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#### **On the Analysis of HHFW Heated Plasmas in NSTX**

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#### Newer TORIC in TRANSP Provides Improved Analysis Tool Supplement analysis with CQL3D

- TRANSP makes use of recent version of TORIC, which can compute HHFW propagation and absorption in NSTX
  - M. Brambilla, Plasma Phys. Control. Fusion 44 (2002) 2423-2443
- TORIC calculates power deposition into all species including fast ions
  - But TRANSP RF Monte Carlo Fokker-Planck operator is not ready
  - Self-consistent calculation of fast ions not available for NBI + HHFW plasmas
- Use CQL3D to estimate neutron rate generated by fast ions
- Analyze two cases
  - HHFW generated high-T<sub>e</sub> plasmas
  - HHFW heating of NBI-induced H-mode plasmas



## TORIC/TRANSP Calculation Assumptions *Power absorption hypothesis*

- TRANSP/TORIC assumes that only the fast wave is propagating
  - AORSA calculations show mode-conversion effects small
- Assume that all the antenna power is absorbed
  - This assumption is not met experimentally, but provides a uniform reference
    - Edge/coupling physics effects have been identified: excitation of surface wave and PDI ion heating, which can absorb up to 30% of the power in the plasma periphery
    - J.C Hosea, et al., Physics Plasmas 15 (2008) 056104
    - T. Biewer et al, Physics of Plasmas **12** (2005) 056108
  - Efficiency will be addressed later by comparing with experimental neutron production rate
- Notation
  - Qa = Volint(qa) where qa is the total power density coupled by the antenna
  - Qe = Volint(qe) where qe is the power density coupled to electrons
  - Qf = Volint(qf) where qf is the power density coupled to the fast ions

#### High T<sub>e</sub> Achieved during HHFW Heating in He and D<sub>2</sub> Plasmas



**NSTX** 

51st Meeting APS-DPP, HHFW Analysis, B.P. LeBlanc, PPPL

### Competition between Electron and Fast-ion Damping *Qe reduced when fast ions are present*

- All HHFW power absorbed by electrons prior to NBI pulse starting at 0.2s.
- After NBI onset, the fast-ion population absorbs HHFW power at the expense of the electrons
  - Long lasting effect since the fastion density n<sub>f</sub> persists beyond 0.4s.
  - TORIC/TRANSP consistent with single time point calculations done with AORSA, GENRAY and TORIC



 $Q_a$ : power delivered by antenna  $Q_e$ : power absorbed by electrons  $Q_f$ : power absorbed by fast ions  $n_f$ : on axis fast-ion density NBI: NBI power



## HHFW Heating of NBI-Induced H-mode Plasma Achieved using k<sub>b</sub>=-13m<sup>-1</sup>

- Previous attempts at HHFW heating of NBI-induced H-mode plasmas were unsuccessful[i], but recent application of k<sub>\u03c0</sub>=13m<sup>-1</sup> HHFW power resulted in measurable change in the stored energy and kinetic measurements.
- Better understanding of edge effects and attention to the edge density were conducive to this power coupling improvement[ii].

 B.P. LeBlanc, 16th RF Conference, AIP Conference Proceedings 787, p.86

• [ii] J.C. Hosea, et al., RF Conference, Ghent, Belgium, 2009



#### Measured W<sub>tot</sub> and T<sub>e</sub> Increase during HHFW Heating of NBI-induced H-mode Plasma





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#### Measured Stored Energy and Neutron Rate Exceed TRANSP Calculations during HHFW Pulses



#### Will look at data in bottom panel



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#### CQL3D Suggests ≈40% of HHFW Power Ultimately Coupled to Plasma Core





## TORIC/TRANSP HHFW Power Deposition Results for Heating of NBI-Induced H-mode Plasma, $k_{\phi}$ =-13m<sup>-1</sup>

- HHFW power, Qa, is divided between the electrons, Qe, and the fast ions, Qf
- Qf decreases with time as the fast ions thermalize
  - (e.g. as fast-ion density decreases)
- Power deposition profiles (qa, qe,qf) become peaked at the end of HHFW phase\_\_\_\_\_\_





## TRANSP with TORIC has been used to Estimate HHFW Heating

- The current implementation of TORIC into TRANSP, although not fully consistent for fast ions, provides revealing information about the time and profile evolution of the HHFW power absorbed by plasma species
- TRANSP analysis has been complemented with single time point CQL3D calculation to determine the ultimate HHFW power coupling based on neutron production
- More work needed to continue validating modeling for HHFW heating in NSTX plasmas

TI3.00002: G. Taylor, Advances in High-Harmonics Physics in NSTX

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# **EXTRA SLIDES**



#### **HHFW Progress during 2009 Campaign**

TI3.00002: G. Taylor, Advances in High-Harmonics Physics in NSTX





#### **HHFW Power Shifted from Electrons to Fast Ions during NBI**



# Surface FW Propagation Supports Surface Loss at Lower k<sub>II</sub>



- Propagation is very close to wall at  $k_{\parallel} = 8 \text{ m}^{-1}$ , on wall at  $k_{\parallel} = 3 \text{ m}^{-1}$  J.C.Hosea
- Losses in surface should be higher for lower k<sub>ii</sub>
- Propagation angle relative to B much less than for lower harmonic case
- Increasing B should move onset farther from antenna, increasing heating

**WNSTX** 

# TORIC VERIFICATION Compare to codes AORSA and GENRAY



 $N\varphi = 12.$  •C.K. Phillips, Nucl. Fusion 49 (2009) 075015

#### NSTX HHFW Heating System Fixed frequency 30 MHz, variable $|k_{\phi}| = 3, 8, 13, 14 \text{ m}^{-1}$

- System parameters:
  - 12 antenna straps, six sources, decoupling loops, phase shifters, and stubs
  - Digital control of power and phase
  - Automated matching calculation



The double-peak  $k_{//}=(14,18)m^{-1}$  spectrum will be referred to as simply  $k_{//}=14m^{-1}$ .



 $W_{mhd}$  and neutron ( $S_n$ ) changes are small and reproducible, but appear related to edge effects



#### NBI into HHFW Pre-heated Plasma Three NBI Pulses



#### HHFW starts before H-mode onset: three NBI pulses









