



# Motivation

- Typical scattered light signals in the NSTX Thomson detectors are in the range 1,000-100,000 photoelectrons. • In many cases, systematic calibration uncertainties predominate over photon statistics in determining the computed errors in the measured signals, which are used as weights for the least squares fits for  $T_e$  and  $n_e$ .
- The most significant contribution from the calibration is the quoted uncertainty in the the traceable calibration of the reference detector used to characterize the spectral response of the filter polychromators.
- A statistical study of the residuals between the measured and fitted signals for a large number of laser shots can illuminate systematic offsets, some of which are consistent with errors in the reference detector responsivity.
- This system from is currently being upgraded from 20 to 30 spatial channels, and a recalibration of the entire system is planned in the near future. This study provides insights into potential improvements in this process.
- This type of analysis has been applied previously to LIDAR Thomson data on JET. "A new method for estimating and identifying systematic error in Thomson scattering diagnostics", H. Fajemirokun, C. Gowers, P. Nielsen, H. Saltzmann, and K. Hirsch. RSI, 61, (1990)
- The above paper describes how this technique was able to discern a change in the window transmission, and there is potential that we could do the same on NSTX to monitor window conditions.

#### Strategy

A) Compare measured  $\chi^2$  distribution with theoretical distribution for 3 degrees of freedom:  $P_k(x) = \frac{(x)^{3/2-1}e^{-x/2}}{\Gamma(3/2)2^{3/2}}$  (1) for  $x_j = \chi_j^2 = \sum_i \frac{(y_{ij} - f_{ij})^2}{\sigma_{ij}^2}$  (2) which peaks at a value of  $\chi^2 = 1$  for our case, assuming we do not use the Rayleigh channel in the fitting. Here y<sub>ii</sub> is the measured number of photoelectrons, f<sub>ii</sub> is fitted value, and  $\sigma_{ii}$  is the modeled uncertainty in spectral channel i of spatial channel j.

If, for a large number of samples, the measured distribution is different from (1) it could mean that one or more the following is untrue:

- a. The assumed fitting function describes the physics b. Calibrations are correct
- c. Estimated errors have been properly assigned
- d. Residual distributions are Gaussian

**B**) In order to obtain more information, it is useful to look at the **normalized residuals**:

For a large sample size (k=1, N) of laser shots,  $r_{ii}^k$  should be distributed with a Gaussian distribution:  $P(z_{ij}^k) = -$ 

If a) through d) above are correct, then the distribution should be centered about  $\mu_{ii} = 0$ , with a width  $s_{ii} = 1$  or equivalently  $FWHM = 2\sqrt{2\ln 2} = 2.35$ . Nonzero values of  $\mu_{ii}$  indicate a failure of one or more of a) through d), the most likely candidate being a systematic calibration error.  $s_{ii} \neq 1$  indicates that errors have not been properly

C) If there are systematic offsets in the normalized residual distributions, and if one assumes these are due to systematic calibration errors of individual detector sensitivity, one can calculate correction factors c

or  $c_{ij} = \sum \frac{f_{ij}^{\kappa}}{\sigma_{ij}^{k}} / \sum \frac{y_{ij}^{\kappa}}{\sigma_{ij}^{k}}$ 

where the summation over k is for thousands of laser pulses.

• Thus knowing the observed signals, the fitted values and the estimated errors permits calculation of correction factors for individual detectors. • Multiplying the observed signal in each detector by the appropriate correction factor and refitting the data allows the process to be iterated to convergence.

## **NSTX Thomson Scattering System** • Operational for several years, the NSTX Thomson system currently has 20 spatial channels, each with a 6 wavelength filter polychromator designed by General Atomics. • The filter passbands are shown at right, convolved with the quantum efficiency of the APD

- Excluding the channel at the laser wavelength used for Rayleigh scattering, this statistical study made use of signals from the 5 other spectral channels, used to fit the scattered spectrum.
- We have used a set of ~ 1000 recent NSTX discharges each with ~ 40 laser shots. We have constrained the data used in this study such that there are at least 2000 photoelectrons in the two short wavelength polychromator channels.



- Because we were interested in discerning systematic effects over a broad see a wide range in temperatures.
- The position range is shown at left, which also shows the change in the the final case 3 analysis.
- Sources of uncertainty in the Thomson signals are shown at the right.

detectors.

- The first row shows the "statistical" errors due to photoelectron statistics in the scattered and plasma light.
- The other rows list the uncertainties resulting from various calibrations.
- The terms in red always dominate the modeled uncertainty

