Threshold Physics Makes an NTM Linearly Stable and Non-Linearly Unstable

$$\frac{\tau_{R}}{r^{2}} \frac{dw}{dt} = \Delta' + \epsilon^{1/2} \frac{L_{q}}{L_{p}} \beta_{p} \begin{bmatrix} \frac{w}{w^{2} + w_{d}^{2}} & -\frac{w_{pol}^{2}}{w^{3}} \end{bmatrix}$$

$$\frac{\sqrt{1}}{ransport} \qquad polarizat$$

tion threshold threshold

- **Transport threshold** (R. Fitzpatrick 1995)
 - \star transport along \vec{B} in island is fast compared to perpendicular
 - helical pressure perturbation washed out if perpendicular transport dominates

$$w_d \approx \left(\frac{L_s^2}{k_{\theta}^2} \frac{\chi_L}{\chi_{II}}\right)^{1/4} \sim 1 \text{ cm}$$

- Polarization threshold (H.R. Wilson et al., 1996)
 - ★ inertial effects are important in frame of E×B equilibrium flow
 - polarization currents induced by island propagation are stabilizing for $\omega(\omega_{*i} - \omega) > 0$

$$w_{pol} \approx (L_q/L_p)^{1/2} \epsilon^{1/2} \rho_{\theta i} \sim 1 \text{ cm}$$

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