scattering system
 - Core fluctuation diagnostic with high sensitivity, spatial localization, k-space selectivity for up to five distinct wavenumbers, and steerable optics.
- Initial measurements and research topics

ETG turbulence and electron thermal transport

Transport sets the plasma size needed to attain fusion-relevant conditions

The perpendicular particle flux (Γ_{\perp}) and perpendicular energy flux (Q_{\perp}) are key transport quantities

0th moment:

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\vec{V}) = S_p \quad \stackrel{[dV]}{\Rightarrow} \quad \frac{\partial N}{\partial t} + \vec{\Gamma} \cdot \hat{r} = S_p V$$
1st moment (SS):

$$\vec{\Gamma}_{\perp} \equiv n\vec{V}_{\perp} = \frac{1}{qB}\hat{b} \times \left[\nabla \cdot \vec{P} - qn\vec{E} - \vec{S}_m\right]$$
2nd moment:

$$\frac{\partial}{\partial t} \left(\frac{3}{2}p + \frac{1}{2}mnV^2\right) + \left[\nabla \cdot \vec{Q}\right] = qn\vec{E} \cdot \vec{V} + S_e \quad \stackrel{[dV]}{\Rightarrow} \quad \frac{\partial E}{\partial t} + \vec{Q} \cdot \hat{r} = H_{tot}V$$
3rd moment (SS):

$$\vec{Q}_{\perp} = \frac{1}{\Omega}\hat{b} \times \left[\nabla \cdot \vec{R} - \frac{5}{2}\frac{q}{m}p\vec{E} - \frac{q}{m}\vec{\pi} \cdot \vec{E} - qn\vec{E} \cdot \left(\frac{1}{2}V^2\vec{I} + \vec{V}\vec{V}\right) - \vec{S}_{ef}\right]$$

terms with substantial fluctuation contributions

(species subscripts suppressed above)

Fluctuation-induced transport typically dominates other contributions

Fluctuation-induced particle and energy transport

$$\vec{\Gamma}_{\perp}^{fI} = \left\langle \widetilde{n} \widetilde{\vec{v}}_{E} \right\rangle \qquad \qquad \vec{Q}_{\perp}^{fI} = \frac{5}{2} \left\langle \widetilde{p} \widetilde{\vec{v}}_{E} \right\rangle$$

The phase between fluctuations is critical.

When \tilde{n} leads $\tilde{\phi}$ (non-adiabatic response), there is a net motion of particles down the density gradient.

The story is the same for energy and $\widetilde{
ho}$.

What drives electron thermal transport?



Electron thermal transport is typically the dominant loss channel in NSTX plasmas



Ion thermal transport is at or near neoclassical.

It is widely held that ExB flow shear stabilizes ITG/TEM turbulence in NSTX.

[see papers by Kaye (07, 06), Levinton (07), Stutman (06), LeBlanc (04), Synakowski (03), Bourdelle (03), Rewolt (96)]

Relation between energy flux Q_s and heat flux q_s :

$$\vec{Q}_{s} = \vec{q}_{s} + \frac{5}{2}p_{s}\vec{V}_{s} + O(\delta^{2})$$

Common thermal transport model:

$$\vec{q}_{s} = \chi_{s} \nabla T_{s}$$

Drift-interchange modes are analogous to the Rayleigh-Taylor instability



[G. Hammett, APS 2007 Review Talk]

Plasma turbulence is anisotropic with $k_{\perp} \gg k_{\parallel}$ NSTX =[Candy, Waltz, and the GYRO team]

 $k_{\perp}\rho_{x} \sim O(1)$ and $k_{\parallel}qR \sim O(1) \Rightarrow$ $kB >> \vec{k} \cdot \vec{B} \approx 0$ and $k_{\perp} >> k_{\parallel}$

Remember this for later...

