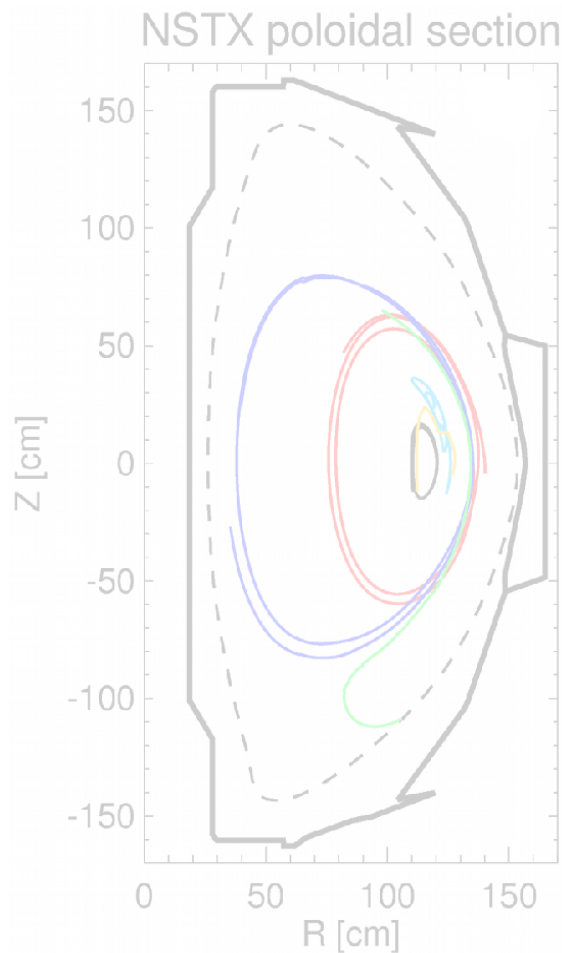


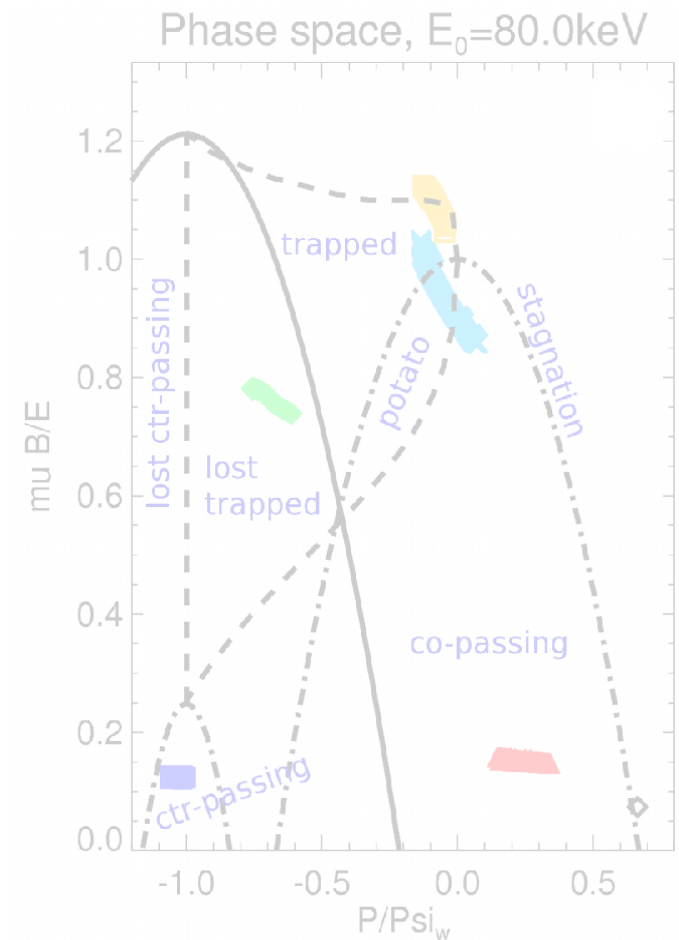
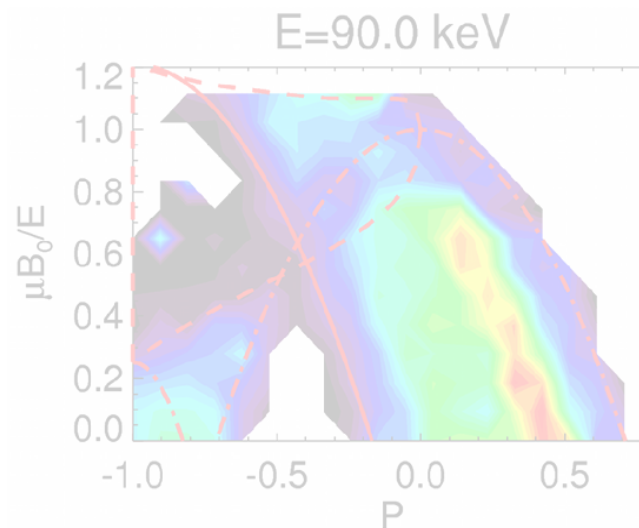
Interpretive model for resonant fast ion transport by Alfvénic instabilities



M. Podestà
and many contributors

PPPL, B252

Aug. 28th 2012



Motivation

- Resonant/stochastic fast ion transport may not be well described by *diffusive* processes, $\Gamma \sim dn/dr$
 - Resonances introduce selectivity in phase space (Energy, pitch)
- Fast ion transport models presently implemented in TRANSP/NUBEAM do not include basic physics of wave–particle resonant interaction
- Codes such as M3D–K do treat resonant interaction correctly, but are not designed for ‘mass production’
- *New “general/reduced” model is required to complement existing models*
 - *Must capture basic mechanisms of wave–particle interaction*
 - *Formulation must be compatible with TRANSP/NUBEAM*

Outline

- Scope of new model
- Main ingredients
 - Choice of variables
 - Constraints from resonant interaction with *AEs
- General implementation
 - *Probability distribution* to evolve particles' orbits
- Questions, comments, open issues

Backup:

- Example#1: computing the 'Mode Amplitude'
- Example#2: deriving the transport coefficients
- Example#3: evolving F_{nb} in time

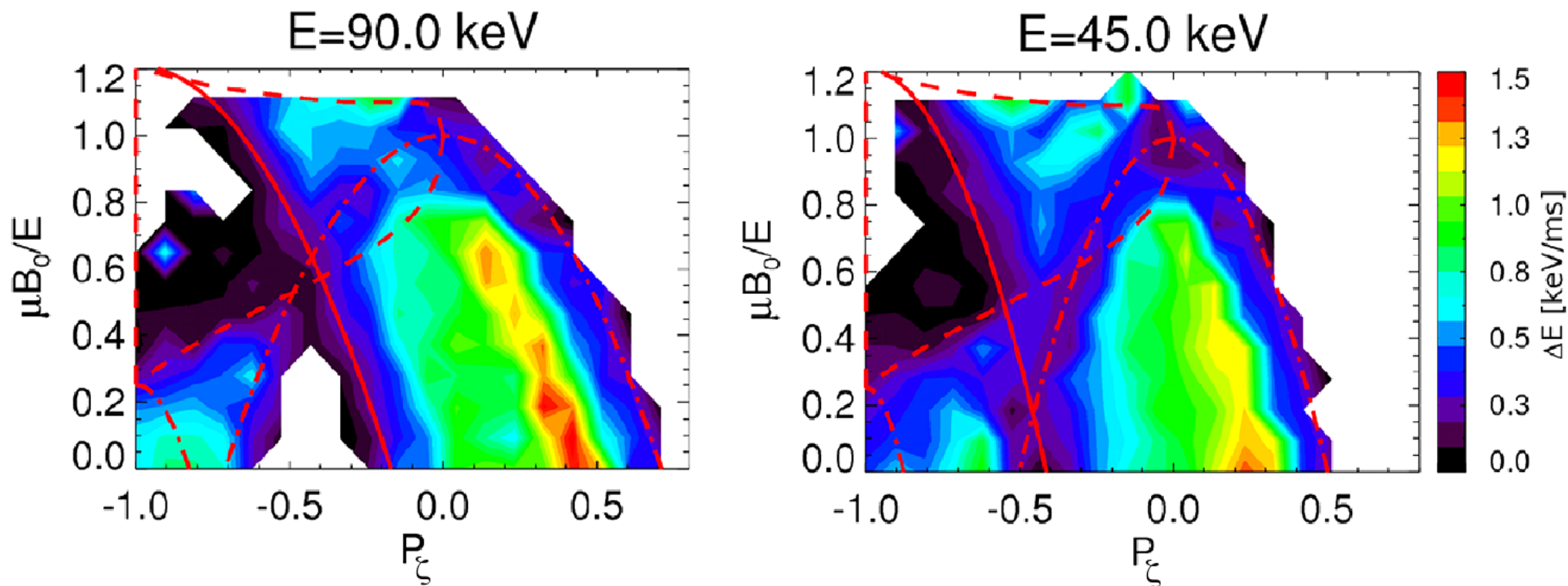
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Goal: simulate *long-term* effects on F_{nb} from Alfvén Eigenmodes (*AEs) with arbitrary amplitude evolution $A_{modes}(t)$



- Example: average energy step at fixed energy, A_{modes}
 - ORBIT simulation with 3 TAE modes from NOVA-K
- Regions with significant resonant interaction well identified in (E, P_ξ, μ) space
- *How can these variations be accounted for in general-purpose codes such as TRANSP?*

All fast ion transport models presently implemented in TRANSP have *diffusive* nature; little/no phase space selectivity

- Four models presently implemented in TRANSP for fast ion diffusion coefficient, D_b

[from <http://w3.pppl.gov/~pshare/help/transp.htm>]

$$1) D_b = k_{ADIFB} \times D_e$$

k_{ADIFB} : multiplier

$D_e(x,t)$: electron particle diffusivity

$$2) D_b = k_{ADIFB} \times D_e^{WP}$$

$D_e^{WP}(x,t)$: electron particle diffusivity from Ware-pinch corrected flux

$$3) \Gamma_{fi} = -D_b \nabla n_b + v_b n_b$$

diffusion/convection model

$$4) D_b(E, t, x) = \sum_k \alpha_k D_k$$

$D_k(E,x,t)$: diffusivity for deeply trapped, barely trapped, co-barely passing, ...