fractional incertainty	description	range of values	comments
$rac{\sigma^2_{N^{ij}_{TS}}}{(N^{ij}_{TS})^2}$	estimate of the photoelectron noise normally dominated by the scattered photoelectrons $N_{TS}$ and/or the plasma signal $N_{pl}$	.005 – 0.10	$\sigma_{N_{TS}^{ij}}^{2} = (N_{o}^{ij})^{2} + [F(M)(N_{TS} + N_{Pl})]^{ij}$ readout excess scat plasma noise noise pe pe
$rac{\sigma_{_{m{v}_{fs}}}}{m{v}_{_{fs}}}, rac{\sigma_{_{V_{fs}}}}{V_{_{fs}}}$	errors in the fast and slow outputs in the calibration that relates fast and slow sensitivities	.001003	statistical measurement error, can reduce with further averaging
$rac{\sigma_{\left(rac{M_x}{M_{sp}} ight)}}{\left(rac{M_x}{M_{sp}} ight)}$	error in the ratio of APD gains used in the TS measurement and the spectral calibration	not yet measured	not currently measured
$rac{\sigma_{_{V_{sp}}}}{V_{_{sp}}(\lambda)}$	errors in the slow outputs during the spectral calibration	.002 - ,004	statistical measurement error, can reduce with further averaging, finer wavelength intervals
$rac{\sigma_{_{V_{ref}}}}{V_{_{ref}}(\lambda)}$	errors in the reference detector signals during the spectral calibration	.002005	statistical measurement error, can reduce with further averaging, finer wavelength intervals
$rac{\sigma_{S_{ref}}}{S_{ref}(\lambda)}$	quoted uncertainty in the responsivity of the reference detector	.01503 (.0306)	systematic error, cannot be reduced by more averaging or finer wavelength intervals (range in parentheses indicates discrepancy between two calibrated reference detectors)
$rac{\sigma_{T_c^j}}{T_c^j(\lambda)}$	errors in the determination of the spectral transmission of the collection optics and fiber bundles	.00050015	statistical measurement error, can reduce with further averaging, finer wavelength intervals

# **Study of Uncertainties in NSTX Thomson Scattering Data**

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spectral range, we confined the study to the central set of positions, which

temperature profile which occurred between the baseline case 1 analysis and

## **Uncertainty Estimates**

# **Baseline Analysis - Case 1**

- $\chi^2$  Distribution: The distribution of  $\chi^2$  for one of the positions is shown in the upper left panel at the left. The experimental distribution in black has a very different shape from the theoretical curve in red, with a higher peak at lower values. This can be understood by looking at the normalized residual distributions.
- **Normalized Residual Distributions**: Distributions of the normalized residuals *z* are shown in the other panels at the right, one for each of the five wavelength channels for this position. The red curves are what one would expect if there were no systematic calibration errors, and if the  $\sigma^k$  are being modeled correctly. Note that z=1 corresponds to a residual equal to one standard deviation. Both experimental and theoretical curves have been normalized to have an integral equal to one. The green curves are Gaussian fits to the observed distributions. Above each P(z) curve are three numbers. The first is the **fitted offset**, the second is the **fitted width**, and the third is the **mean number of photoelectrons** for the data in the distribution. Note that the experimental distributions in black are much **narrower than expected**, indicating that the **modeled**  $\sigma^k$  are too large, consistent with the experimental  $\chi^2$  peak being too low.
- **Correcting the uncertainty model:** Motivated by the above discrepancies, we reexamined our model for the uncertainties  $\sigma$ . We found two partially offsetting problems. First, there was an error in coding the uncertainty due to the various calibration contributions. Second, the error contribution from the reference detector responsivity was incorrectly treated as a statistical measurement error rather than as a systematic error.
- Impact of the improved analysis: The panel at left below shows the fractional error due to scattered photoelectrons,  $\sigma_{N_{TS}}/N_{TS}$ , as a function of time for the 5 spectral channels at a central position. The central panel shows the fractional error from the old calculation of the combined calibration components  $\sigma_{cal}/N_{cal}$  from all but the first row of the table at the lower left. The right panel shows the corrected analysis for  $\sigma_{cal}/N_{cal}$ , showing a significant decrease in this contribution. In the corrected analysis, the error due to photon statistics dominates the total  $\sigma$ .



- analyzed using the improved uncertainty model.
- normalized residual distributions P(z).
- For most channels, the width of P(z) is much closer to the centers of these distributions from z=0.
- theoretical  $\chi^2$  distributions.







# **Analysis Applying Statistical Correction Factors - Case 3**

- wavelength channels.





- P(z) distributions result. This needs to be understood.
- of the plasma light component of the uncertainty.





## Conclusions

• We have identified errors in our previous model for uncertainties used as weights in the least square fits for  $T_e$  and  $n_e$ .

• The combined results of the improved uncertainty model and application of the correction factors is shown in the left two panels below. The reduction in the fractional error  $\delta T_e / T_e$  is shown for a central position in the plot at right below. The correction factors have a common spectral dependence across many sample positions in the plasma, consistent with



• It should be possible to test the sensitivity of this technique to various spectral 'shapes' of systematic error.

• A database of 100-200 discharges are needed to discern correction factors of  $\sim 1\%$ . It would be interesting to analyse data from periods where window coatings were known to cause a relative change in the window transmission by > 5% across the spectrum.

• If the data is filtered to permit laser shots resulting in signal levels lower than 1000 pe in the widest channels, distinctly non-Gaussian

• The trend toward narrower P(z) distributions at lower wavelength needs to be investigated. Perhaps in indicates an error in the treatment

• We have obtained a new calibrated reference detector, and it should be used, along with the statistical tools developed in this study, to investigate whether we can better understand the true source of the systematic errors.