ETG instability exhibits a critical gradient

ETG critical gradient from linear gyrokinetic simulations using GS2 [Jenko, PoP 8, 4096 (2001)]

$$\left(\frac{R}{L_{Te}}\right)_{crit} = \mathsf{Max}\left\{\left(1 + \frac{Z_{eff}T_{e}}{T_{i}}\right)\left(1.33 + 1.91\frac{\hat{s}}{q}\right)\left(1 - 1.5\varepsilon\right)\left(1 + 0.3\varepsilon\frac{d\kappa}{d\varepsilon}\right), \ 0.8\frac{R}{L_{ne}}\right\}$$

ETG critical gradient extends to nonlinear saturation and electron thermal transport





ETG turbulence can saturate at higher amplitude than ITG/TEM turbulence due to zonal flow dynamics $\boxed{0} NSTX = 0$





ExB flow shear

ETG turbulence generates zonal flows less efficiently than ITG/TEM turbulence because the adiabatic-ions in ETG turbulence have large gyro-orbits compared to ETG-driven zonal flows.

Consequently, ETG turbulence saturates at higher amplitude and produces more normalized transport than ITG/TEM turbulence. [see Dorland, PRL (2000) and Nevins, PoP (2006)]

$$\frac{\chi_e^{ETG}}{\chi_e^{gB}} \sim 2 - 5 \qquad \qquad \frac{\chi_i^T}{\chi_i^g}$$

$$\frac{\chi_i^{ITG}}{\chi_i^{gB}} \sim 0.5 - 2$$

Scattering measurements of plasma turbulence

Layout of a scattering system

- The intersection of the probe beam and receiving beam defines the scattering volume
- The scattering angle provides k-space selectivity through the Bragg relation: k = 2k_i sin(θ/2)



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Conservation relations describe three-wave coupling among EM waves and a plasma fluctuation

An incident EM wave (k_i, ω_i) and a plasma fluctuation (k, ω) couple to produce a scattered EM wave (k_s, ω_s) .

Momentum conservation:



+ gives anti-Stokes interaction- gives Stokes interaction

Energy conservation:

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$$\omega_{\rm s} = \omega_{\rm i} \pm \omega$$

Bragg relation follows from high frequency incident wave:

$$\omega_i \gg \omega \Longrightarrow k_i \approx k_s$$
 and

 $k = 2k_i \sin(\theta/2)$



The **plasma** performs the **Fourier transform** from real space to k-space Anisotropic turbulence and magnetic field variation combine to improve spatial localization



For scattering on the midplane, k's are purely radial. Some k's within the overlap volume produce scattered light (k_s) that does not intersect the receiver, so the scattering volume is constricted.

see Mazzucato, PPCF 48, 1749 (2006)

Density fluctuations encoded into scattered power spectrum



Scattered power due to a single, coherent fluctuation:

$$P_{s} = \frac{1}{4} P_{i} r_{e}^{2} \lambda_{i}^{2} \tilde{n}^{2} L_{v}^{2}$$

Thomson scattering \rightarrow EM wave scattering due to unbound electrons M NSTX =

Rayleigh scattering \rightarrow EM wave scattering due to bound electrons

Scattering cross-section, σ , derives from the dipole radiation of an electron, with natural frequency ω_0 , driven by an EM wave with frequency ω :



Classical electron radius:

 $r_{\rm e} = \frac{e^2}{4\pi a_{\rm e} m_{\rm e}^2}$

Two varieties of Thomson scattering: collective and incoherent

- Incoherent Thomson scattering
 - Scattering wavelength $\rightarrow \lambda \ll \lambda_D$ (λ =2 π /k)
 - Random electron motion \rightarrow incoherent radiation
 - Dipole powers additive \rightarrow $\rm P_{s} \propto n_{e}$
 - Near-IR sources $\rightarrow \lambda \thicksim 1 \ \mu m$
 - Scattering angle $\rightarrow 90^{\circ}$
 - Useful for n_e and T_e measurements
- Collective Thomson scattering
 - Scattering wavelength $\rightarrow \lambda \gtrsim \lambda_{\text{D}}$
 - Collective electron motion \rightarrow coherent radiation
 - Dipole powers multiplicative $\rightarrow P_{s} \propto n_{e}^{-2}$
 - Far-IR to mm-wave sources $\rightarrow \lambda \thicksim$ 10 μm to 4 mm
 - Scattering angles $\rightarrow 1^\circ$ to 40°
 - Useful for density fluctuation measurements

The NSTX high-k scattering system

NSTX "high-k" scattering system measures fluctuations up to $k_{\perp}\rho_{e} \cong 0.6$