New model introduces selectivity in phase space, generalizes “diffusive transport” paradigm

- Info on phase space dynamics paramount for Verification&Validation of codes, theory–experiment comparison : *need phase–space representation*
- Steady–state (“fluid”) equilibrium does not provide all information we need : *need E, pitch details*
 - E.g. phase space transiently (and selectively) modified by *AE bursts : effects propagate during slowing–down
 - ⇒ *Solutions for F_{nb} based on integral quantities (f.i. density, neutrons, E_r /rotation, NB–driven I_p , ...) not unique*
- Time–scale τ_{res} for resonant/stochastic transport can be faster than diffusion time : *implementation?*
$$\tau_{transit} \leq \tau_{res} < [d(\log A_{modes})/dt]^{-1} \ll \tau_{diff} \text{ (typically)}$$
- *The new model must be “simple” enough to be included in TRANSP/NUBEAM for routine use*
 - *Retain only minimum required amount of information to represent resonant/stochastic fast ion response to *AEs*
 - *e.g., no attempts to follow “single–particle” orbits*

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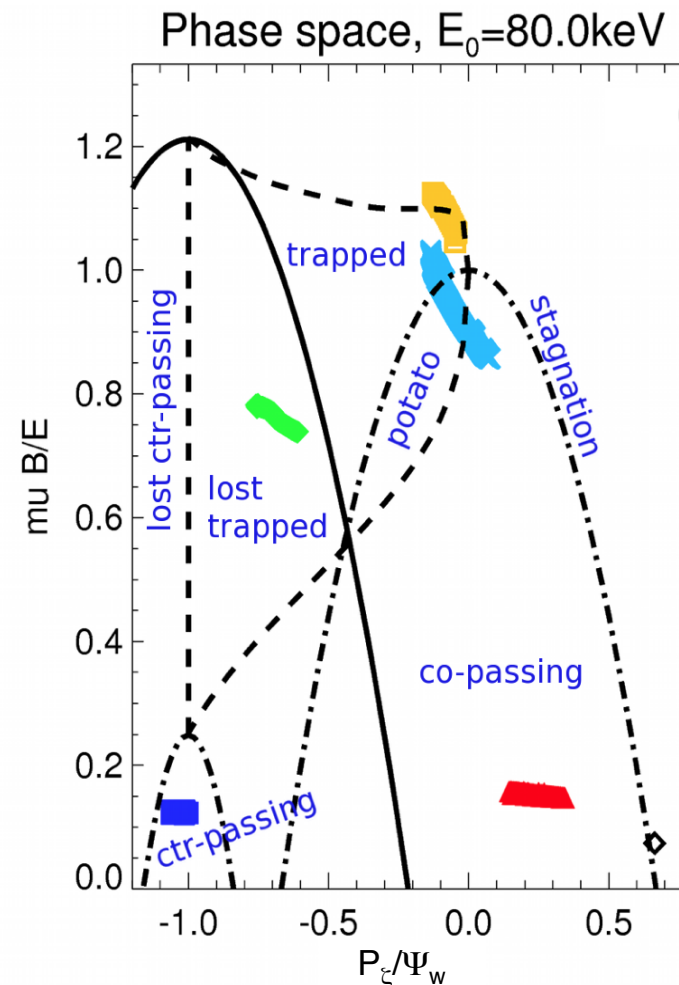
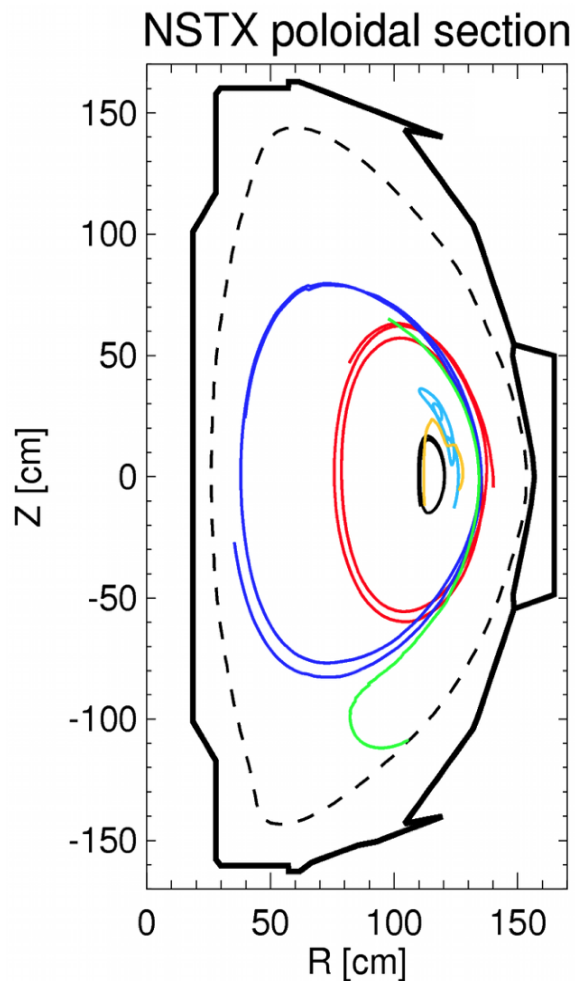
Backup:

- Example#1: computing the 'Mode Amplitude'
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'Constants of motion' are the natural variables to describe resonant wave-particle interaction

- Each orbit fully characterized by:

E , energy
 P_ζ , canonical angular momentum
 μ , magnetic moment



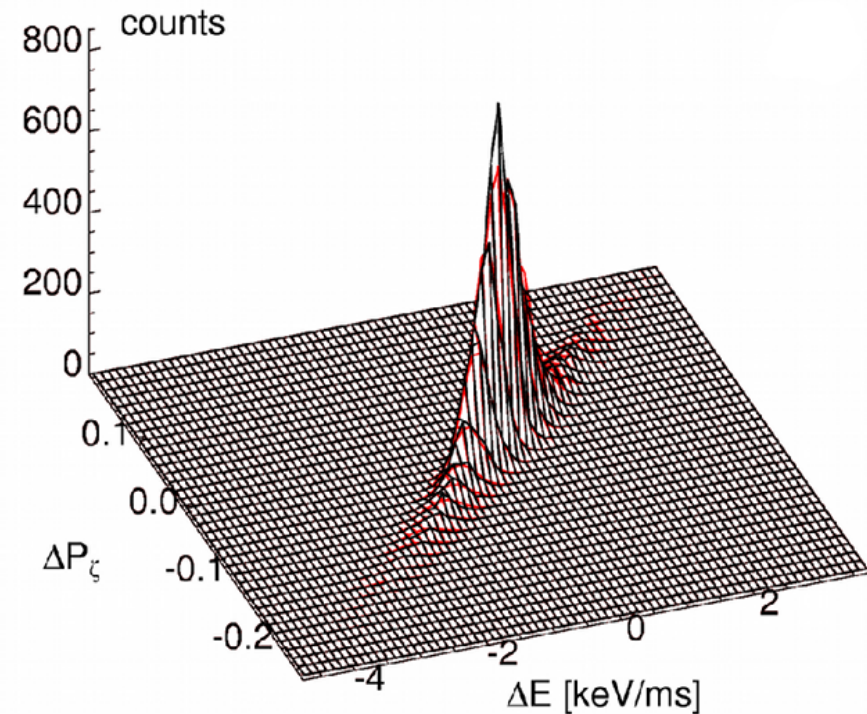
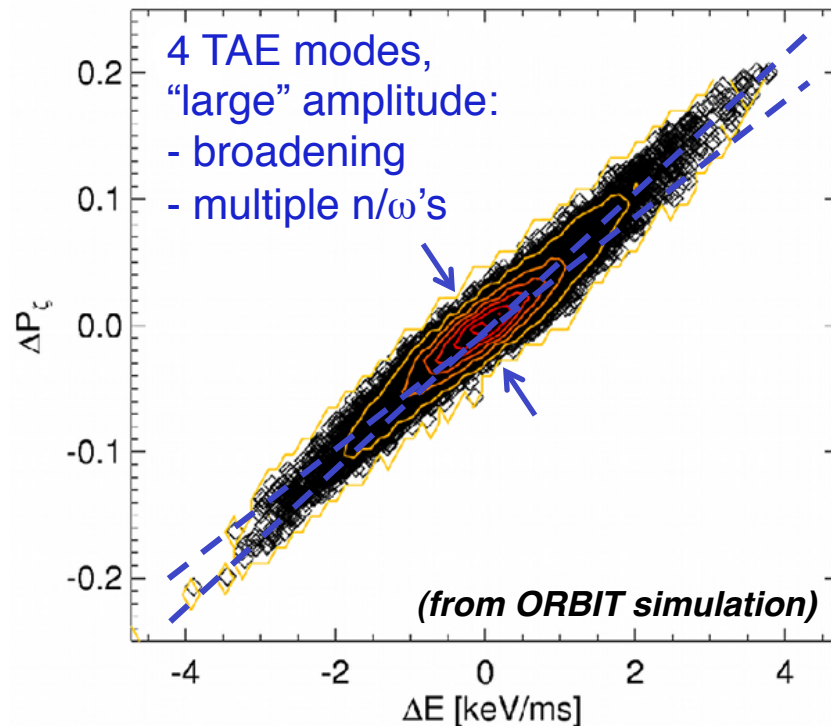
Single (isolated) resonances introduce simple constraints on particle's trajectory in (E, P_ζ, μ)

- From Hamiltonian formulation:

$$\omega P_\zeta - nE = \text{const.} \implies \Delta P_\zeta / \Delta E = n / \omega$$

$\omega = 2\pi f$, mode frequency

n , toroidal mode number



Presence of multiple modes/resonances distorts the 'ideal' (linear) relationship

For low-frequency *AEs with $\omega \ll \omega_{ci}$ such as TAEs, magnetic moment μ is conserved (...well, *almost*)

- *In these notes, it is assumed that $\Delta\mu=0$*
- However: $\Delta\mu=0$ hypothesis may break down if
 - $\rho_f \sim$ radial width of the modes
 - $\rho_f \sim$ scale-length of equilibrium profiles
 - ⇒ *Both conditions are likely to be met in spherical tokamaks (e.g. NSTX)*
- ⇒ **Proposed model can be generalized to cases where μ is *not* conserved**
 - Also important for ω_{ci} -range instabilities: GAE/CAEs

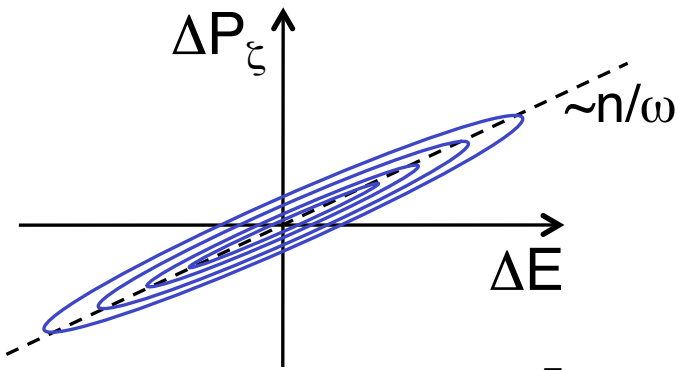
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Single-resonance case: introducing the *probability distribution function* for particle transport



For each *bin* in (E, P_ξ, μ) , steps in $\Delta E, \Delta P_\xi$ can be described by bivariate $p(\Delta E, \Delta P_\xi | P_\xi, E, \mu, A)$

$$p = p_0 e^{-\frac{1}{2(1-\rho)} \left[\frac{(\Delta E - \Delta E_0)^2}{\sigma_E^2} + \frac{(\Delta P_\xi - \Delta P_{\xi 0})^2}{\sigma_{P_\xi}^2} - 2\rho \frac{(\Delta E - \Delta E_0)(\Delta P_\xi - \Delta P_{\xi 0})}{\sigma_E \sigma_{P_\xi}} \right]}$$

$$p_0 = \frac{1}{2\pi \sigma_E \sigma_{P_\xi} \sqrt{1 - \rho^2}} \quad \text{normalization}$$

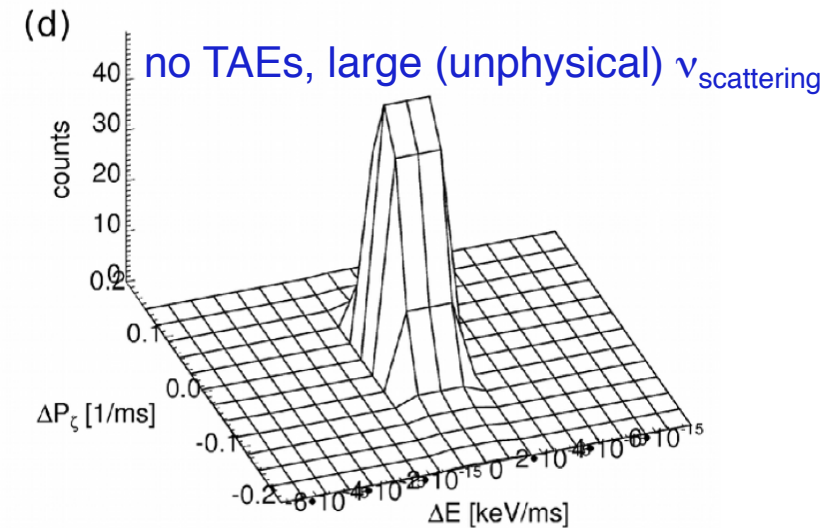
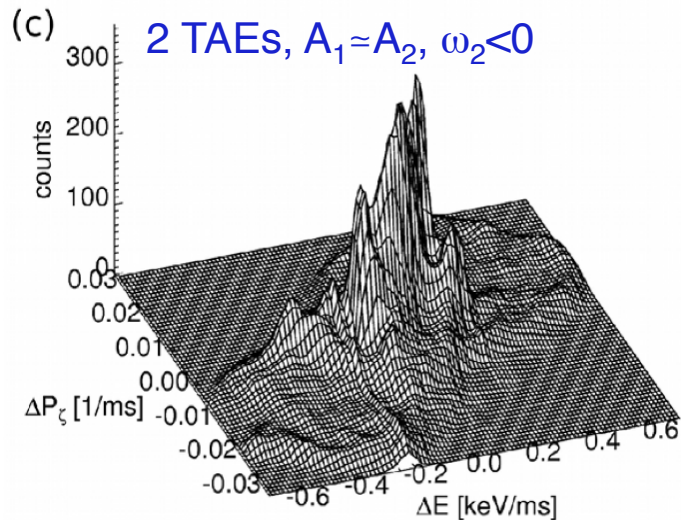
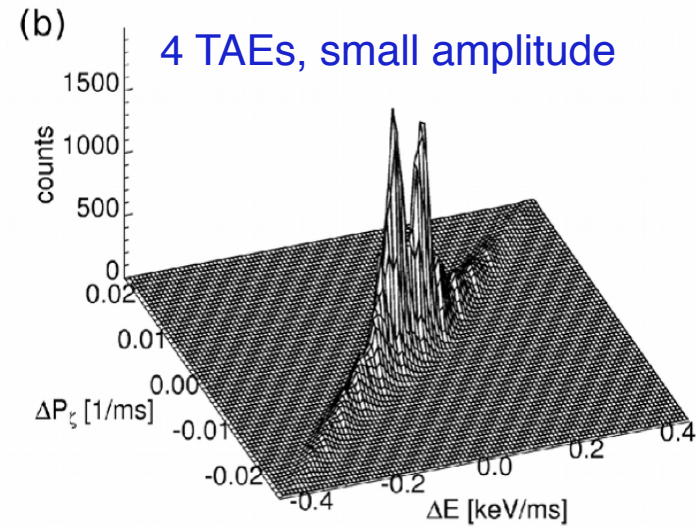
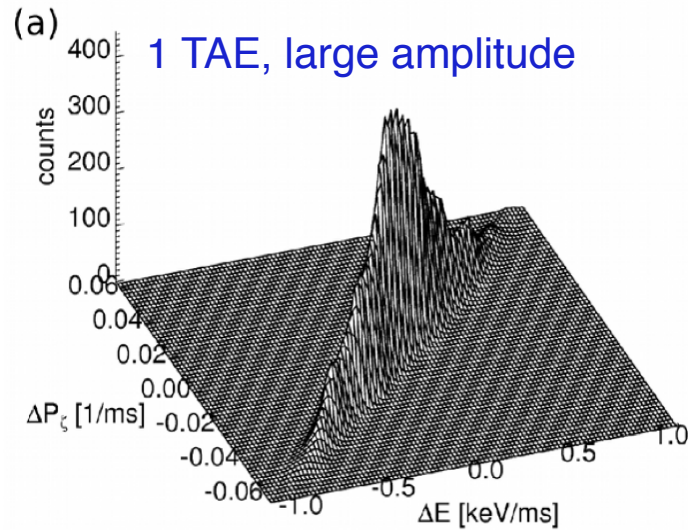
$$\rho = \frac{\langle (\Delta E - \Delta E_0)(\Delta P_\xi - \Delta P_{\xi 0}) \rangle}{\sigma_E \sigma_{P_\xi}} \quad \text{correlation parameter,}$$

such that $\Delta P_\xi(\Delta E) = \Delta P_{\xi 0} + \text{sign}(\rho) \times \frac{\sigma_{P_\xi}}{\sigma_E} (\Delta E - \Delta E_0)$

...and all parameters depend, in principle, on the mode amplitude $A=A(t)$

Analytical formulation *could* be extended to multi-mode case – but it would be quite unpractical