- 280 GHz (λ =1 mm) scattering system
- Instrumental minimum detectable fluctuation is ñ_e/n_e~10⁻⁵
- Five detection channels
 - k_{\perp} spectrum at five discrete k_{\perp}
 - ω spectrum from time domain sampling
- Tangential scattering
 - Probe and receiving beams nearly on equatorial midplane
 - System sensitive to radial fluctuations
- Steerable optics
 - Scattering volume can be positioned throughout the outer half-plasma
- First data during FY06 run campaign



Five detection channels can measure fluctuations at five discrete wavenumbers





Steerable optics enable good radial coverage

Intermediate $\rho = 0.4$

 $\longrightarrow \bigcirc NSTX = -$

Outboard $\rho = 0.75$ $k_{\perp}\rho_e$ up to 0.2



 $k_{\perp}\rho_{e}$ up to 0.7 $k_{\perp}\rho_{e}$ up to 0.3 $k_{\perp}\rho_{e}$ up to 0.3 $k_{\perp}\rho_{e}$ up to 0.3 $k_{\perp}\rho_{e}$ up to 0.3 $k_{\perp}\rho_{e}$ up to 0.4 $k_{\perp}\rho_{e}$ up to 0.4

-1

-2.0

-1.5

-1.0

X (m)

-0.5

0.0

2 Vacuum Vessel LCFS Magnetic Axis (آ ح -1 -2.0 -1.5 -1.0 -0.5 0.0 X (m)

Inboard $\rho = 0.05$

∠ک

Scattering system layout



- BWO source
 - outside test cell
 - ~100 mW at 280 GHz

NSTX =

- Overmoded, corrugated waveguide
 - low-loss transmission
- Steerable optics
 - quasi-optical design
- Heterodyne receiver
 - five channels
 - reference signal extracted from main beam

Scattering system pictures

waveguide and launch optics



collection optics



NSTX =



heterodyne receiver

collection mirror





Quasi-optical design positions beam waist at scattering location and couples scattered light into receiver



- 3 cm beam waist (1/e² intensity radius) at the scattering location
- k-space resolution $\implies \Delta k \approx 0.7 \text{ cm}^{-1}$

NSTX=

Quadrature heterodyne receiver samples at 7.5 MS/s





Ray tracing calculations are needed to design configurations prior to experiments



Launch mirror (probe beam) Horizontal actuator: 1.117 in. Horizontal angle wrt X-axis: 83.0° Vertical actuator: 1.217 in. Vertical angle wrt midplane: -6.0°

Collection mirror Horizontal actuator: 2.144 in. Horizontal angle wrt X-axis: 232.5° Vertical angle wrt midplane: 5.4° (Vertical angle not adjustable)

Exit window angles						
	Ch. 2	Ch. 3	Ch. 4	Ch. 5		
Hor. Ang.	-6.4°	-3.9°	-0.7°	2.5°		
Vert. Ang.	-0.2°	-0.7°	-1.1°	-1.5°		

Ray tracing calculations are needed to ensure k-space alignment



	Ch. 2	Ch. 3	Ch. 4	Ch. 5
r/a	0.27	0.28	0.29	0.30
d _{min} (cm)	0.1	0.1	0.1	0.1
k _∥ (cm ⁻¹)	0.1	0.0	0.2	0.0
k _r (cm ⁻¹)	6.9	11.0	14.6	17.8
k _θ (cm⁻¹)	-1.6	-3.4	-4.4	-5.5
k _⊥ (cm ⁻¹)	7.1	11.5	15.2	18.6
k _θ /k _r	0.23	0.30	0.30	0.31
$k_{\perp} \rho_e$	0.23	0.38	0.51	0.62
$k_{\perp} \rho_s$	14	22	30	37
k _T (cm ⁻¹)	-0.4	-0.7	-1.2	-1.3
f _D (MHz)	-0.95	-1.78	-2.98	-3.26
d _{tan} (cm)	-14.5	-18.4	-23.5	-25.5