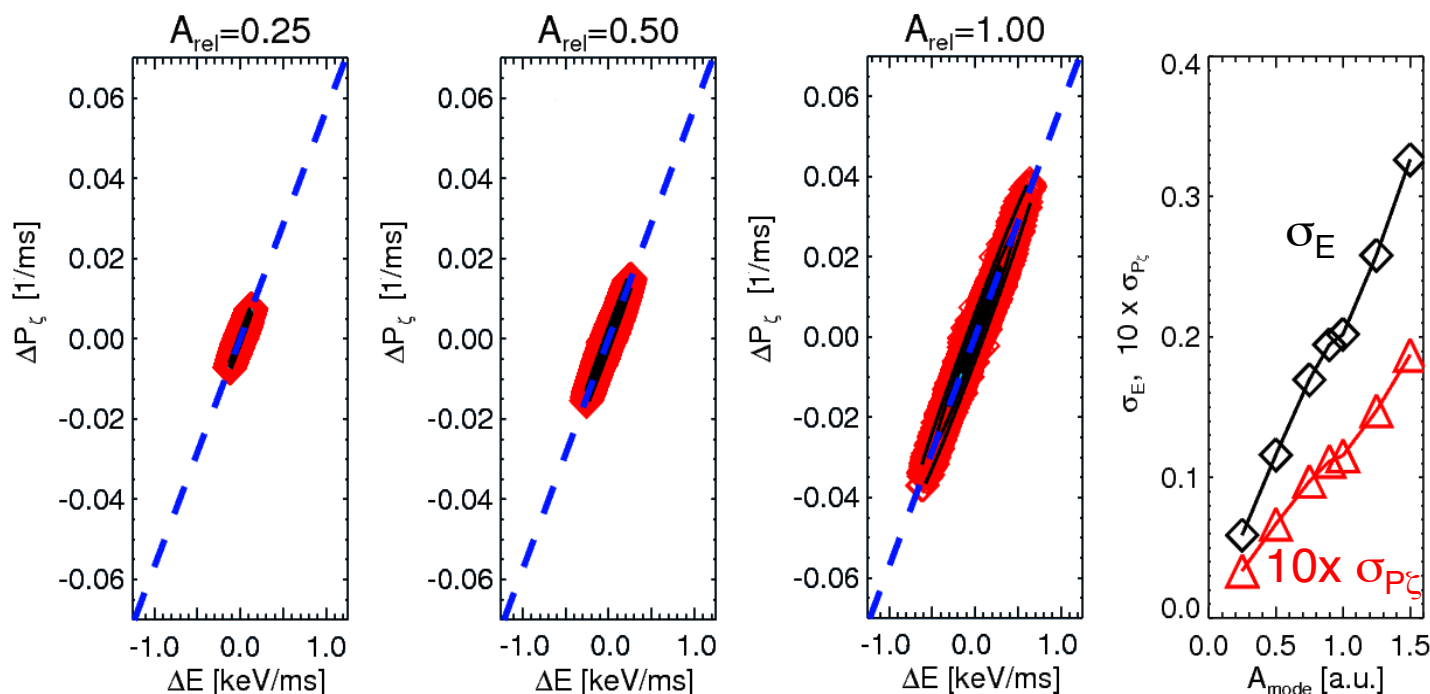
E.g., use: $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu) = \sum_{i=1}^N w_i p_i(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu)$



In practice, too many parameters to be passed to TRANSP/NUBEAM

General case: reduce number of parameters; dependence on $A=A_{\text{modes}}(t)$ can be simplified; use raw $p(\Delta E, \Delta P_\xi)$ directly

- ‘Convective’ terms $\Delta E_0, \Delta P_{\xi 0}$ are =0 if slowing down is accounted for elsewhere
- Variances σ_E, σ_{P_ξ} are roughly linear with A_{modes} :
 - Specify $p(\Delta E, \Delta P_\xi)$ for $A_{\text{modes}}=1$ only, then re-scale $\sigma \rightarrow \sigma' = A_{\text{modes}} \times \sigma$



⇒ Pass the raw matrix $p(\Delta E, \Delta P_\xi | P_\xi, E, \mu)$ to TRANSP/NUBEAM directly - no analytical form required

Layout of TRANSP/NUBEAM implementation (as seen from an experimentalist...)

- Compute $p(\Delta E, \Delta P_\zeta | P_\zeta, E, \mu)$
 - ORBIT, SPIRAL, theory
 - *Generalize to 6D matrix $p(\Delta E, \Delta P_\zeta, \Delta \mu | P_\zeta, E, \mu)$ if μ is not conserved
- Pass ‘Ufiles’ to TRANSP/NUBEAM
 - 2 files: 5D matrix* for $p(\Delta E, \Delta P_\zeta, \Delta \mu | P_\zeta, E, \mu)$, vector for $A_{\text{modes}}(t)$
- Evolve F_{nb} at each step:
 - I. Re-normalize bins (P_ζ, E, μ) based on q-profile, fields, ...
 - II. Identify ‘bin’ in (P_ζ, E, μ) for current ‘particle’ (i.e. orbit)
 - III. Extract $\sigma_E, \sigma_{P_\zeta}$ (σ_μ) from multivariate $p(\Delta E, \Delta P_\zeta, \Delta \mu)$
 - IV. Calculate step magnitude, based on $A_{\text{modes}}(t)$
 - V. Compute slowing down, scattering
 - VI. Advance particle’s trajectory in phase space
(steps II–VI actually divided in sub-steps)

Loop over particles

Outline

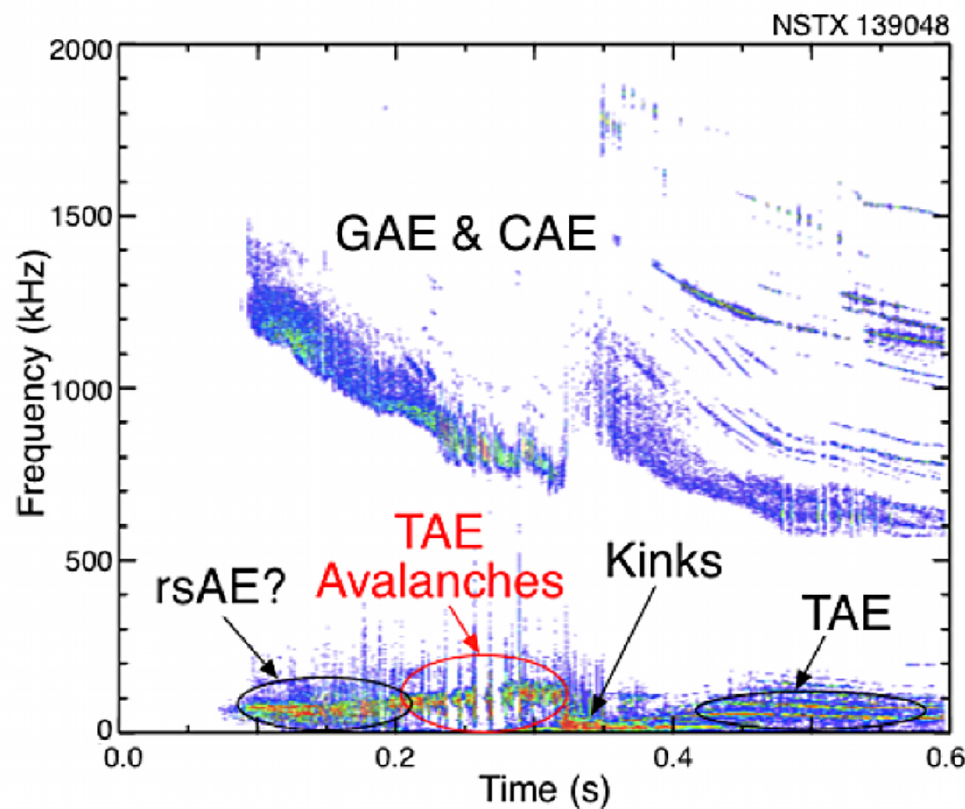
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Generalize to multiple MHD classes?

- Scenarios with presence of more than one MHD type of modes are quite common
 - NSTX: kink + GAE/CAEs + TAEs
- Formulation discussed so far can only deal with one class at the time
- However, this general case is of great interest
 - More realistic F_{nb} evolution when modes have comparable amplitude
 - Can mock-up scenarios such as 'fast ion channeling'
- To extend the model to multi-class case, each class can be represented by its $p_k(\Delta E, \Delta P_\xi, \Delta \mu)$
 - MonteCarlo scheme becomes more complicated: how to decide which class is 'active' for a specific bin (P_ξ, E, μ)?
 - Is the effect of different classes simply additive?
 - *Is this something worth pursuing from the beginning?*



Use the model in 'predictive' mode?

- As it is, the model would be OK to analyze 'real' discharges
 - Need mode structure to calculate $p(\Delta E, \Delta P_{\xi})$, e.g. from NOVA-K and reflectometers' data
 - Need data (Mirnovs, neutrons, etc.) to infer $A_{\text{modes}}(t)$
 - Two possibilities for 'predictive' runs:
 - Find reasonable guess for unstable modes (e.g. NOVA-K)
 - Explore different scenarios w/ scan of $A_{\text{modes}}(t)$: weak *AE activity, bursts/avalanches, etc.
- or:
- Have an additional module to compute $A_{\text{modes}}(t)$ self-consistently
 - Still requires mode structure, probably estimates for γ_{drive} , γ_{damp}
 - Let $A_{\text{modes}}(t)$ evolve according to F_{nb} evolution - how?
 - *Couple to Quasi-Linear model? Use Q-L instead?*

Additional questions, comments, open issues

Scope/approach:

- Which particular aspects targeted by new model are not included in existing models?
- Are there too many (T)AE resonances that can be accounted for by the model?
- Alfvénic time scale too fast for diffusion, look for steady-state solution?
- Approach looks OK, but is it providing too much detailed info?

Implementation:

- Should $\Delta\mu \neq 0$ case be included by default?
- What new code is required in TRANSP/NUBEAM? Need to include parts from ORBIT (SPIRAL, else)?
- How to optimize sub-steps for orbit calculation? Take $A(t)$ time-scale as reference?
- Is ‘normalization’ of $p(\Delta E, \Delta P_\xi)$ vs. time using B_0, Ψ_w, \dots OK?
- Is simple re-scaling $\sigma_E' = A_{\text{mode}} \times \sigma_E$ OK, or need to specify a separate scaling function $\sigma_E' = f(A_{\text{mode}}) \times \sigma_E$?

Test/check-out:

- May need to start with “simple” case, e.g. internal kink w/ $\Delta E = 0, \Delta P_\xi \neq 0$ and no energy dependence for given resonance.

BACK UP

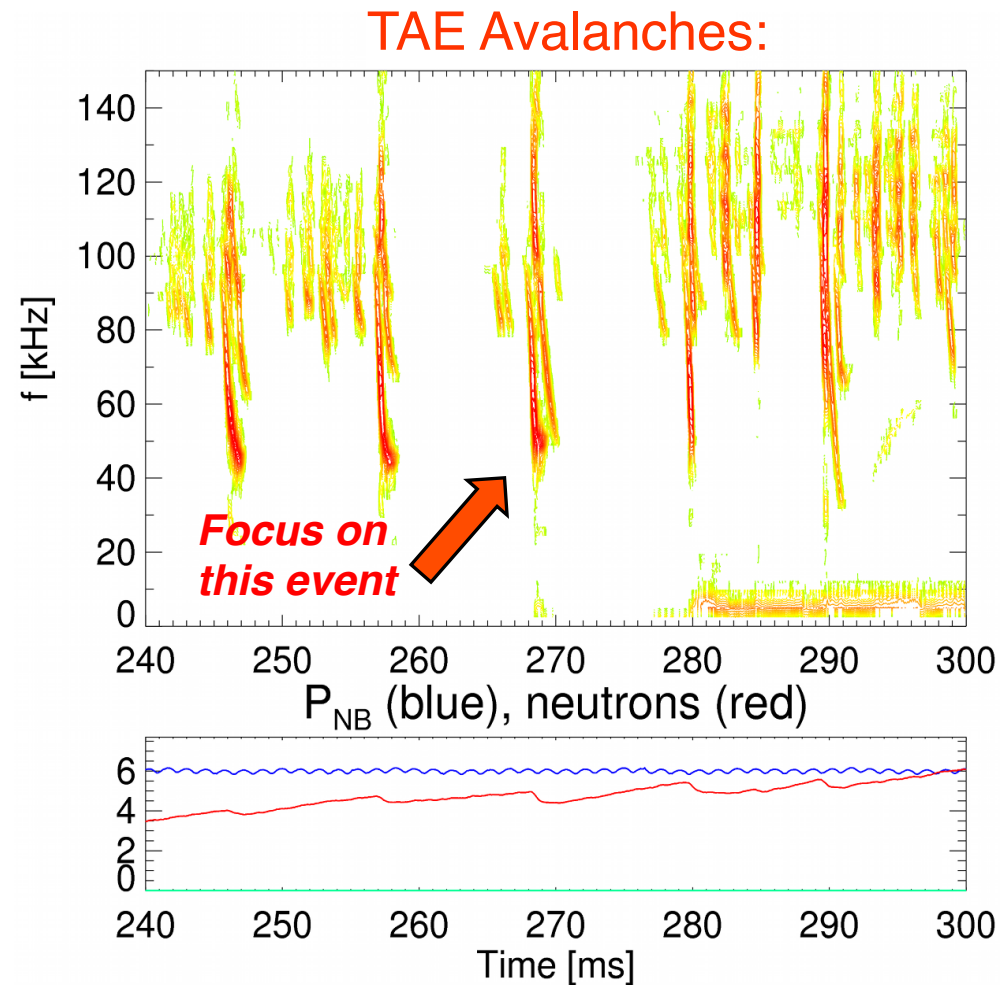
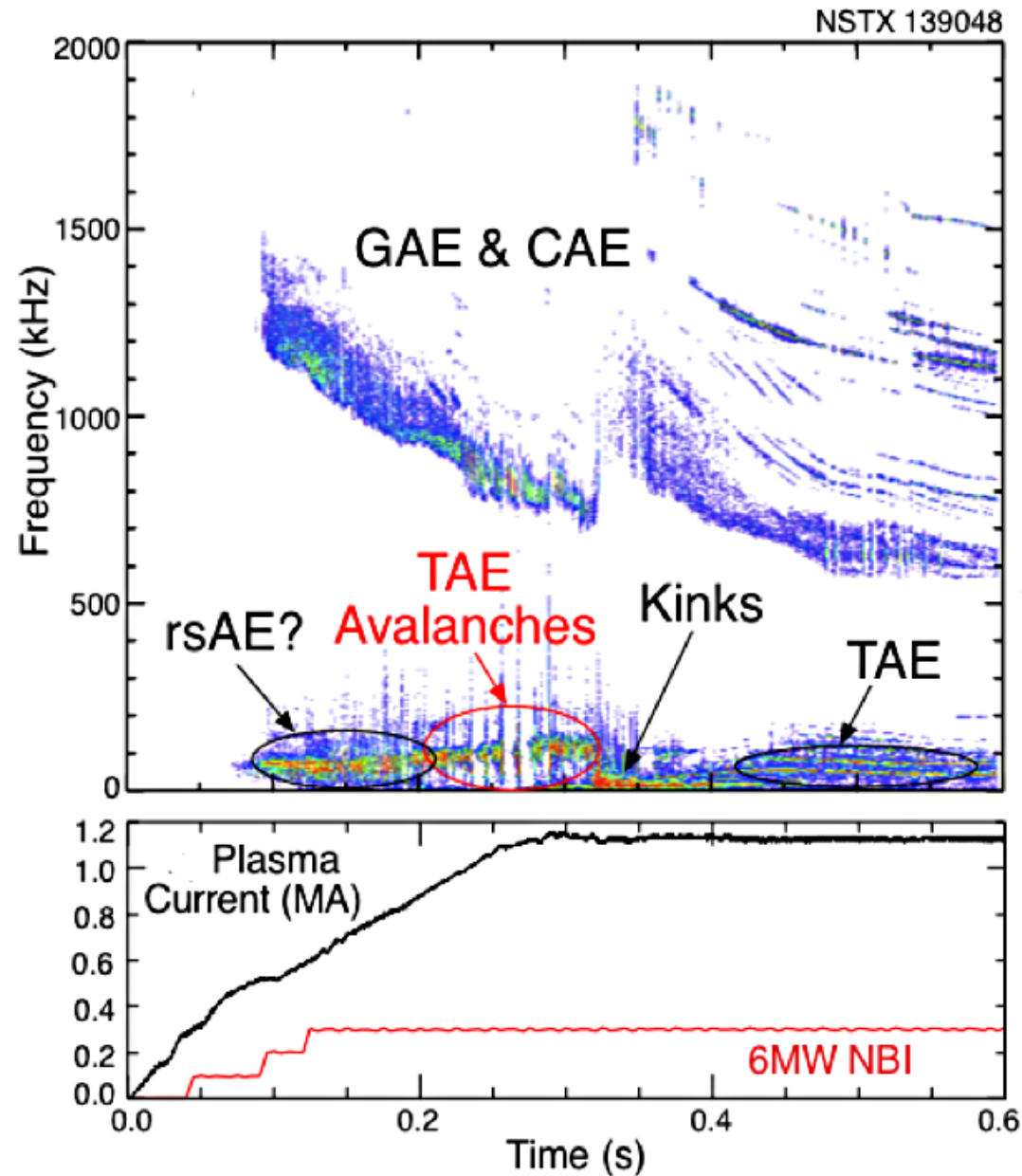
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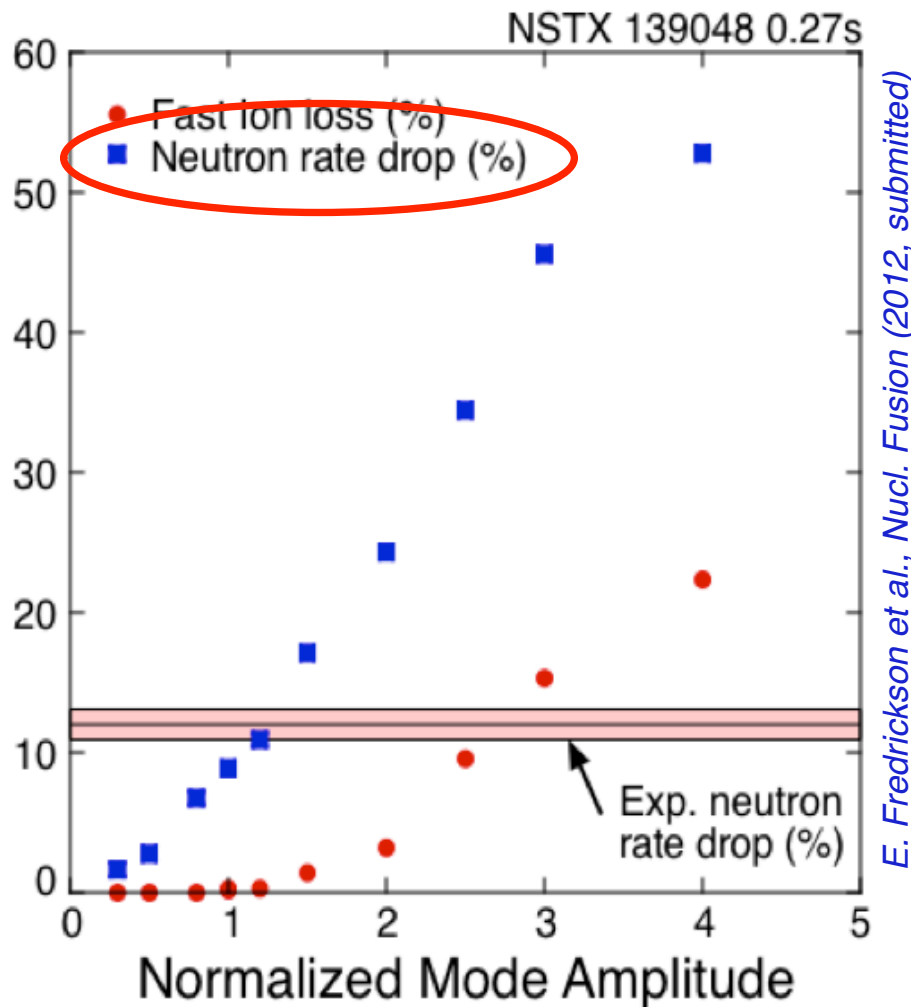
Experimental scenario for tests presented in these notes: NSTX H-mode plasma with bursts of TAE activity