Data is not junk

NSTX=___

Electron diamag. direction

Doppler shift in ion direction Offset from PB tangency

Ray tracing calculations are needed to determine scattering volume size NSTX =Localization along **Probe Beam** Vacuum Vessel 1.0 F k_{perp} = 22.1 cm⁻¹ LCFS $\theta = 21.8^{\circ}$ Magnetic Axis 0.0 $= 12.4 \text{ cm}^{-1}$ beam $\mathsf{k}_{\mathsf{perp}}$ Arbitary Units $\theta = 12.2^{\circ}$ ۲ (m) overlap instrument 0.0 selectivity = 4.4 cm⁻¹ **k** due to -1 $= 4.4^{\circ}$ $k_{\perp} \gg k_{\parallel}$ 0.0 0 -50 50 -2.0 -1.5 -1.0 -0.5 0.0

Dist. from Probe Beam Tang. (cm)

X (m)

30

Ray tracing calculations are needed to account for refraction



NSTX =

Ray tracing calculations are needed to discriminate good data from junk data



Ray tracing calculations are needed to interpret measurements

The time evolution of turbulence parameters, such as $k_{\mu}\rho_{e}$ and k_{θ}/k_{r} , are needed for comparing data to simulations.



NSTX = -

Initial measurements and research topics

Stray light spectral peak and Doppler-shifted fluctuations are common features

High-k measurements at R \simeq 135 cm and r/a \simeq 0.68 with k $_{\rm l}\rho_{\rm e}$ ~ 0.2-0.3



Fluctuations initially appear in the **electron direction**, then Doppler-shift to the **ion direction** due to V_{T} .

NSTX = -

Prominent, persistent fluctuations observed in core

High-k measurements at R \simeq 113 cm and r/a \simeq 0.25 k_⊥ $\rho_e \sim 0.35$ -0.40 for channel 5



Features appear in the ion direction due to Doppler shift

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Core fluctuations grow when R/L_{Te} is near critical value



Prominent, persistent fluctuations observed in outer-plasma



Ion/electron directions are reversed from previous slide. Fluctuations again experience a Doppler-shift to ion direction.

Using only HHFW heating, unshifted fluctuations observed in core



Without toroidal rotation from NBI to produce a Doppler-shift to the ion direction, fluctuations appear in electron direction.

In some HHFW discharges, peak T_e obtained when R/L_n reverses and fluctuations drop

High-k measurements at R \simeq 120 cm and r/a \simeq 0.3



This is a preliminary observation; counter examples exist.

Core fluctuations change character after RS collapse

High-k measurements at R \simeq 120 cm and r/a \simeq 0.35 $k_{1}\rho_{e} \sim 0.15-0.25$ for channel 4 and 0.25-0.35 for channel 5 124942 - Ch. 4 124942 - Ch. 5 124942 -15 8 Frequency (MHz) 252.0 ms ⁻requency (MHz) 2 2 -20 292.0 ms Power (dB) 6 -25 0 0 σ4 -30 -2 -2 -35 0.24 0.28 0.32 0.24 0.28 0.32 0.20 0.20 100 110 120 130 140 150 Time (s) Time (s) Radius (cm) 124942 - Ch. 4 124942 - Ch. 5 124942 2.0 -15 -15 249.9 ms 249.9 ms 1.5 289.9 ms 289.9 ms -20 -20 Power (dB) Power (dB) 1.0 Shear -25 0.5 -25 0.0 -30 -30 -0.5 100 110 120 130 140 150 -2 -2 0 2 2 0 Radius (cm) Frequency (MHz) Frequency (MHz)

This is a preliminary observation due to stray light concerns.

NSTX = -

The NSTX high-k system is the only core fluctuation diagnostic capable of addressing these questions:

k-space isotropy

- Are fluctuations isotropic in the k_{θ} - k_r plane?
- Steerable optics provide the capability to address this.

Mode coupling

- What is the phase coherence among three turbulent fluctuations that satisfy frequency and wave vector matching conditions?
- Multiple detection channels provide the capability to address this.

Summary



- Transport is important
- ETG turbulence may be important
- Collective scattering is a powerful technique for measuring fluctuations with spatial localization and k-space selectivity
- The NSTX high-k system is versatile due to steerable optics and multiple detection channels
- Electron gyro-scale fluctuations have been observed exhibiting a variety of dynamics in a variety of NSTX plasmas

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