E. Fredrickson et al., Nucl. Fusion (2012, submitted)

Ex#1: compute the 'mode amplitude', $A(t)$ / 1

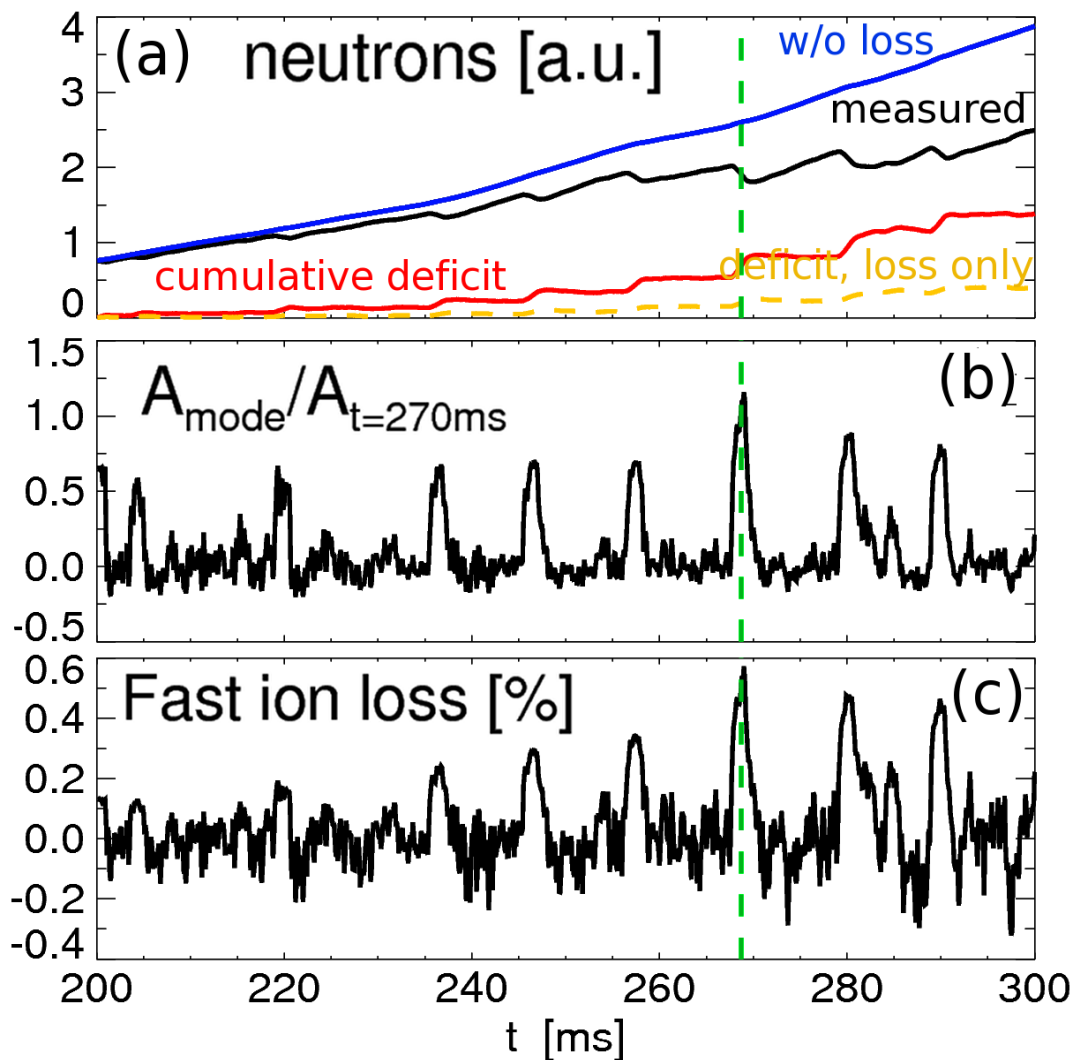
- First step is to obtain the relation between mode amplitude and 'transport':



- Ideal modes from NOVA-K
- Rescale modes based on comparison with reflectometers
- Compute transport, expected neutron drop with ORBIT
- Scan mode amplitude w.r.t. experimental one, $A=1$

Ex#1: compute the 'mode amplitude', $A(t)$ /2

- Get $A(t)$ from measured neutrons + 'table look-up':



- Compute fractional R_n drops vs. time
- Use figure from previous slide to find corresponding (normalized) mode amplitude

Outline

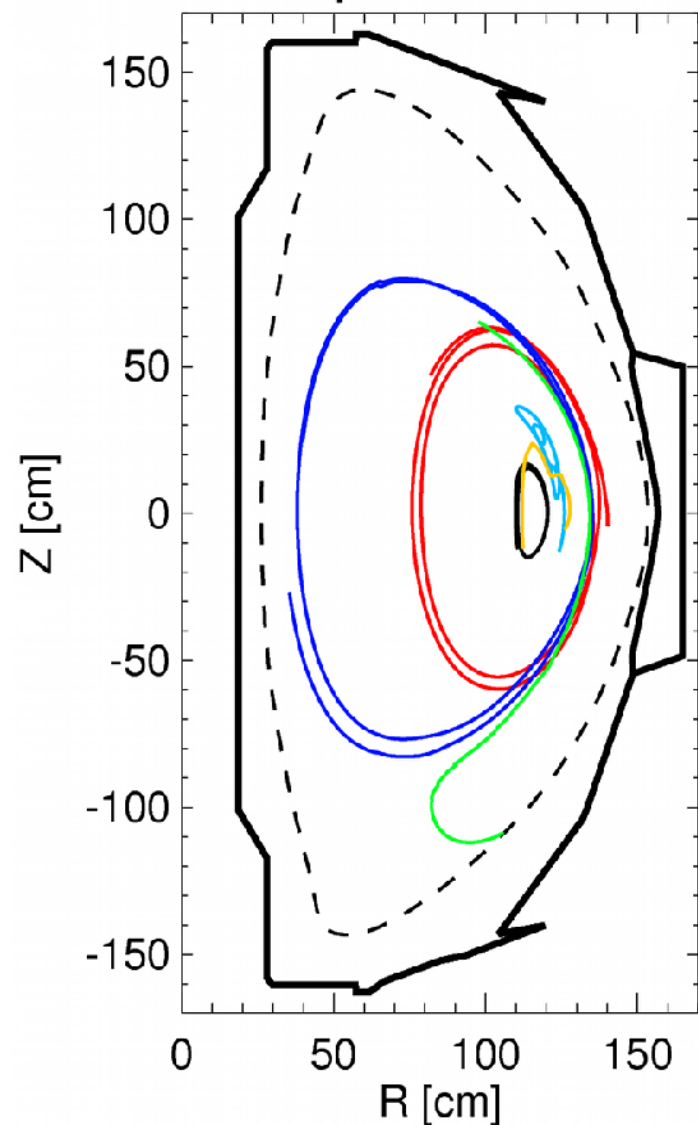
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Ex#2: deriving the transport coefficients

NSTX poloidal section



- Run ORBIT with $A=1$, constant
- Simulation time long enough (~ 1 ms) to capture *AE effects
- Track (P_{ξ}, E, μ) in time for each particle, steps $\delta t_{\text{sim}} \sim 50 \mu\text{s}$
- Compute $\Delta E, \Delta P_{\xi}, D\mu$
- Re-bin over (P_{ξ}, E, μ) space
- Get $p(\Delta E, \Delta P_{\xi} | P_{\xi}, E, \mu)$ for each bin

Outline

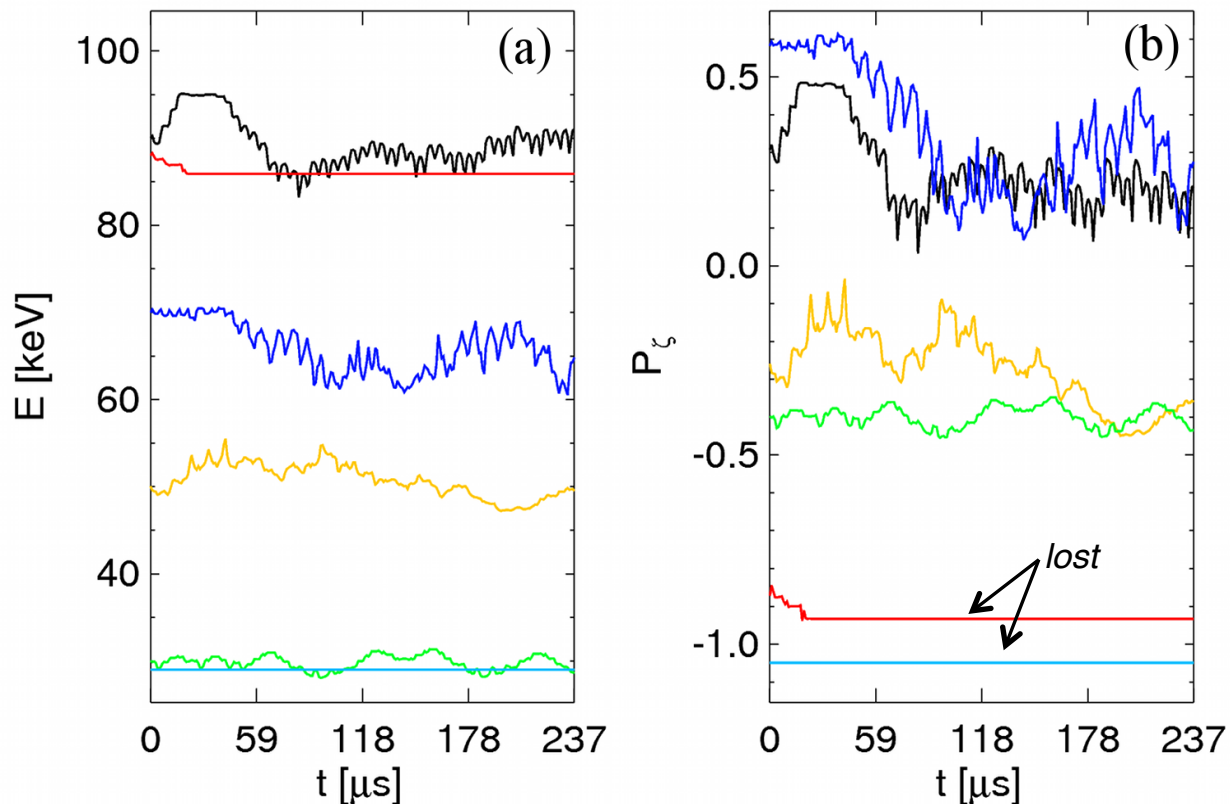
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- **Example#3: evolving F_{nb} in time**

Ex#3: evolving F_{nb} in time

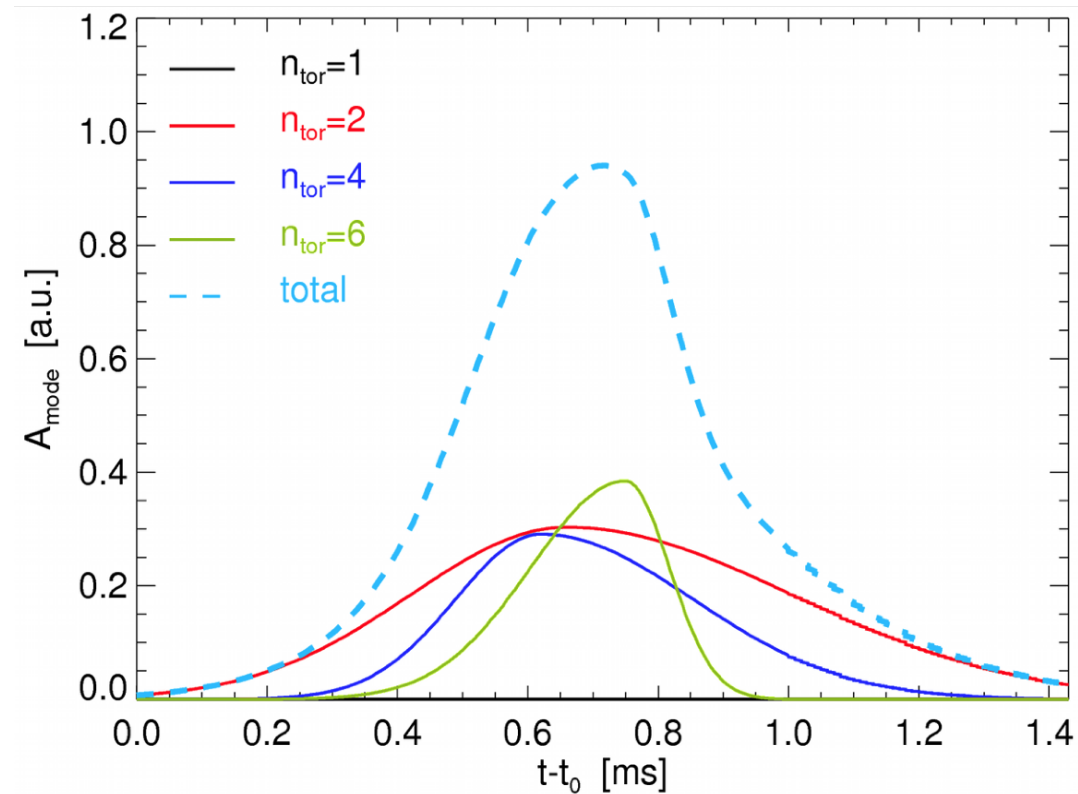
- Particle's motion is characterized by different time-scales:
 - Oscillation in wave field - *neglected*
 - 'Jumps' $\Delta E, \Delta P_\zeta$ around instantaneous energy, P_ζ
 - Slow drift from initial energy, P_ζ



Mode amplitude can evolve on time-scales shorter than typical TRANSP steps $\sim 5\text{--}10$ ms

- Need to represent F_{nb} evolution as a sequence of sub-step

- Duration δt_{step} sufficiently shorter than time-scale of mode evolution
- Examples here have $\delta t_{\text{step}} \sim 25\text{--}50 \mu\text{s}$
- *Is this compatible with NUBEAM scheme?*



Discrete bins in (P_ξ, E, μ) can contain both *resonant* and *non-resonant* particles

- ‘Resonant’ particles undergo large ΔE , ΔP_ξ variations
- ‘Non-resonant’ particles have small fluctuations around initial energy, P_ξ
- To keep track of particle’s class:
 - Sample steps σ_E , σ_{P_ξ} at first step only
 - exception: particle move to a different bin \rightarrow re-sample

$p(\Delta E, \Delta P_\xi | P_\xi, E, \mu)$ can be skewed to positive/negative $\Delta E, \Delta P_\xi$, causing overall drift of $F_{nb}(P_\xi, E, \mu)$

- Introduce ‘random sign’ for i -th step in MC procedure, $S_{r,i}$
- For each particle (e.g. pair of correlated steps σ_E, σ_{P_ξ}), calculate $S_{r,i}$ from probability of positive vs. negative steps

- From $p(\Delta E, \Delta P_\xi)$ compute

$$p_+ \doteq p(\sigma_{E,i}, \sigma_{P_\xi,i}) \quad ; \quad p_- \doteq p(-\sigma_{E,i}, -\sigma_{P_\xi,i})$$

- Then define f_{sign} :

$$f_{\text{sign}} = \frac{p_+}{p_+ + p_-}$$

- Finally, use $0 < f_{\text{sign}} < 1$ to bias random extraction of $S_{r,i} = +1, -1$

Putting all together

- At each ‘macroscopic’ TRANSP step:

I. Re-normalize bins (P_ξ, E, μ) based on q-profile, fields, ...

II. Identify ‘bin’ in (P_ξ, E, μ) for current ‘particle’ (i.e. orbit)

III. Extract steps σ_E, σ_{P_ξ} (σ_μ) from multivariate $p(\Delta E, \Delta P_\xi, \Delta \mu)$

IV. Compute sign $S_{r,i}$ from $p(\Delta E, \Delta P_\xi)$

V. Rescale steps based on $A_{\text{modes}}(t)$:

$$\overline{\Delta E}_i = S_{r,i} \times A_{\text{mode}}(t = \bar{t}) \times \sigma_{E,i}$$

VI. Advance E, P_ξ (μ):

$$E_i = E_i + \overline{\Delta E}_i$$

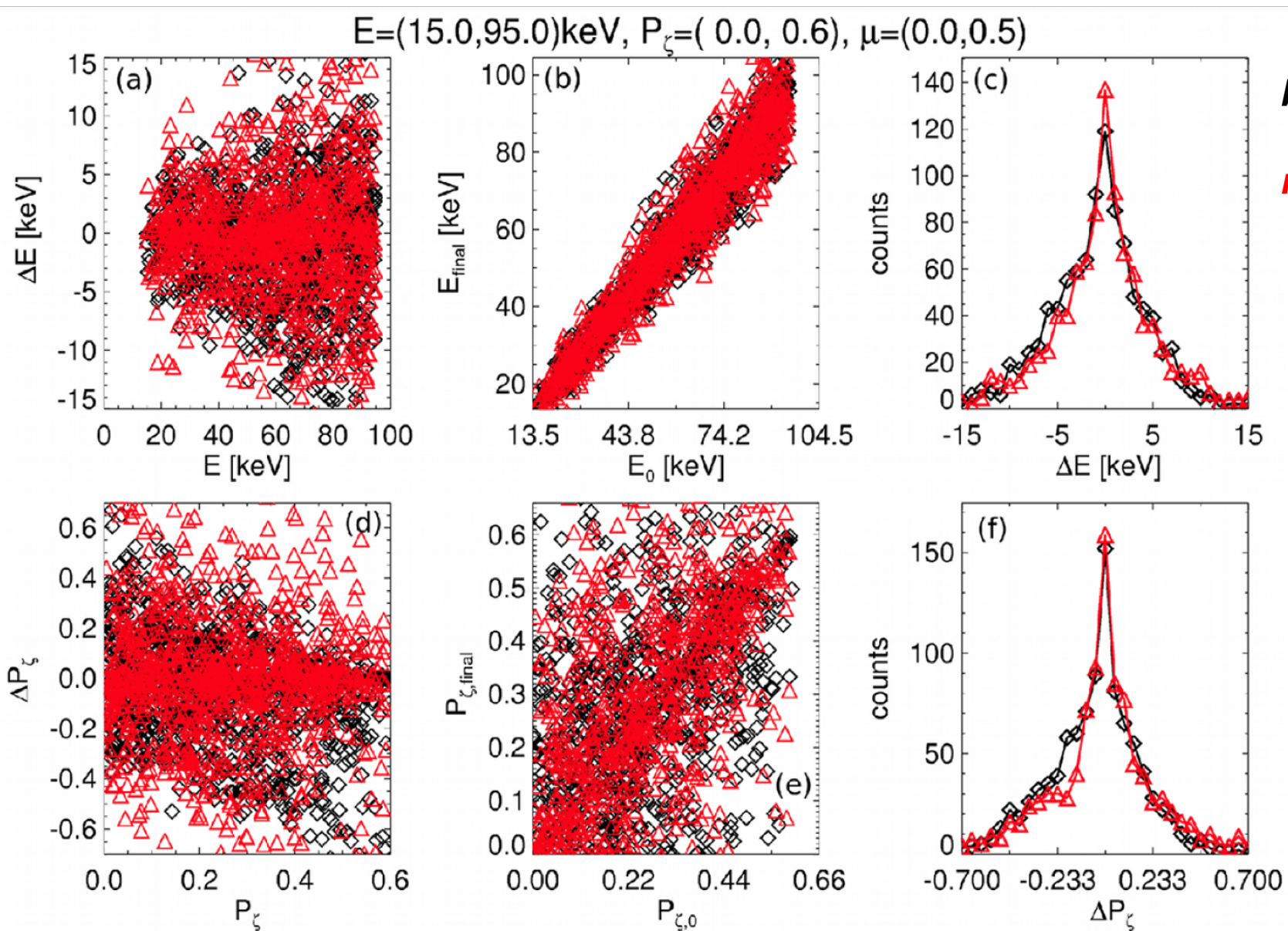
VII. Compute slowing down, scattering

VIII. Advance particle’s trajectory in phase space

Loop over particles

Steps III–VII divided in sub-steps for each particle

Example: evolving F_{nb} over 270 μs in 5 sub-steps



black: ORBIT

red: re-constructed

Reconstruction works (decently) for different classes: co-, counter-, trapped

co-passing

counter-passing

trapped

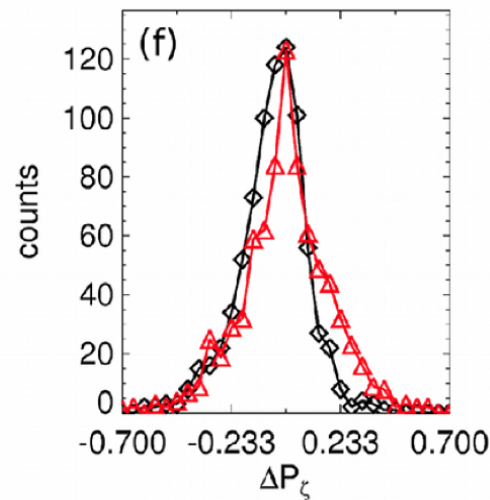
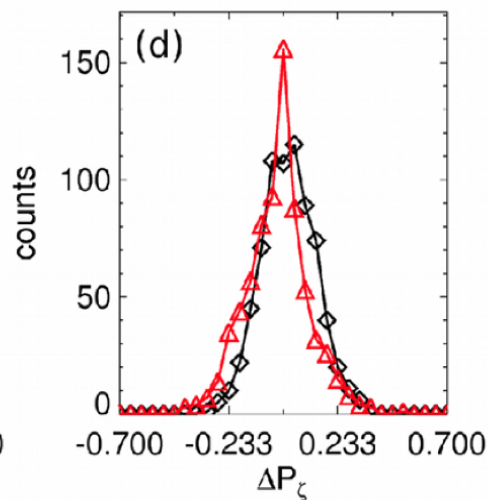
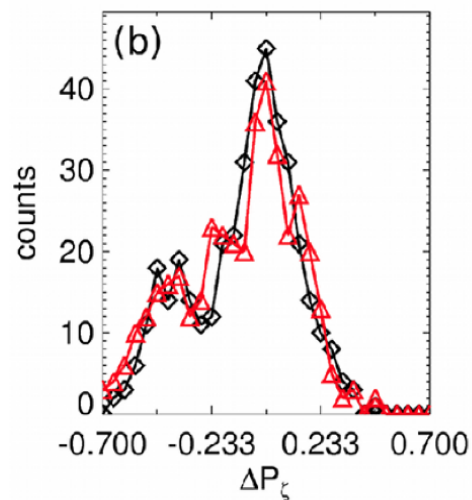
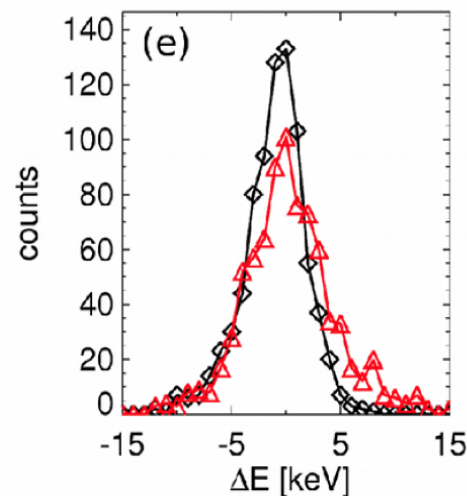
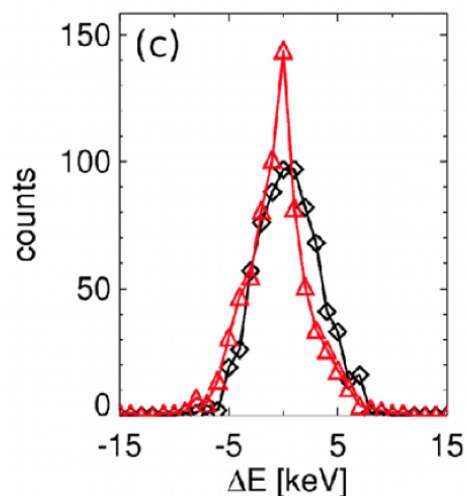
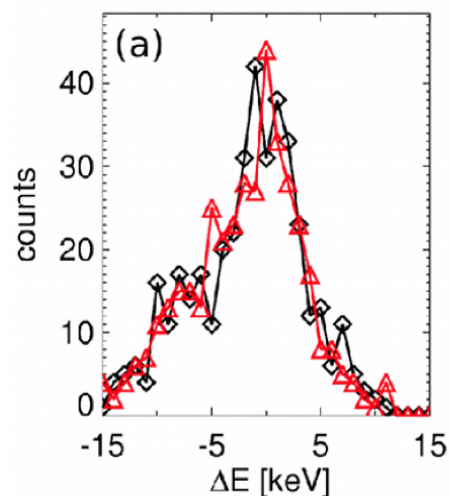
$E=(15.0, 95.0)\text{keV}$
 $P_\zeta=(0.2, 0.6)$
 $\mu=(0.0, 0.4)$

$E=(15.0, 95.0)\text{keV}$
 $P_\zeta=(-1.2, -0.7)$
 $\mu=(0.0, 0.4)$

$E=(15.0, 95.0)\text{keV}$
 $P_\zeta=(-0.6, 0.0)$
 $\mu=(1.0, 1.4)$

black: ORBIT

red: re-constructed



Reconstruction works (decently) at different sub-steps

black: ORBIT **red: re-constructed**

